

Heavy Photons, Axions, and WISPs (+ sub-GeV dark matter)

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8th Patras Workshop on Axions, WIMPs, and WISPs

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What motivates a new particle/interaction?

- Deep (real) puzzles

Examples:

- mechanism of EWSB?

major progress with LHC discovery of
a new boson at ~ 125 GeV!

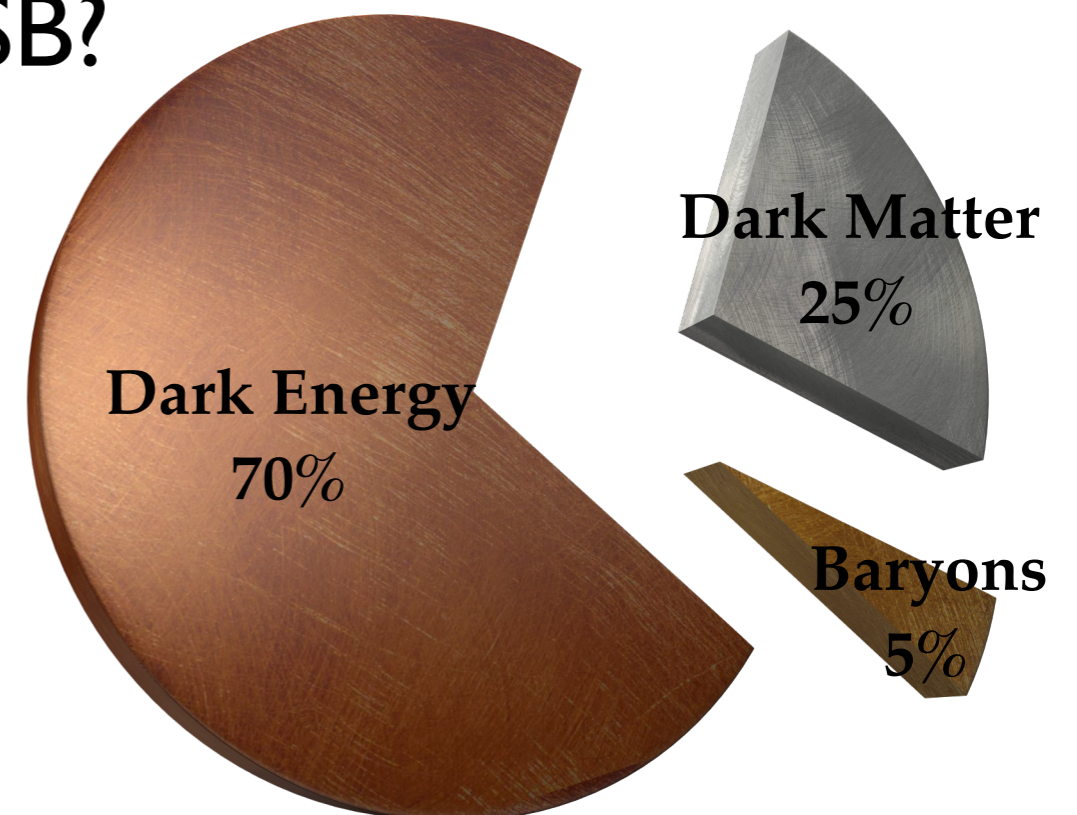
SM Higgs!?

What motivates a new particle/interaction?

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Examples:

- mechanism of EWWSB?
- dark matter?
- dark energy?



What motivates a new particle/interaction?

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Examples:

- mechanism of EWWSB?
- dark matter?
- dark energy?
- hierarchy problem/naturalness

e.g. Supersymmetry

No evidence currently that Weak scale is “natural”

What motivates a new particle/interaction?

- Deep (real) puzzles

Examples:

- mechanism of EWSB?
- dark matter?
- dark energy?
- hierarchy problem/naturalness
- strong CP problem

(more later)

What motivates a new particle/interaction?

- Deep (real) puzzles
- Intriguing observations

Example:

- WIMP miracle + hierarchy problem

main motivation for direct, indirect,
and collider searches for DM

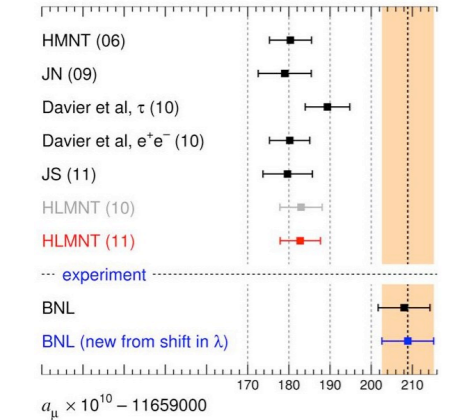
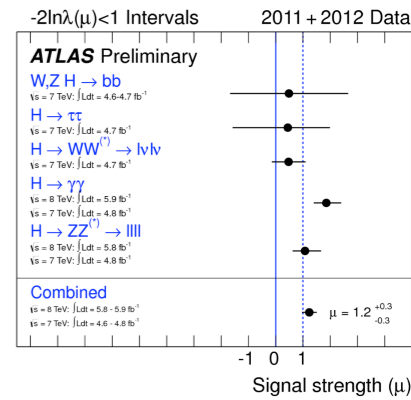
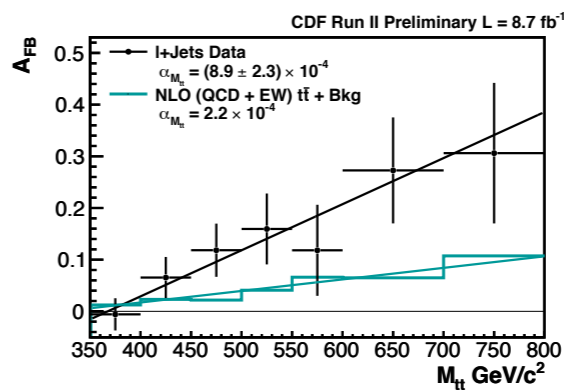
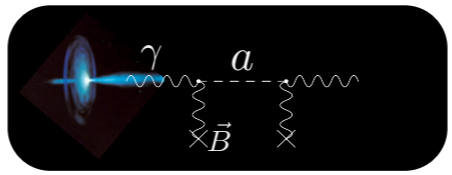
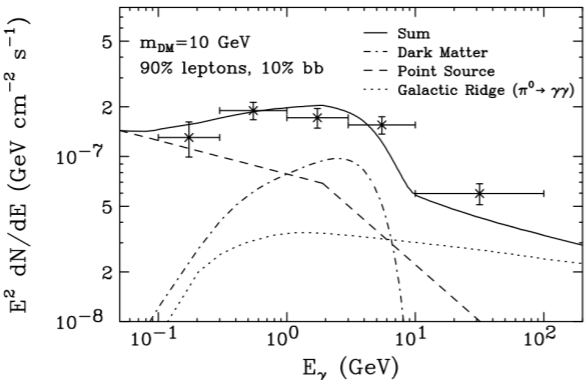
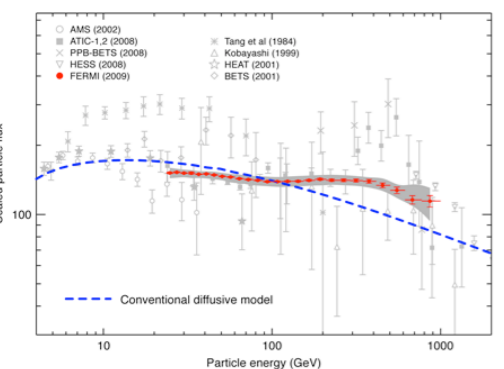
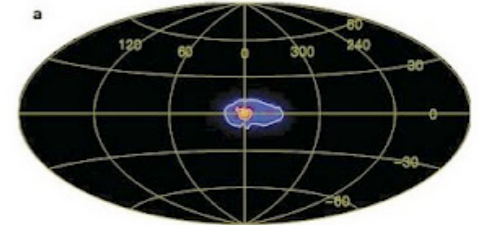
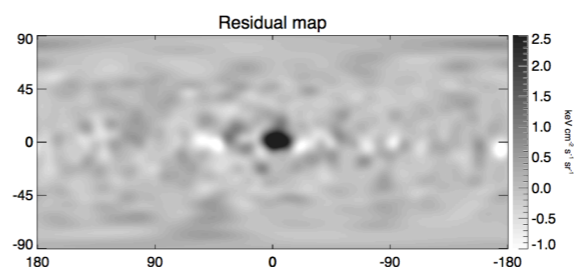
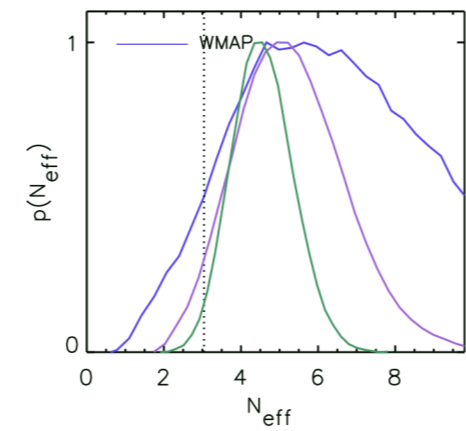
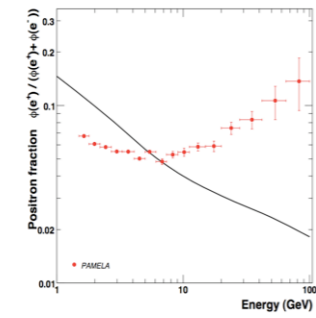
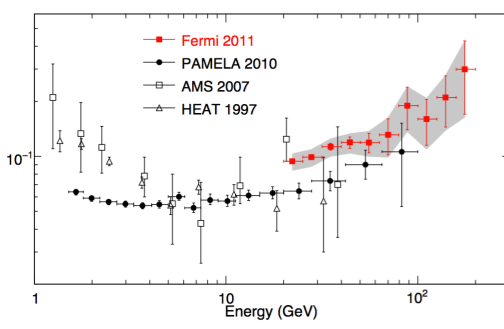
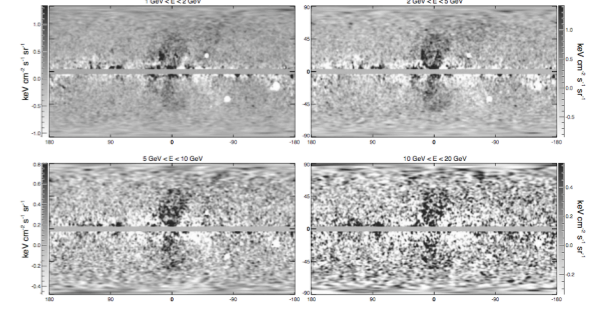
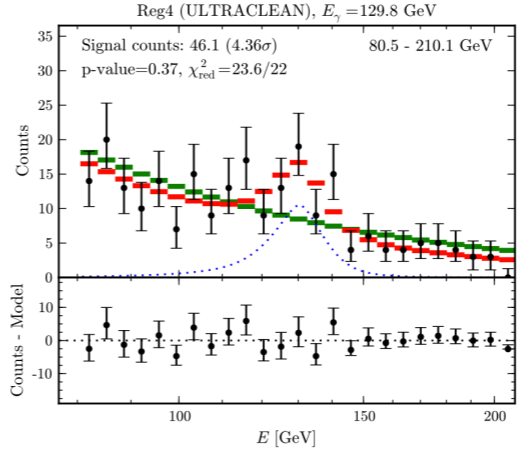
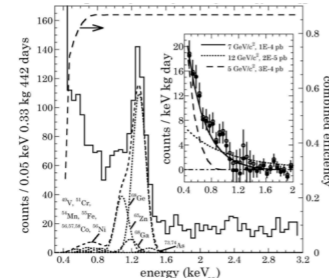
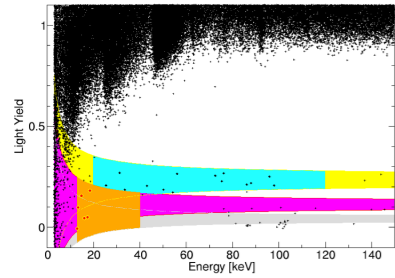
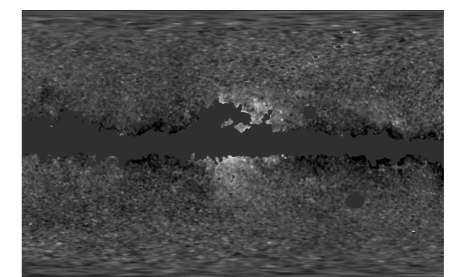
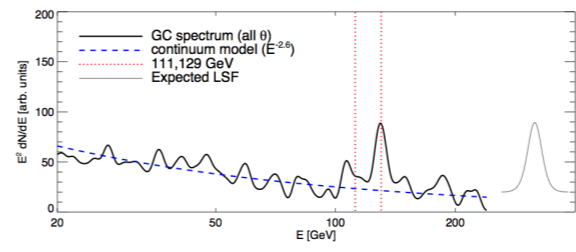
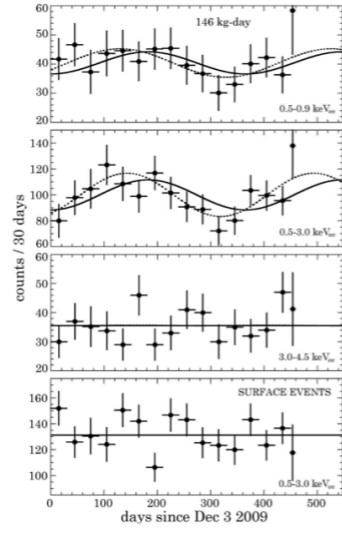
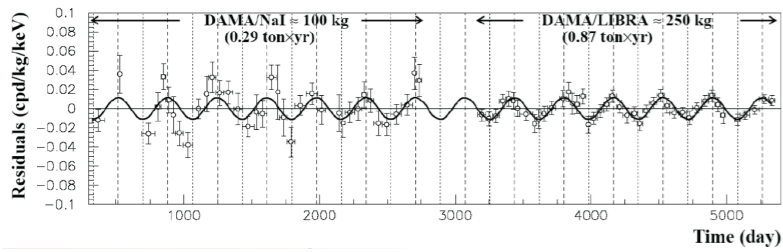
(but may be totally misleading us!)

What motivates a new particle/interaction?

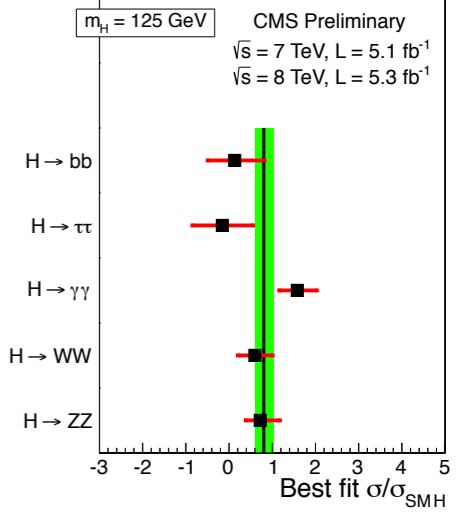
- Deep (real) puzzles
- Intriguing observations
- data anomalies

Examples:

- many!



I'm $>5\sigma$ confident that not all these are real...



What motivates a new particle/interaction?

- Deep (real) puzzles
- Intriguing observations
- data anomalies
- theoretically viable

theorists have a lot of imagination...

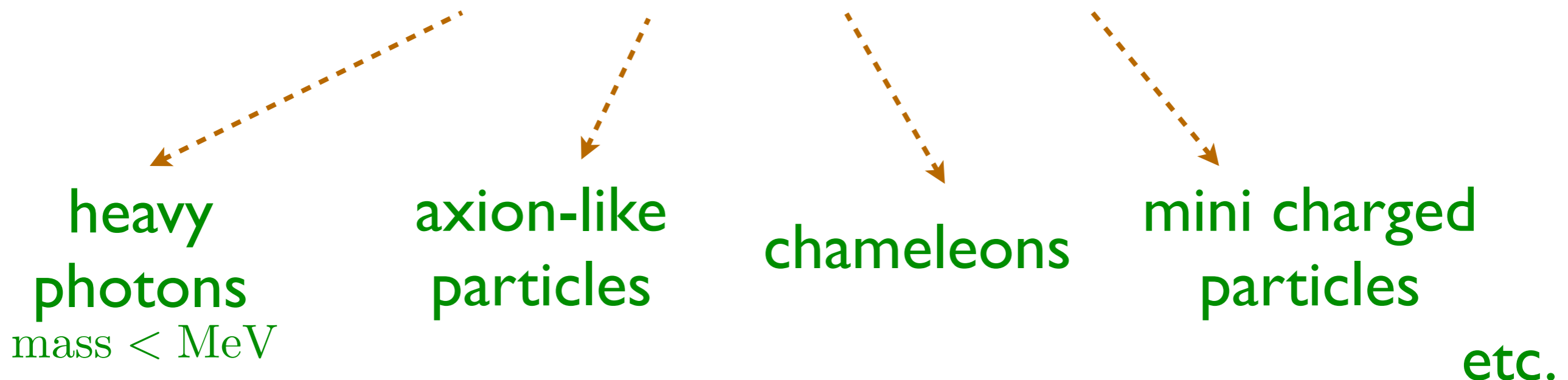
sometimes, exploring a new experimental frontier can lead to amazing discoveries

Heavy Photons, Axions, and WISPs

strongly motivated from theory, strong CP problem,
dark matter, dark energy, muon $g-2$,
and/or various astrophysics anomalies

“WISPs”

Weakly **I**nteracting “**S**ub-eV” (or “**S**lim”) **P**articles



Heavy Photons, Axions, and WISPs

strongly motivated from theory, strong CP problem,
dark matter, dark energy, muon $g-2$,
and/or various astrophysics anomalies

light particles, with weak coupling to ordinary matter

probe with high intensity beams of electrons,
protons, or photons, and sensitive detectors

experiments are *much* smaller scale than LHC

offer incredible pay-off if they find something,
for relatively small investment

Outline



- axions & axion-like particles
- heavy photons
- direct detection of sub-GeV DM

What are they?
Motivation?
Opportunities?

no time to talk about chameleons,
but see talks by A. Upadhye and K. Zioutas

Axions: solve strong CP problem

theoretically, QCD should violate CP: $\mathcal{L} \supset \frac{\alpha_s}{4\pi} \theta \text{tr } G\tilde{G}$

experimentally, no CP violation: $\theta \lesssim 6 \times 10^{-10}$

\implies strong CP problem

Solution: set $\theta = \frac{a}{f_a}$ with new symmetry, $U(1)_{\text{PQ}}$,

spontaneously broken at high scale f_a

\implies new particle (axion) and $\theta \rightarrow 0$

$$m_a \sim \frac{\Lambda_{\text{QCD}}^2}{f_a} \simeq 0.6 \text{ meV} \frac{10^{10} \text{ GeV}}{f_a} \quad \text{naturally light}$$

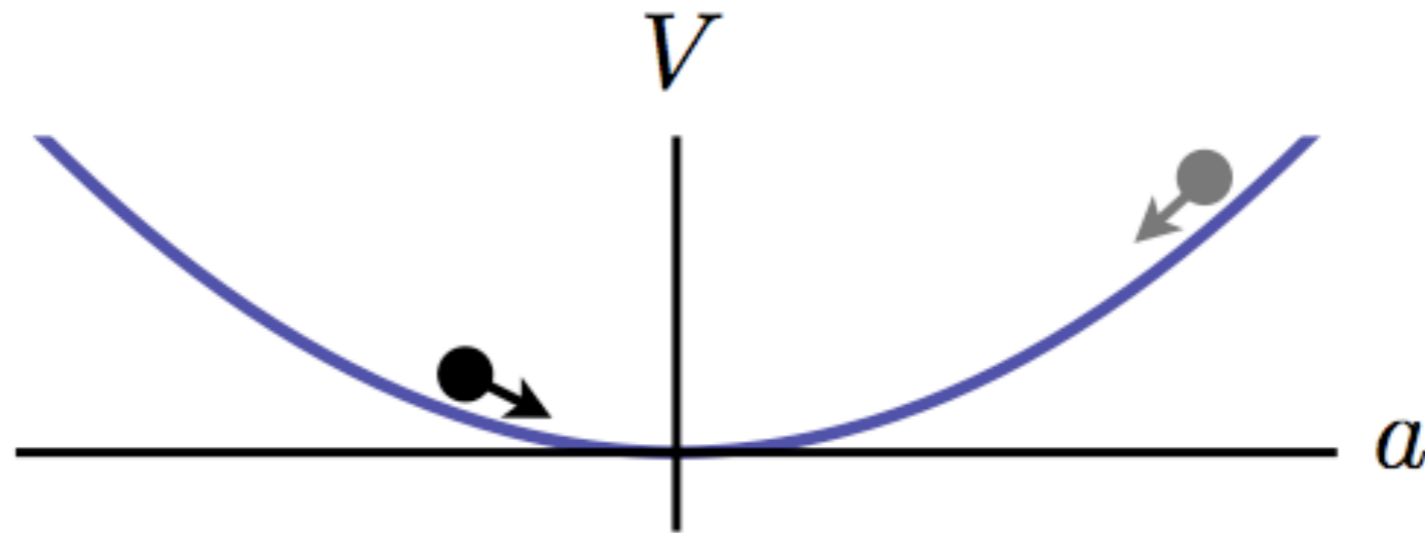
Axion-Like Particles (ALPs)

breaking of PQ symmetry gives axion...

general: breaking of symmetry at high scale can give
Axion-Like Particles with small masses

generic in many well-motivated extensions of Standard
Model, and top-down approaches (string theory)

Axions & ALPs can be dark matter



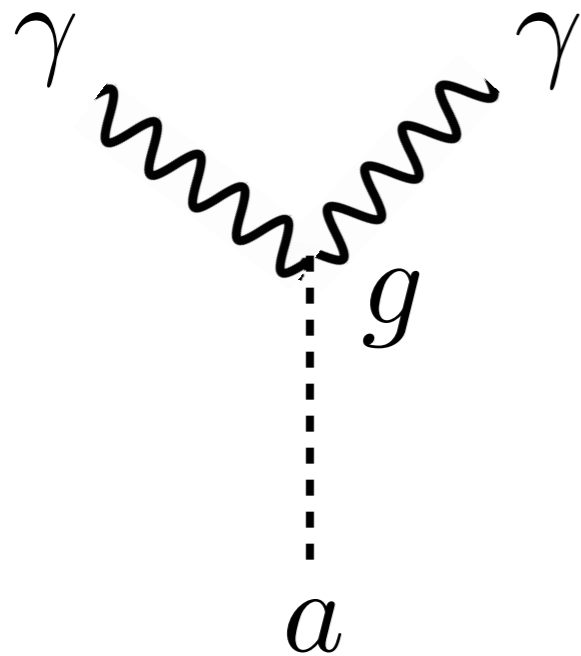
QCD effects provide a potential to axion field,
driving it to CP conserving minimum

oscillations of axion field around
minimum act as cold DM

Couplings to ordinary matter

axions couple to fermions & photons

e.g. coupling to photons:

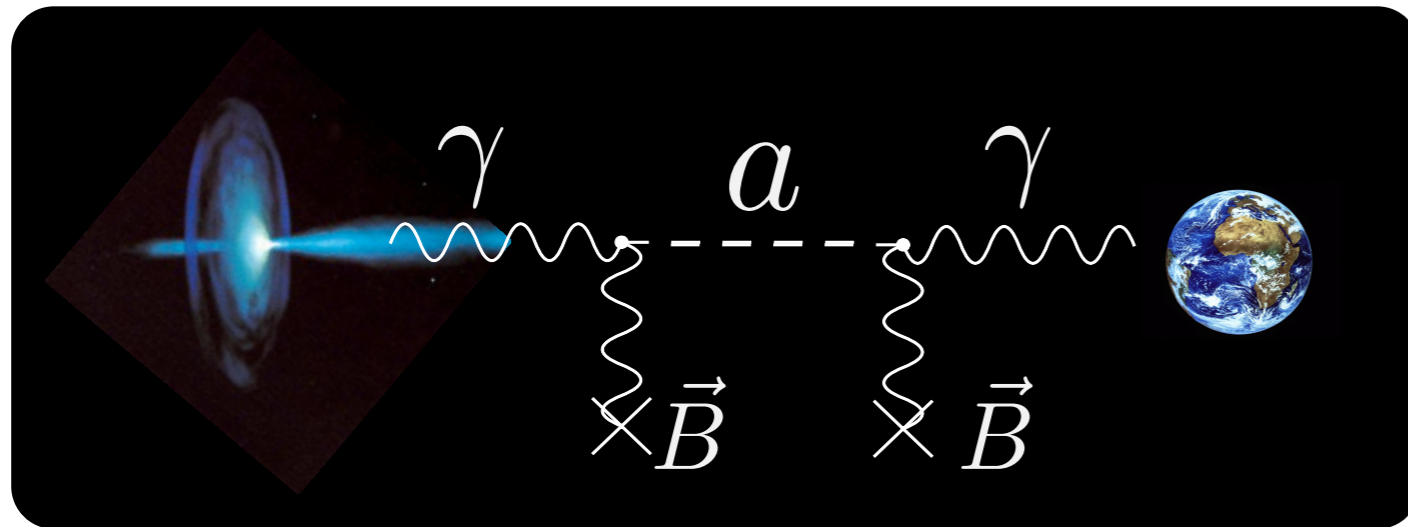


$$g \sim 10^{-13} \text{ GeV}^{-1} \left(\frac{10^{10} \text{ GeV}}{f_a} \right)$$

coupling suppressed by f_a

for ALPs, coupling to photons
can be different (even zero)

Axion/ALPs can solve astro puzzles

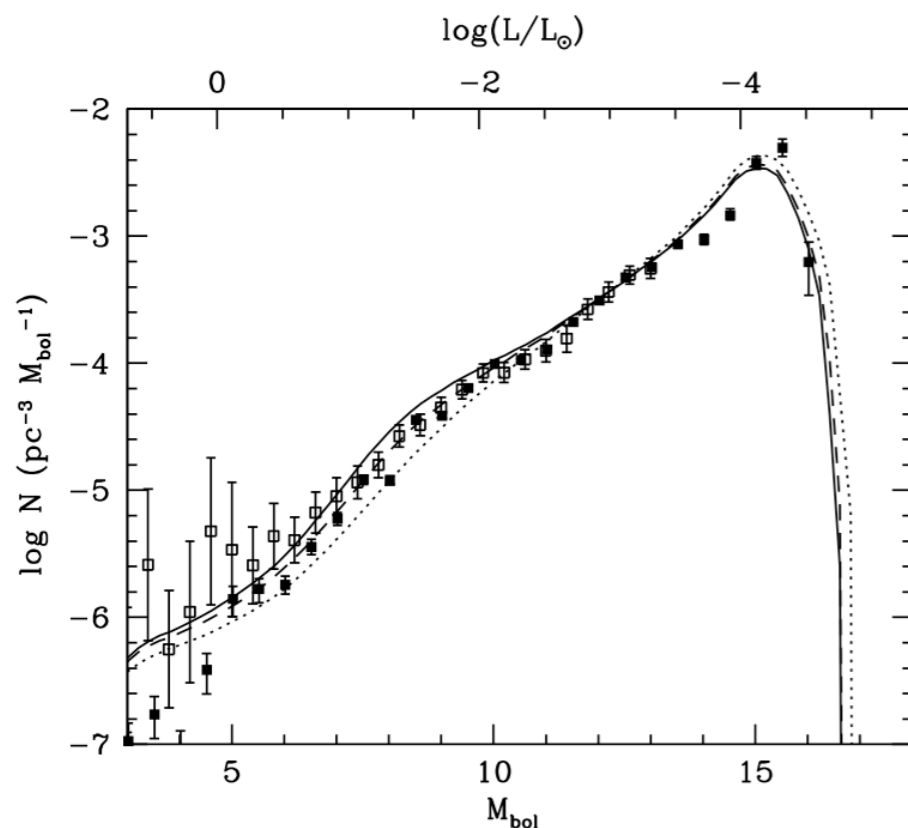


Is universe more transparent than expected to high energy γ -rays?

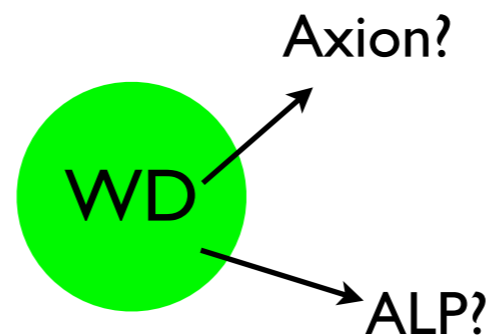
γ -ALP conversion?

Roncadelli, de Angelis, ...

see talks by M. Meyer



Do white dwarf stars cool faster than expected?



cooling enhanced by axion/ALP radiation?

Isern, Garcia-Berro, Torres, Catalan

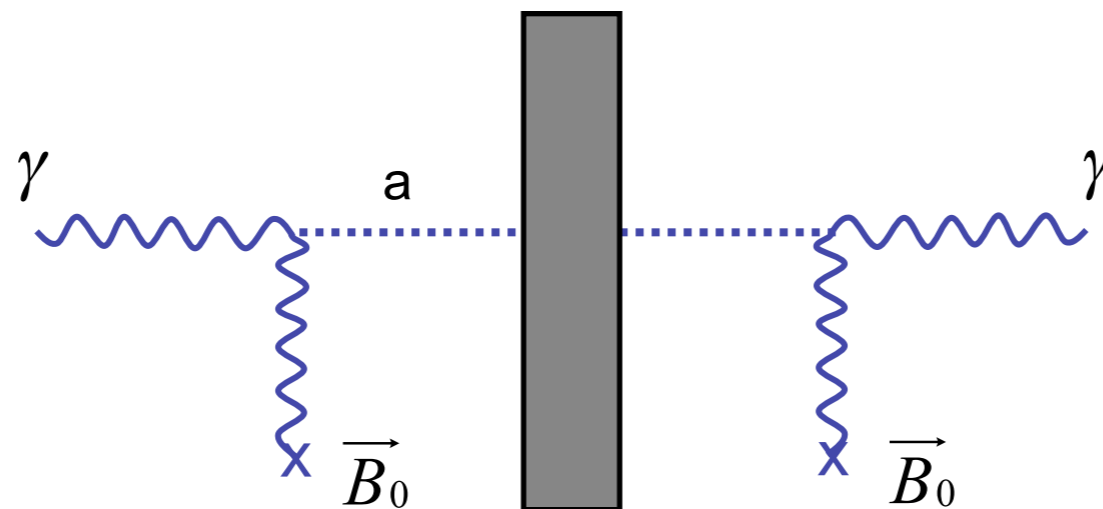
see talks by J. Isern & R. Gill

How to look for Axion and ALPs?

Best probes from γ -axion/ALP conversion

“Light-shining-through-walls”

Okun; Sikivie; Anselm; van Bibber;



LIPSS (Jlab) , BFRT (BNL), BMV (LULI), GammeV (Fermilab),
ALPS (DESY), OSQAR (CERN), PVLAS (INFN), ...

Need large magnets, powerful lasers, optical cavities

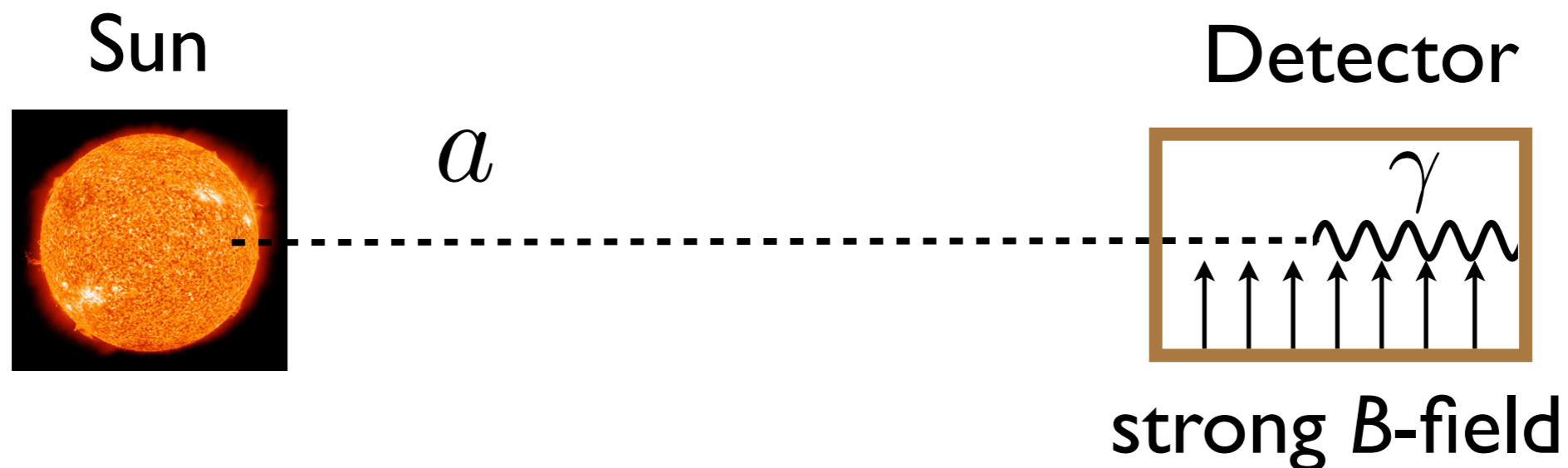
see talks by R. Baehre, M. Betz, M. Sulc

How to look for Axion and ALPs?

Best probes from γ -axion/ALP conversion

Helioscopes: stare at the sun

Sikivie; ...



SHIPS, CAST, SUMICO, IAXO, ...

Need large magnets, sensitive detectors

see talks by J.Vogel (IAXO) & J. Ruz Armendariz (CAST)

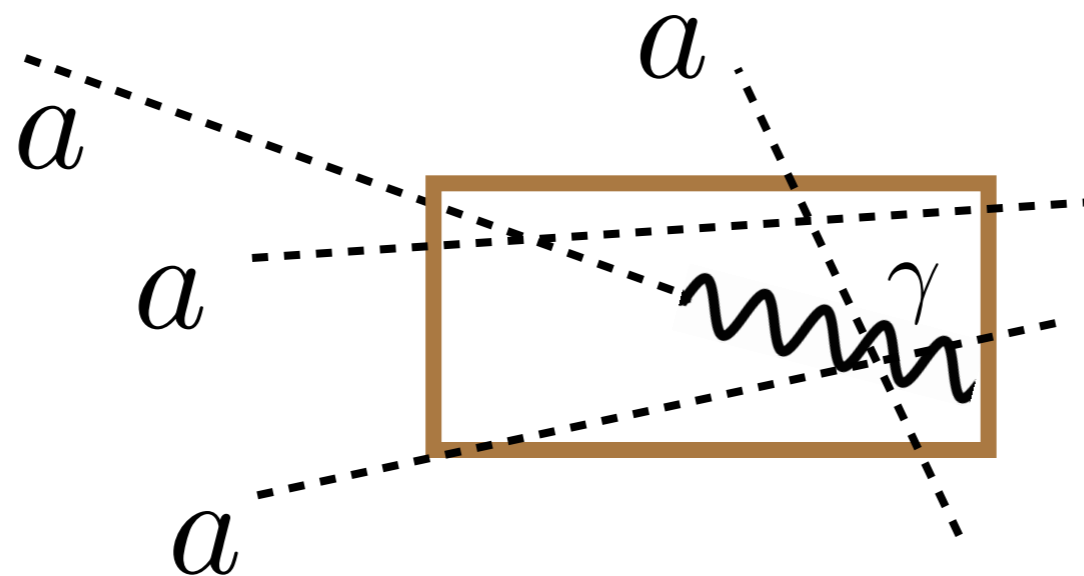
How to look for Axion and ALPs?

Best probes from γ -axion/ALP conversion

Resonant Cavities with Large Magnetic Field

Sikivie;

assume
axions are
dark matter



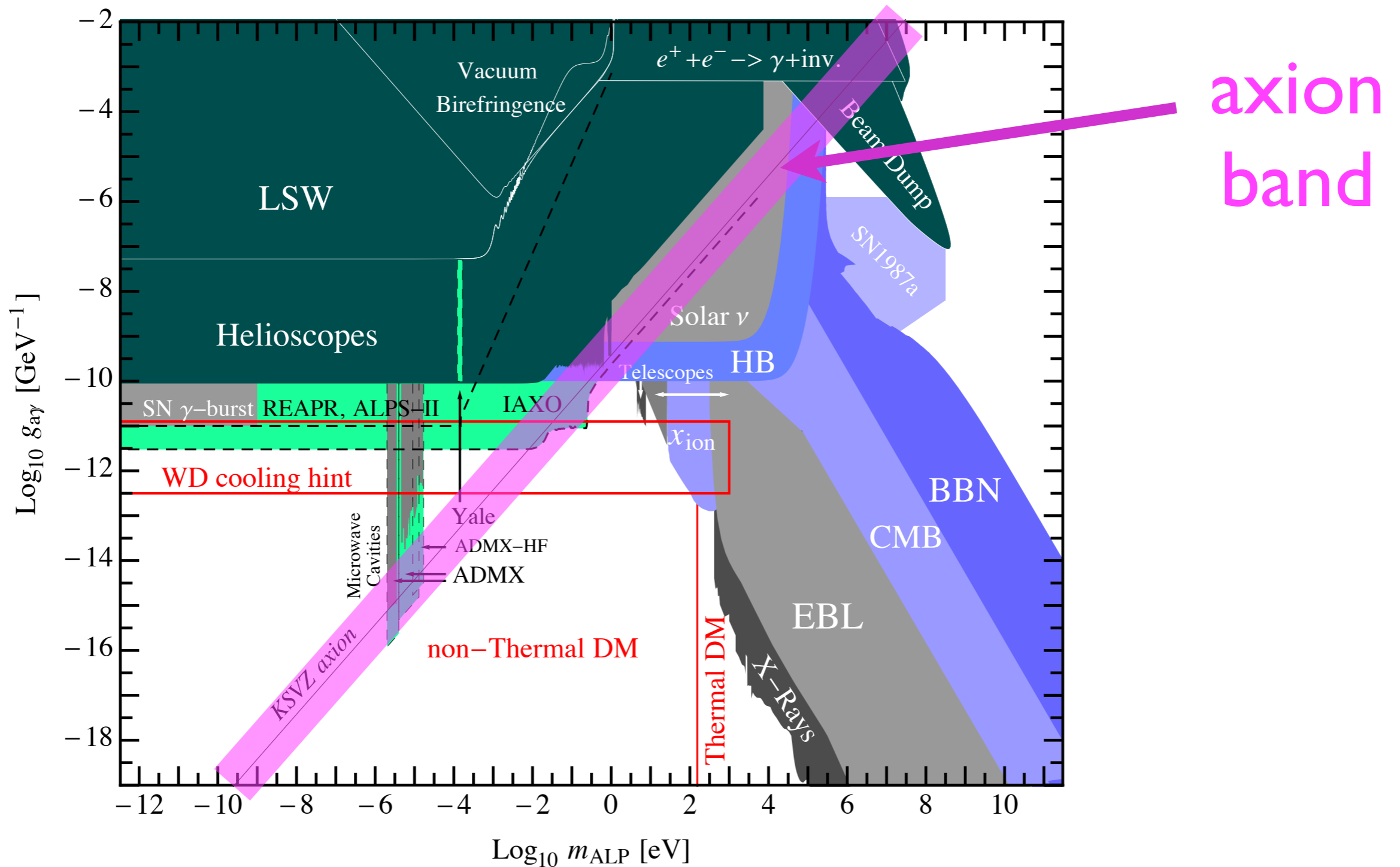
tunable
Resonant
Cavity

ADMX, ADMX-HF, ...

see talk by K. van Bibber on ADMX

Current Limits & Opportunities

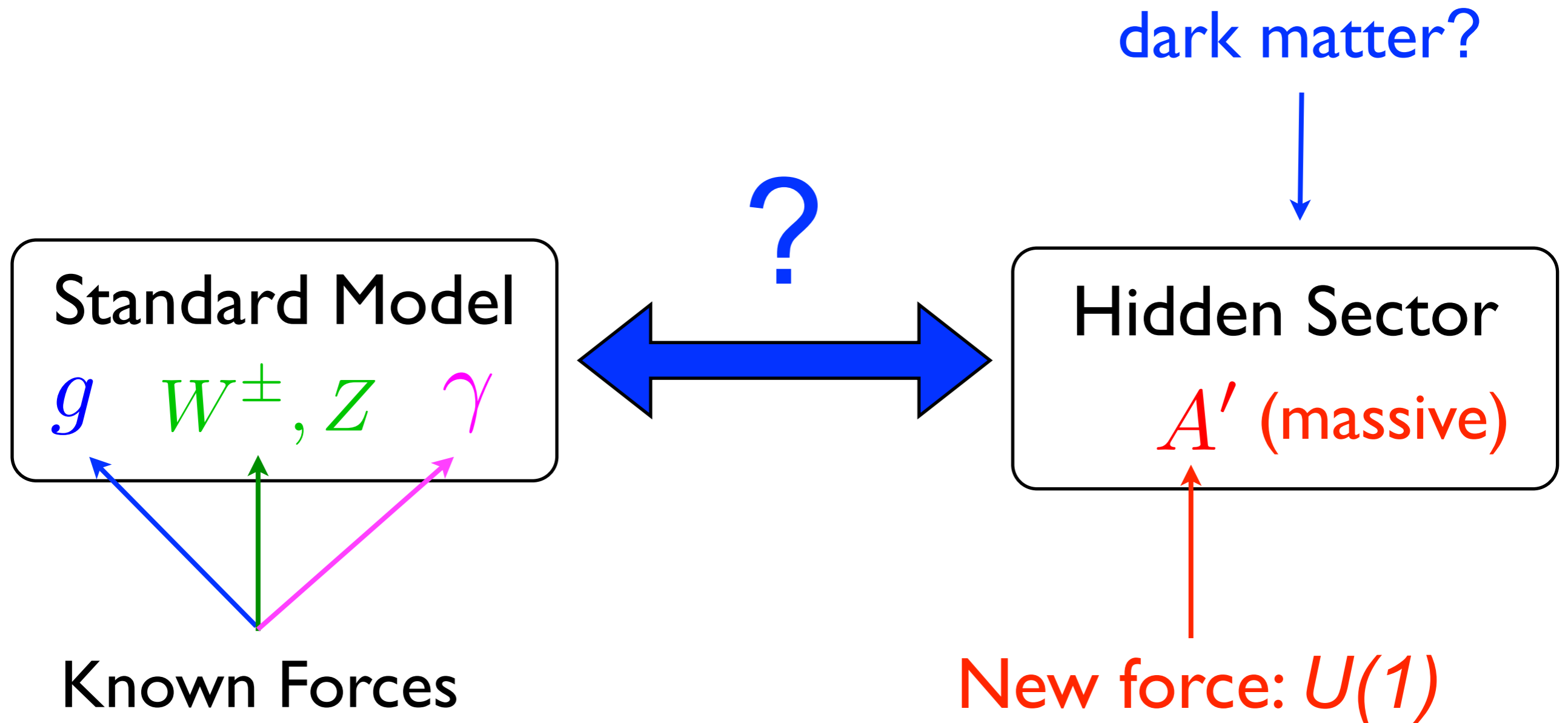
Axions & ALPs



Outline

- axions & axion-like particles
 - heavy photons
 - direct detection of sub-GeV DM
- 

Heavy Photons

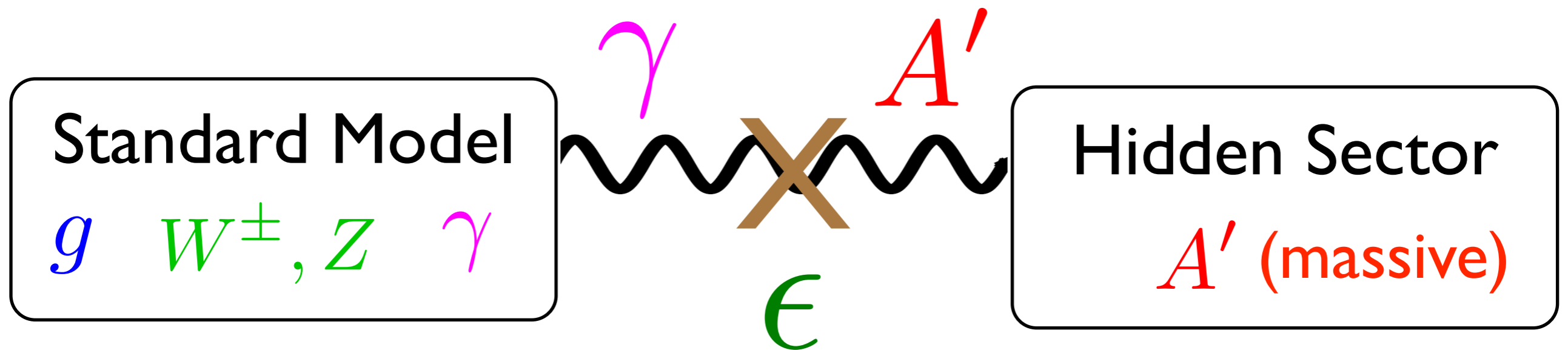


“Hidden Valleys”

Strassler,
Zurek, et.al.

Heavy Photons

ordinary photon & A' can mix

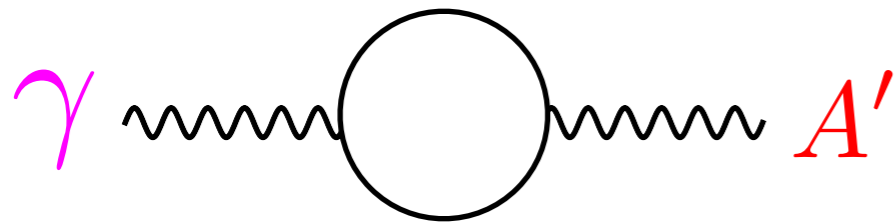


$$\Delta\mathcal{L} = \frac{\epsilon}{2} F^{Y,\mu\nu} F'_{\mu\nu} \quad \text{“Kinetic Mixing”}$$

Generating Kinetic Mixing

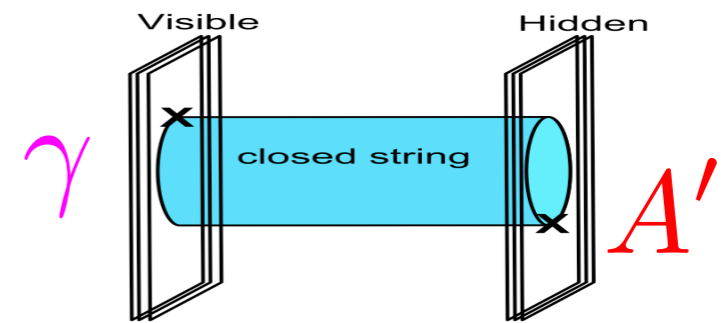
Examples:

loops of heavy particles



$$\epsilon \sim 10^{-8} - 10^{-2}$$

string theory



ϵ much smaller is possible

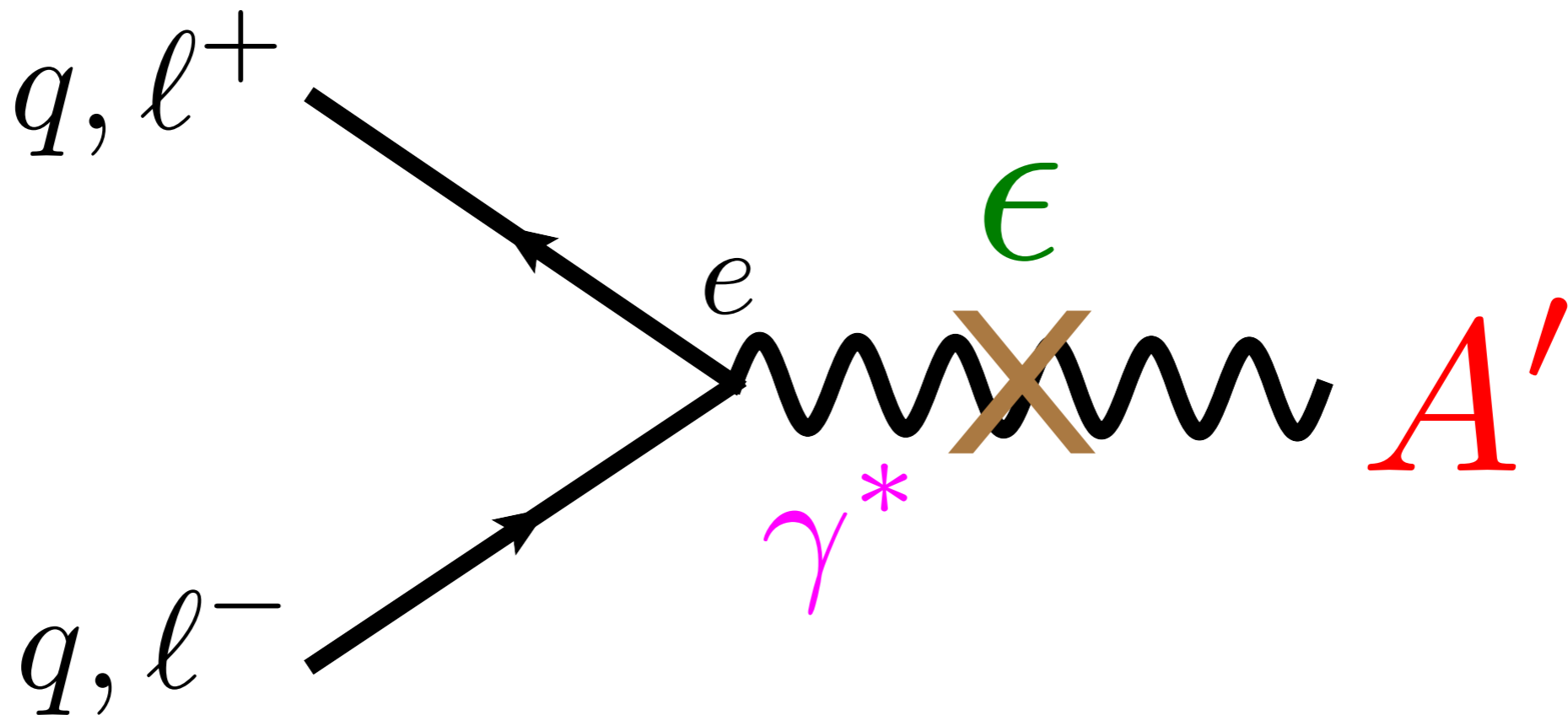
e.g. Goodsell, Jaeckel, Redondo, Ringwald

Mixing with photon allows:

$A' \leftrightarrow \gamma$ “oscillation”

and

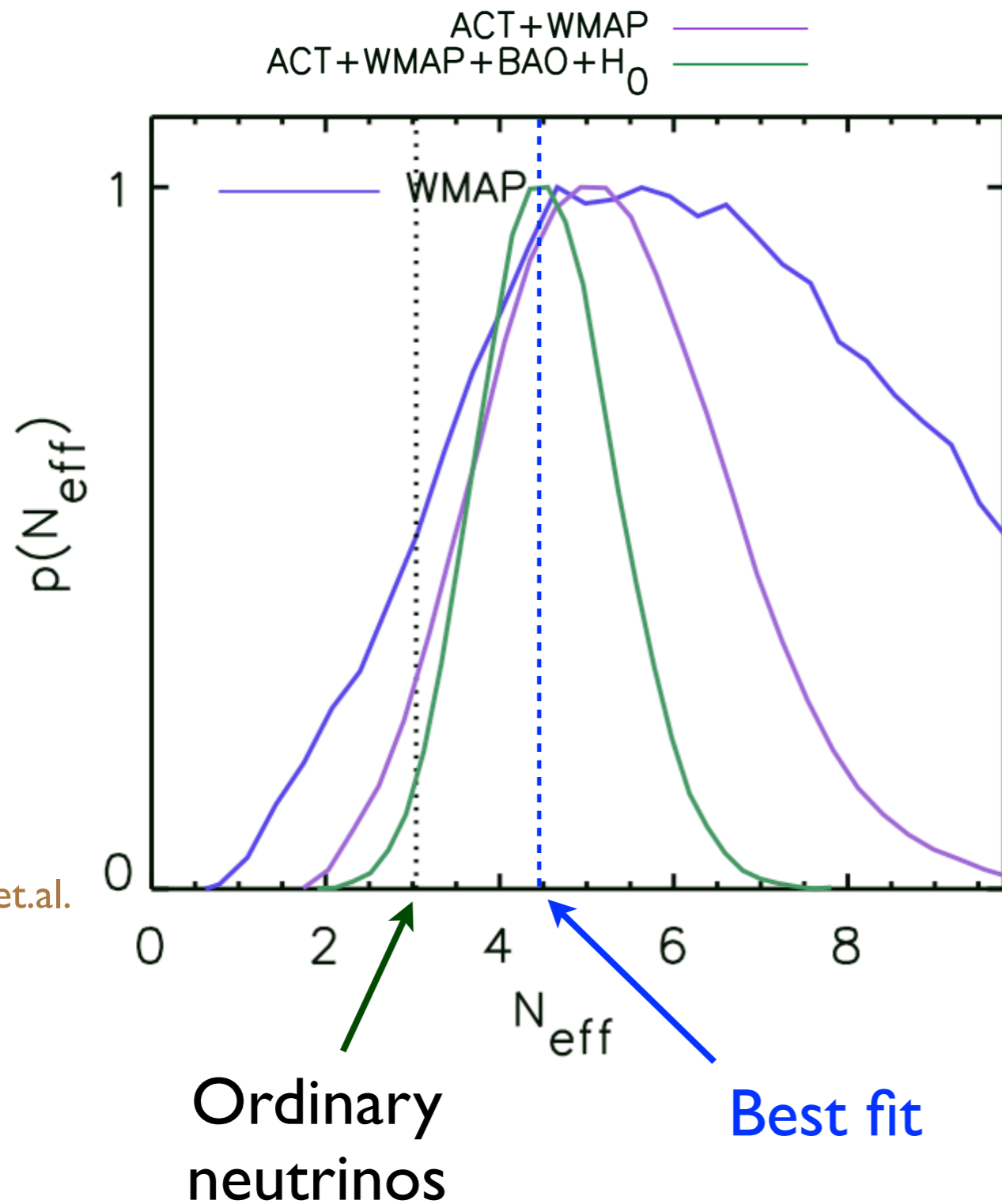
A' coupling to quarks and charged leptons:



A' WISP:

mass $< 2m_e \sim \text{MeV}$

A' WISP hints from N_{eff} from CMB?



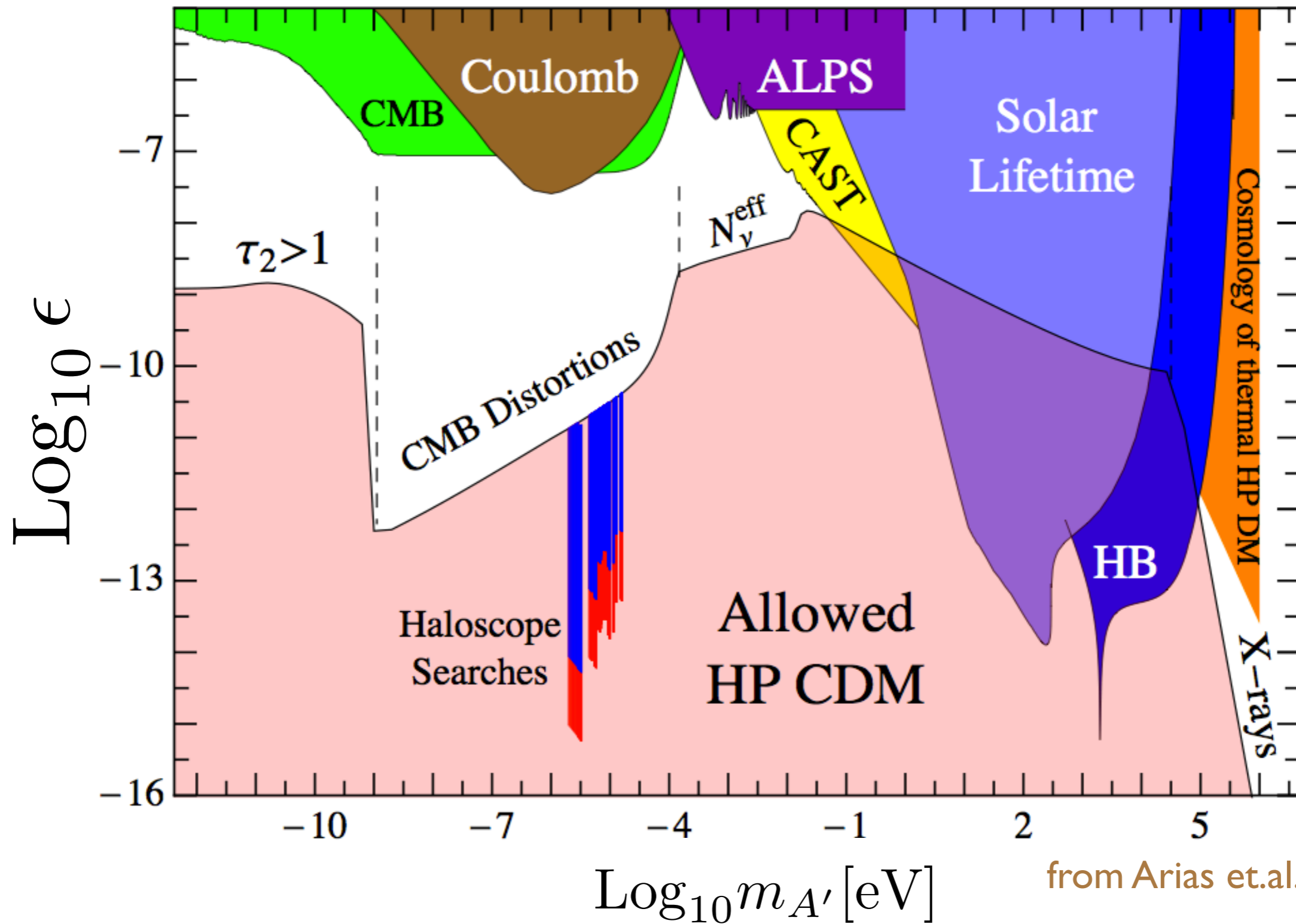
suggestive of
additional relativistic
degrees of freedom
(N_{eff})?

$$m_{A'} \sim \text{meV} ?$$

Jaeckel, Redondo, Ringwald

Dunkley et.al.
(ACT)

A' WISPs as Dark Matter

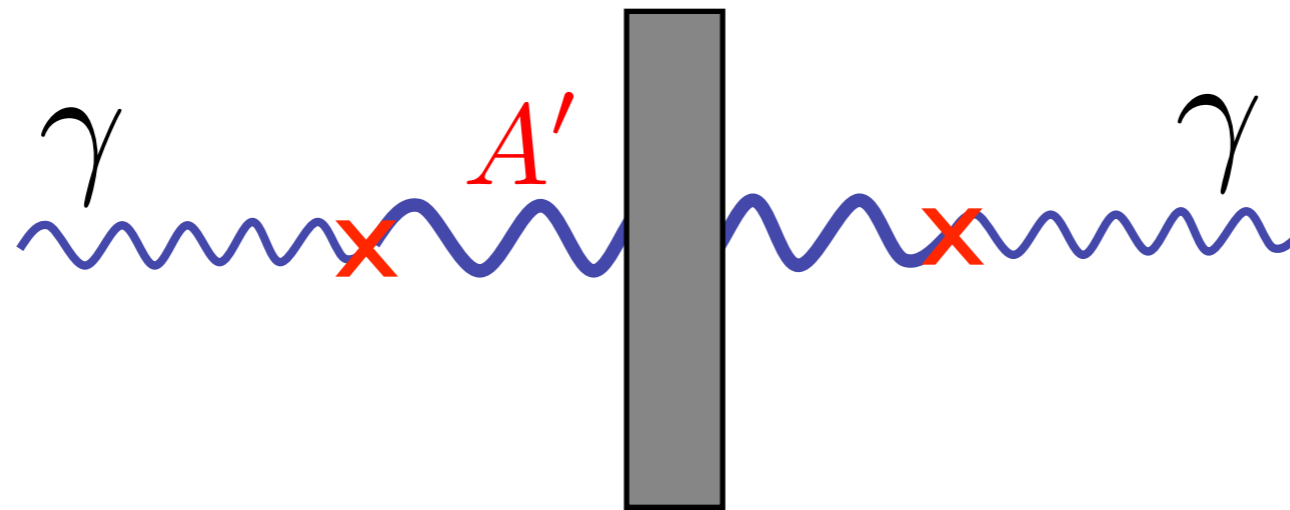


see talk by J. Redondo

How look for A' WISPs?

“Light-shining-through-walls”

(cf. axions)



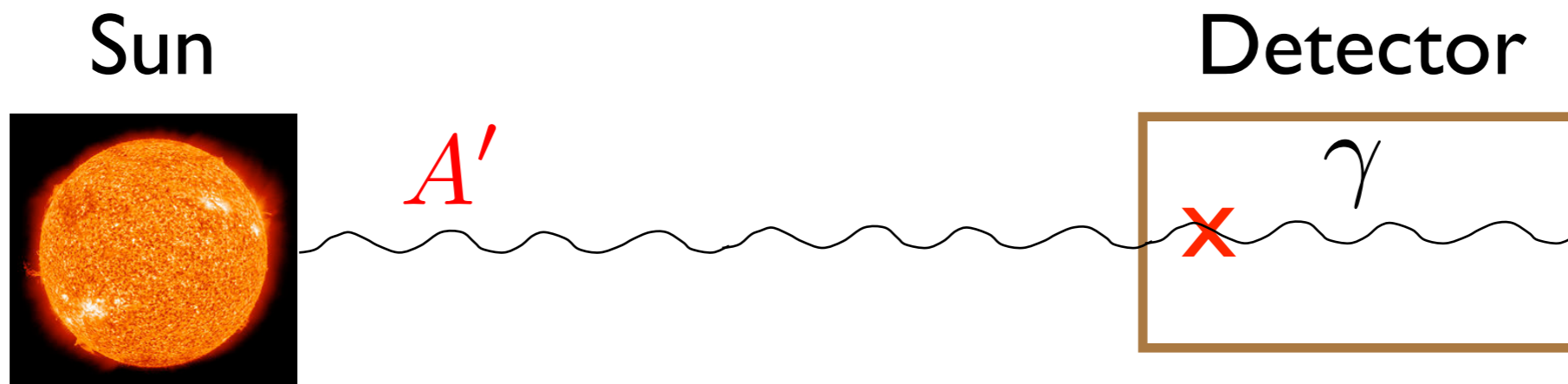
LIPSS (Jlab) , BFRT (BNL), BMV (LULI), GammeV (Fermilab),
ALPS (DESY), OSQAR (CERN), PVLAS (INFN), ...

Need powerful lasers but no magnets

How look for A' WISPs?

Helioscopes: stare at the sun (cf. axions)

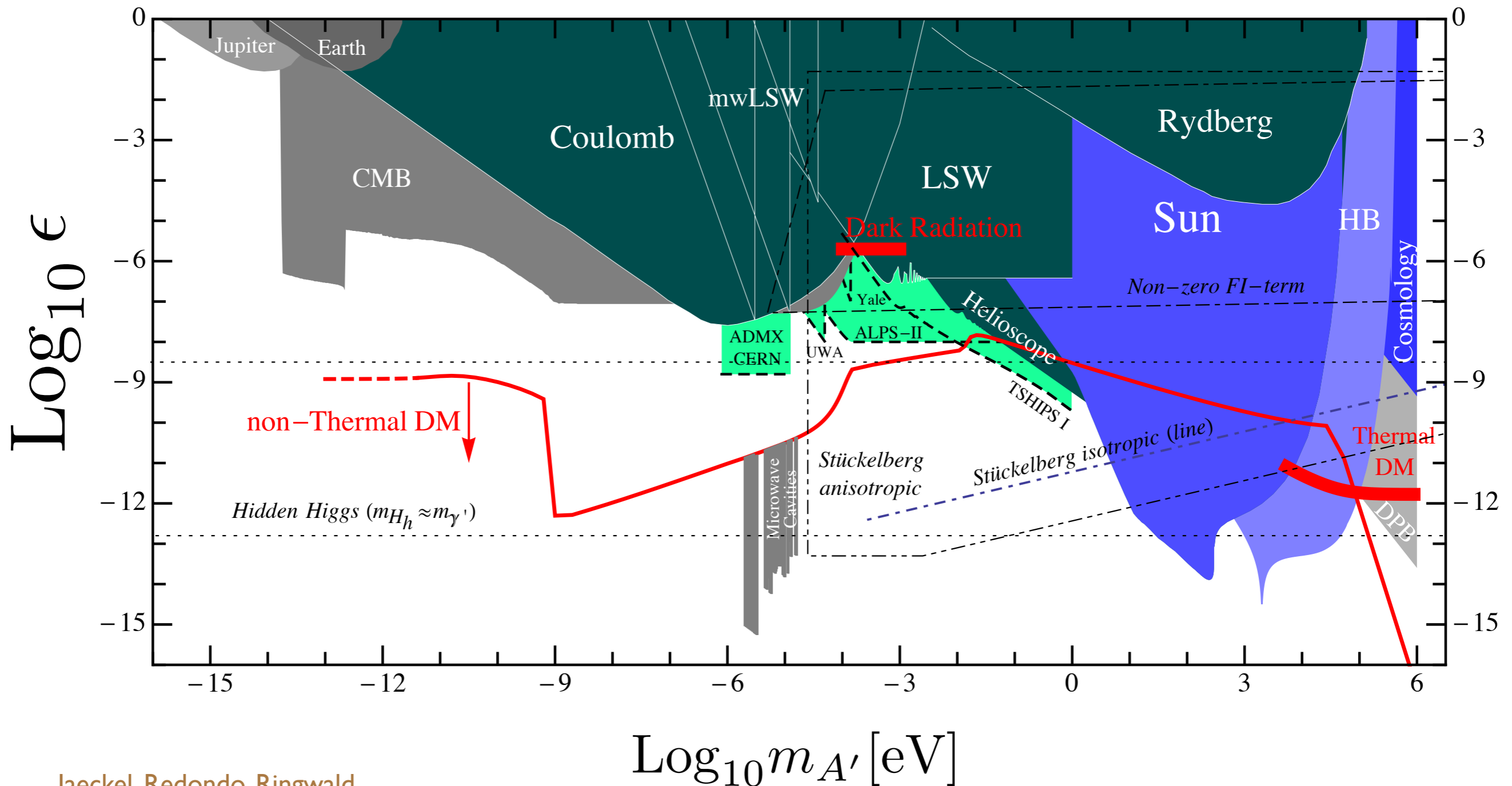
Okun, ...



TSHIPS, CAST, SUMICO, IAXO, ...

see talk by M. Schwarz

A' WISP parameter space



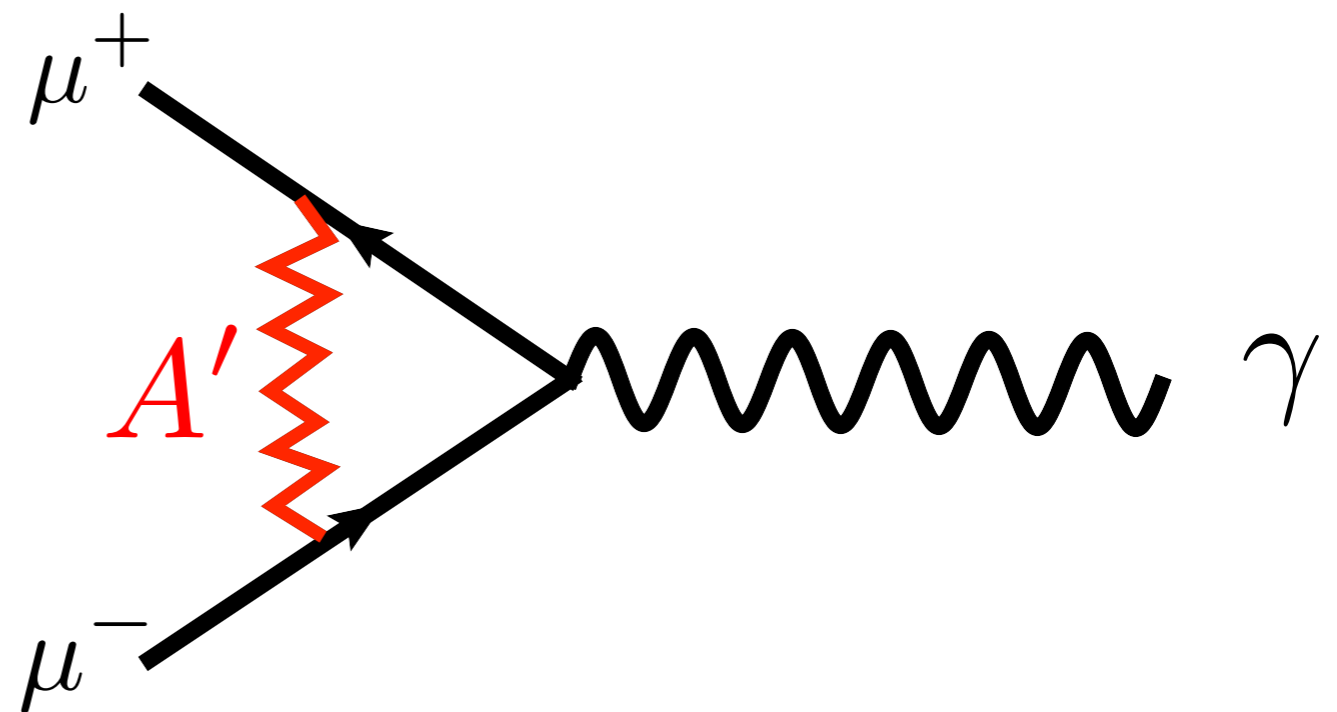
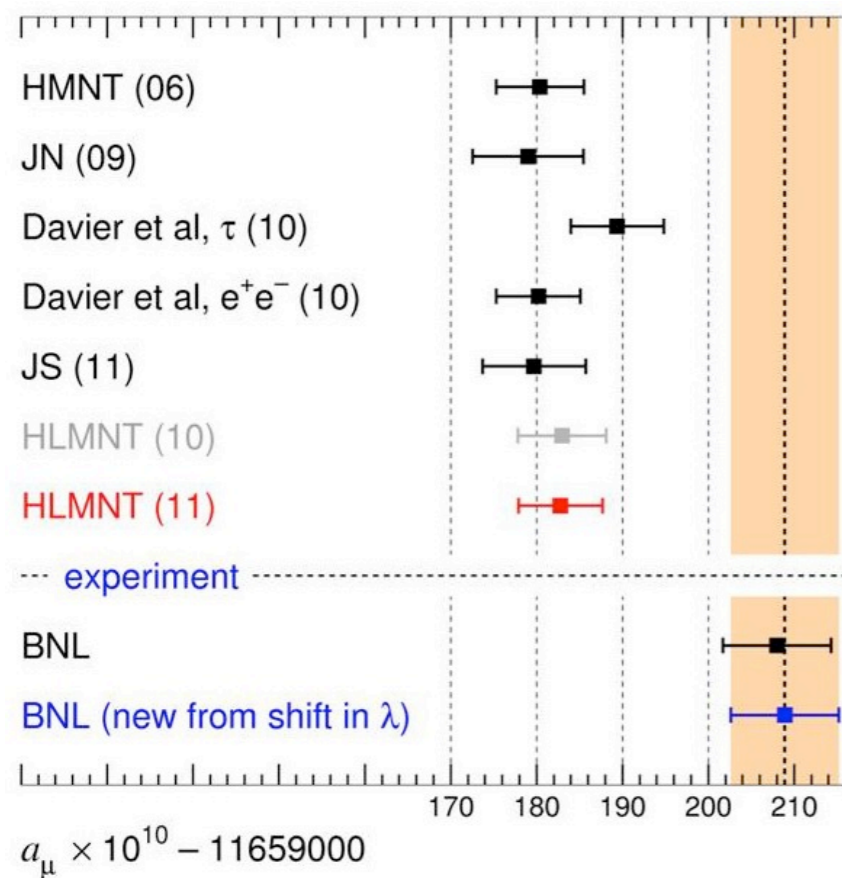
A' :

mass \sim MeV-GeV

Hints for A' with MeV-GeV mass?

anomalous muon $g-2$?

Pospelov
Boehm, Fayet

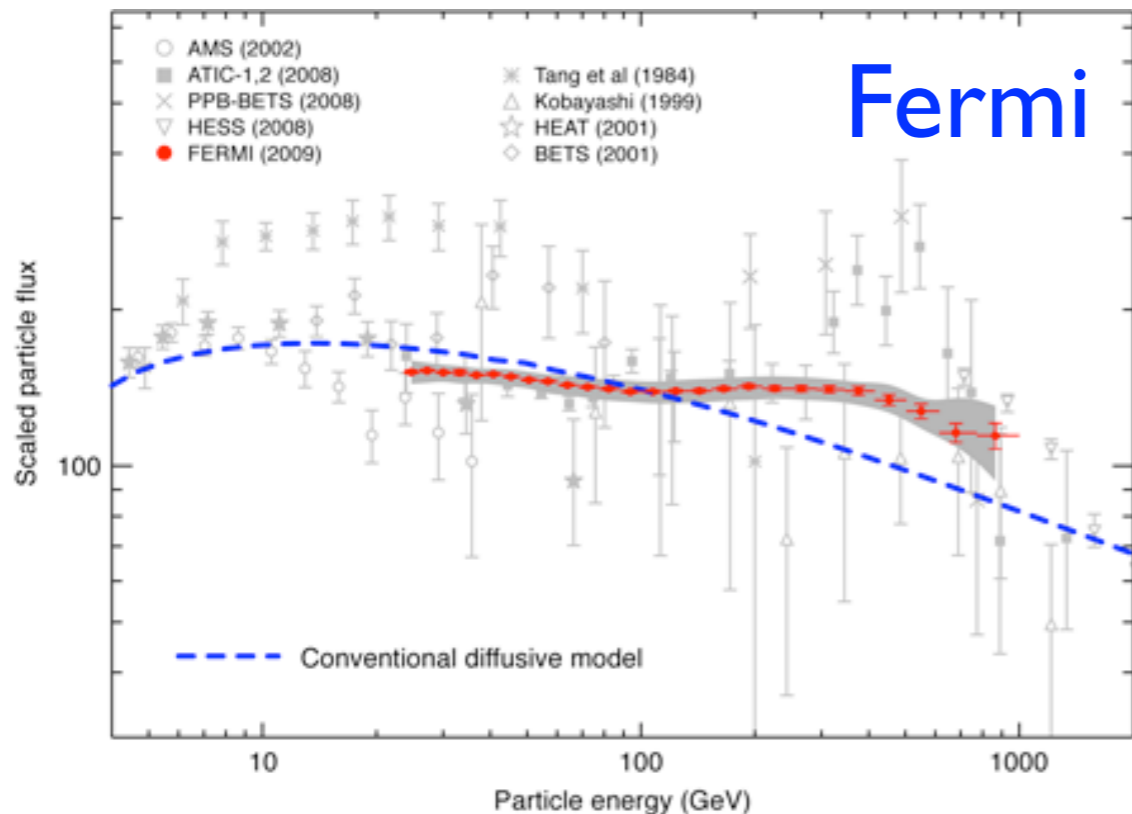
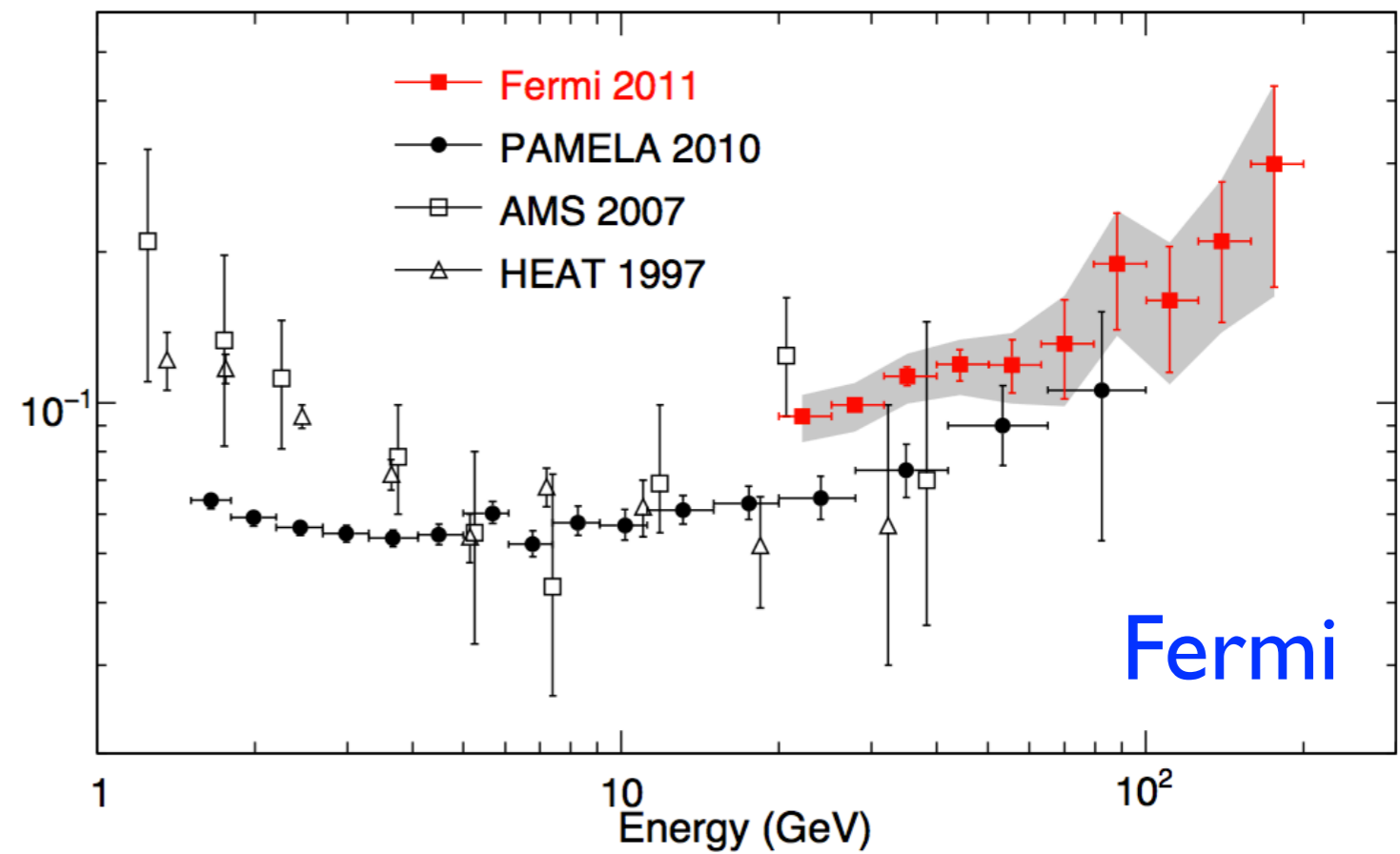
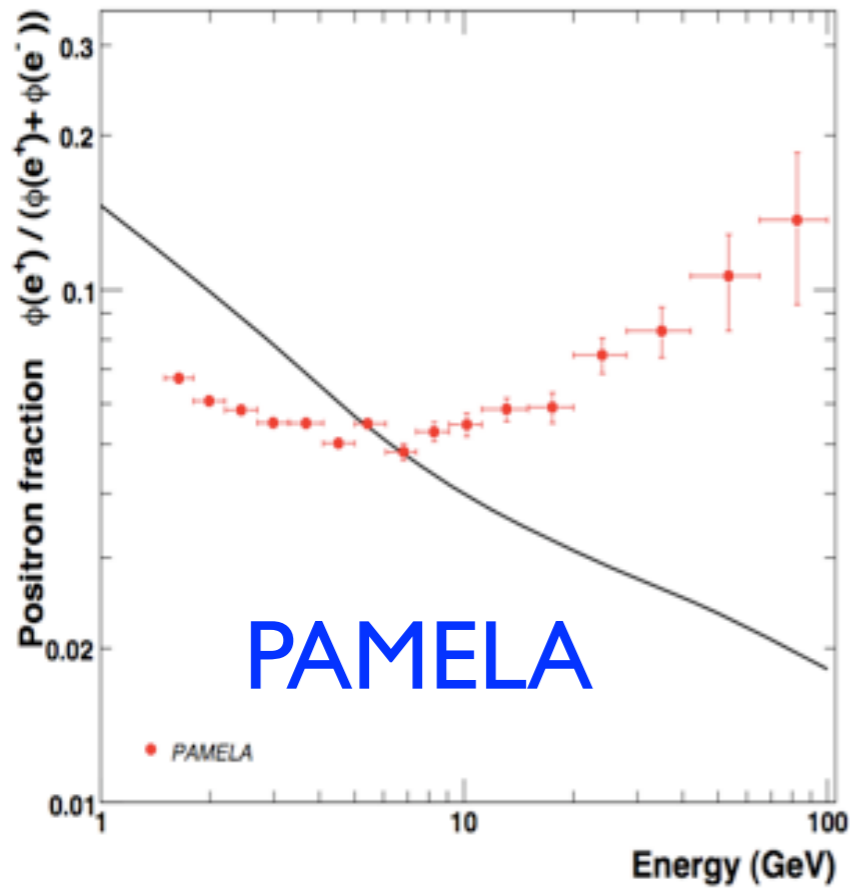


A' may explain observed $(g_s - 2)_\mu$!

Hints for A' with MeV-GeV mass?

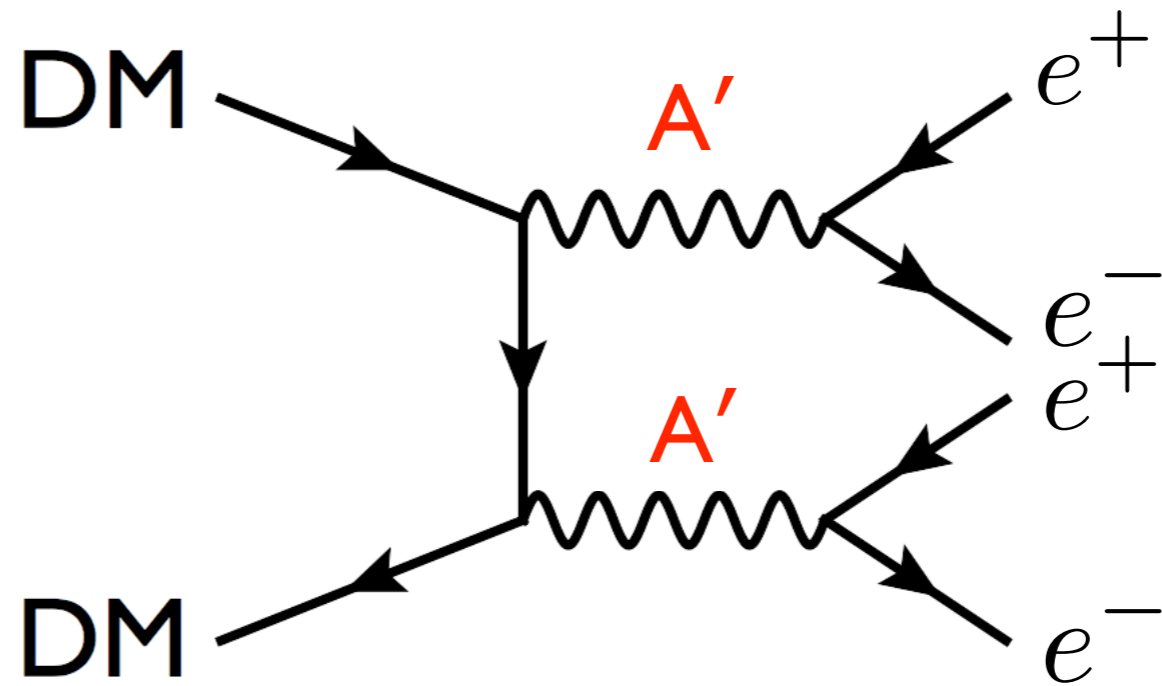
New dark matter interactions?

Many, many DM related anomalies



cosmic-ray e^+ & e^- excess
but no anti-proton excess
(restricts DM interactions)

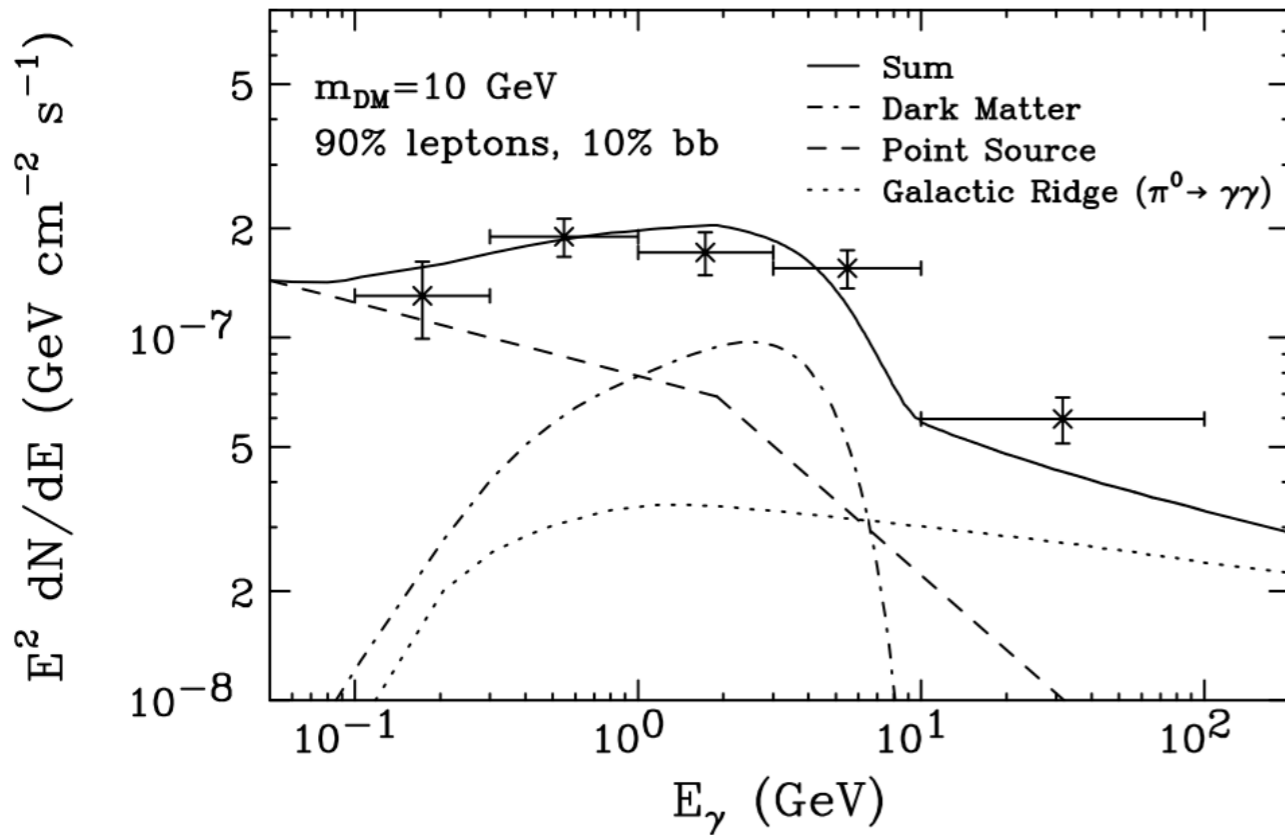
100 GeV - 1 TeV



produces e^+ , no \bar{p}

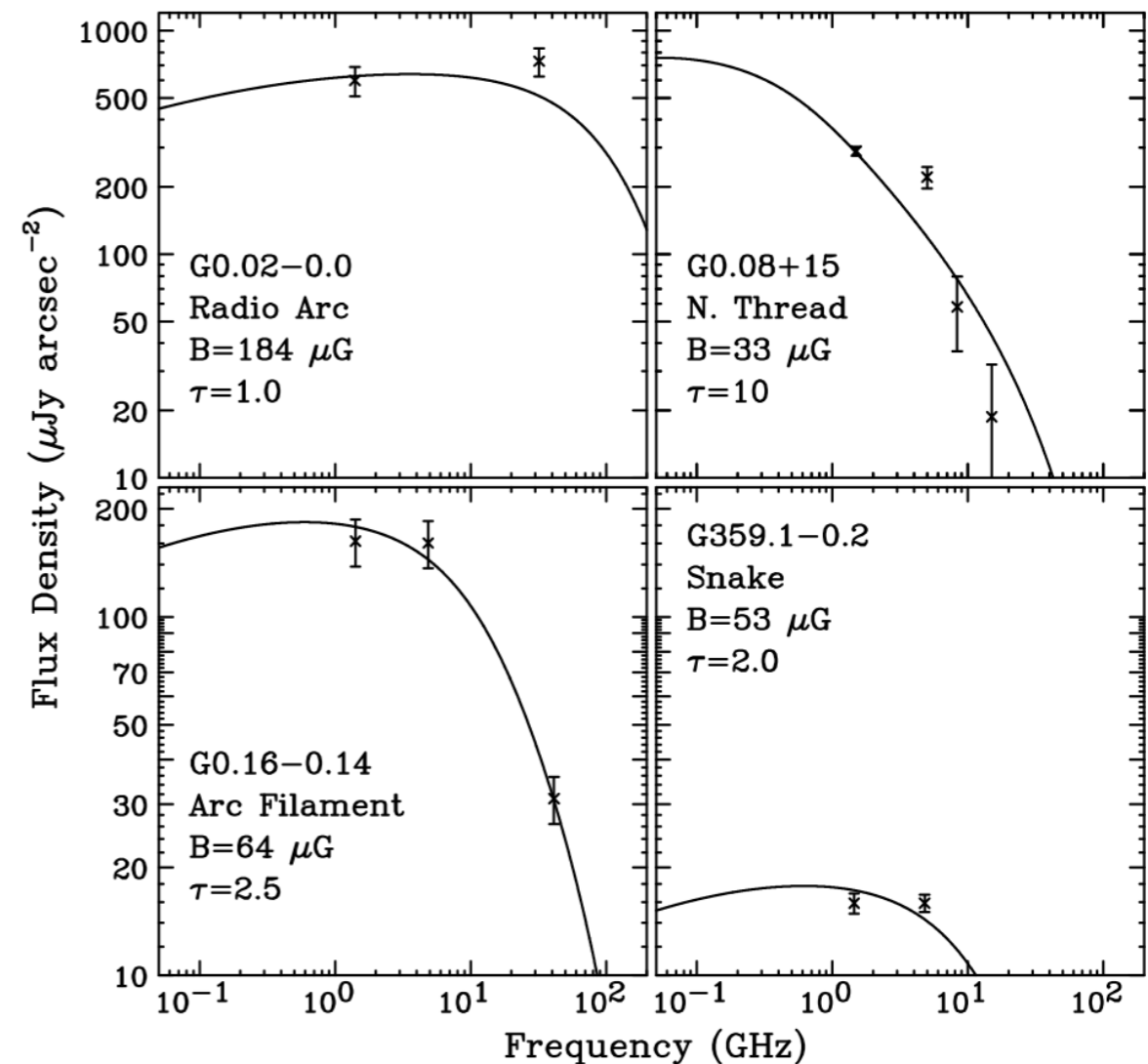
Many, many DM related anomalies

Fermi photon spectrum near GC

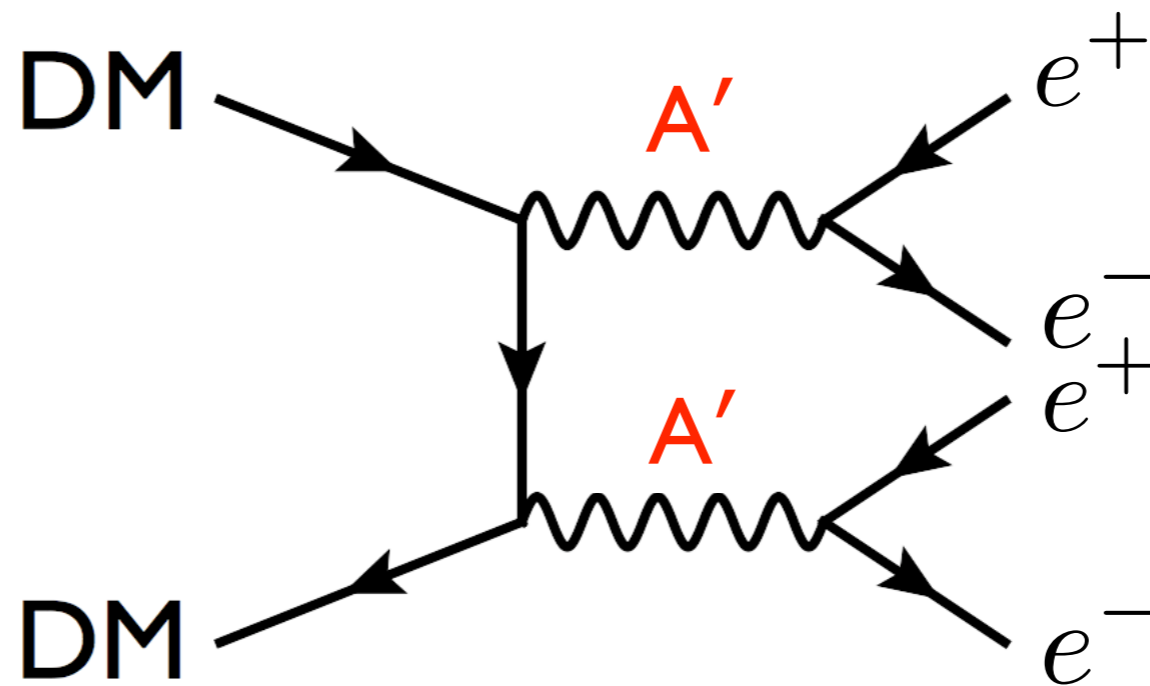


see talk by D. Hooper

Synchrotron emission from inner galaxy's radio Filaments



~10 GeV



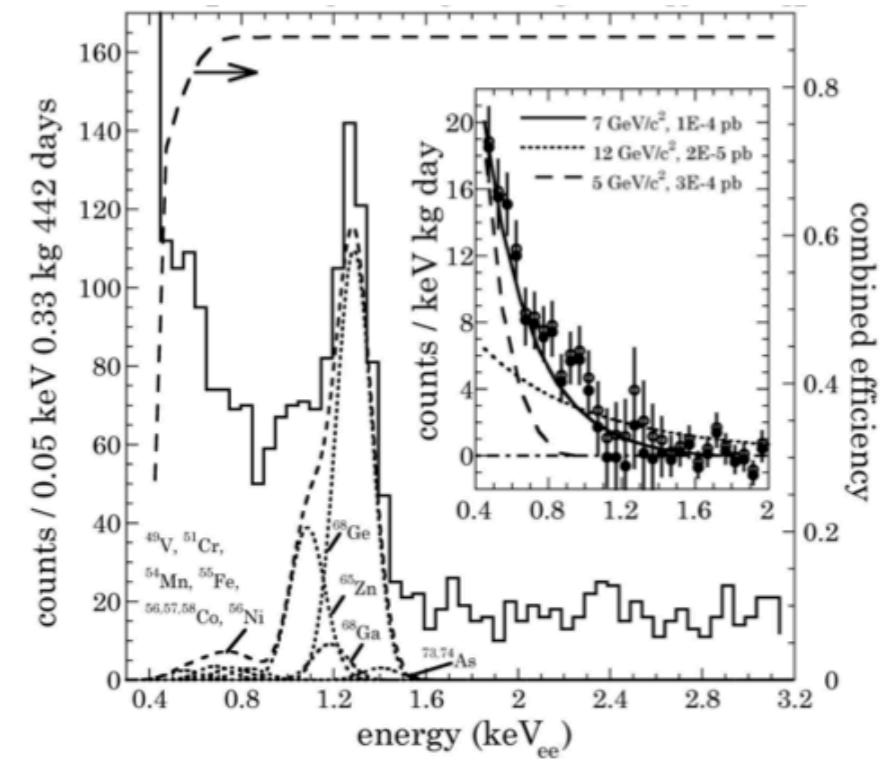
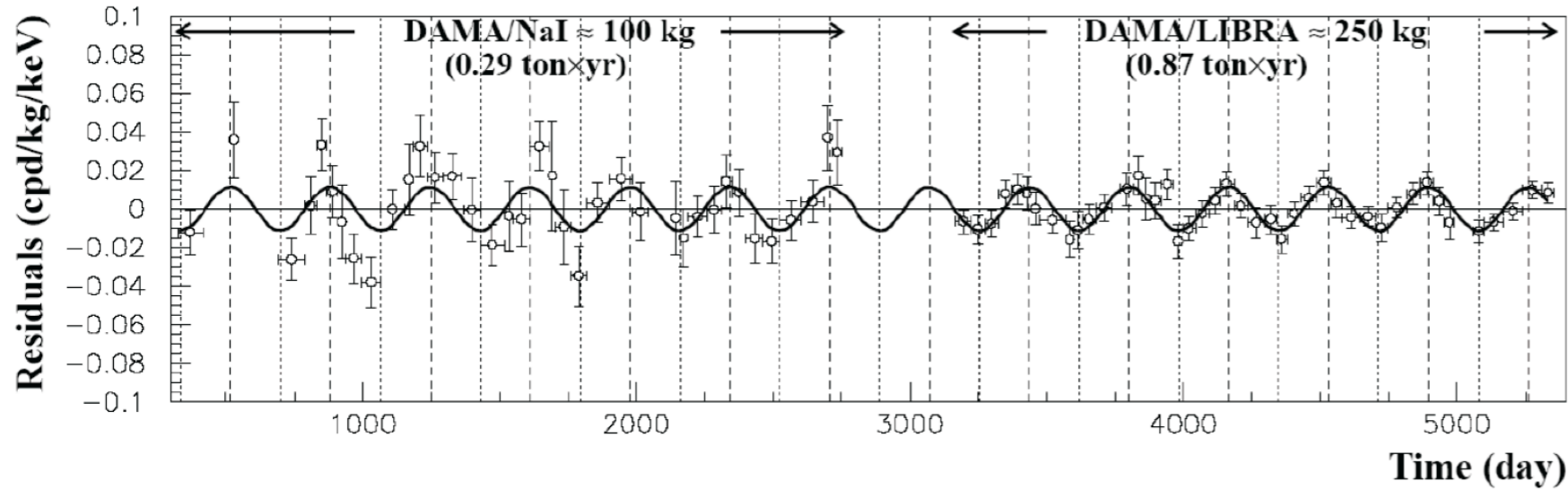
produces e^+

Hooper, Weiner, Xue

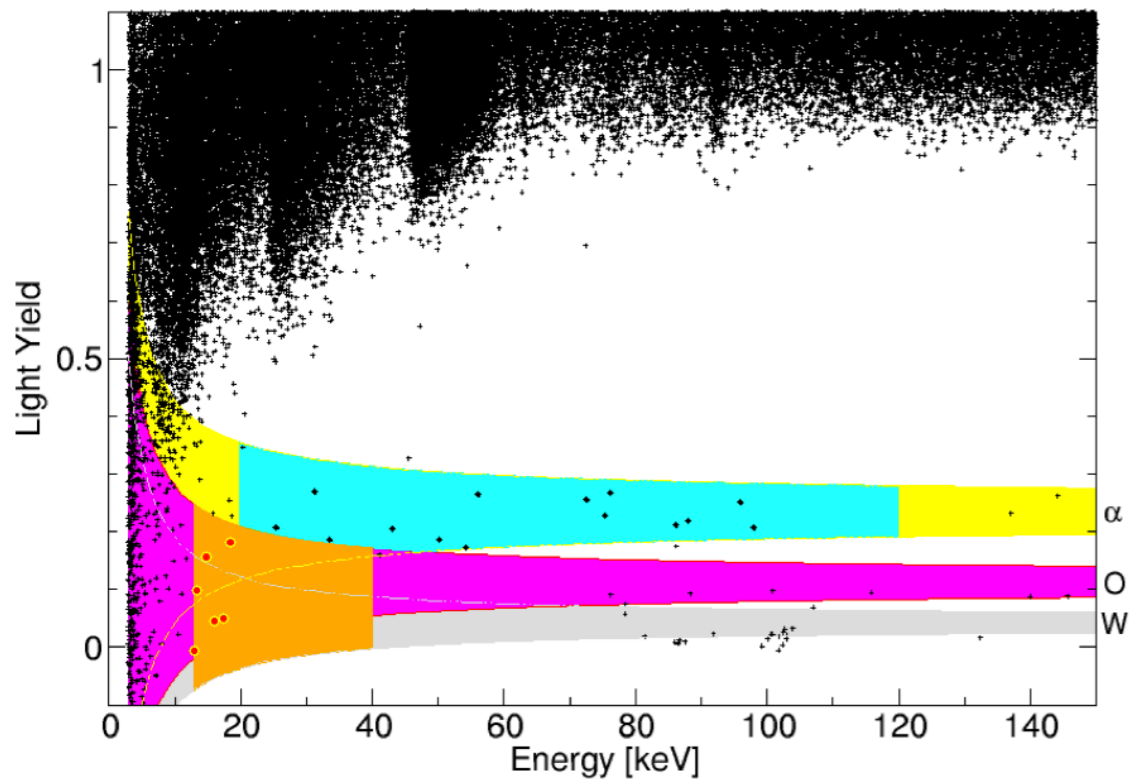
see talk by D. Hooper

Many, many DM related anomalies

DAMA



CRESST

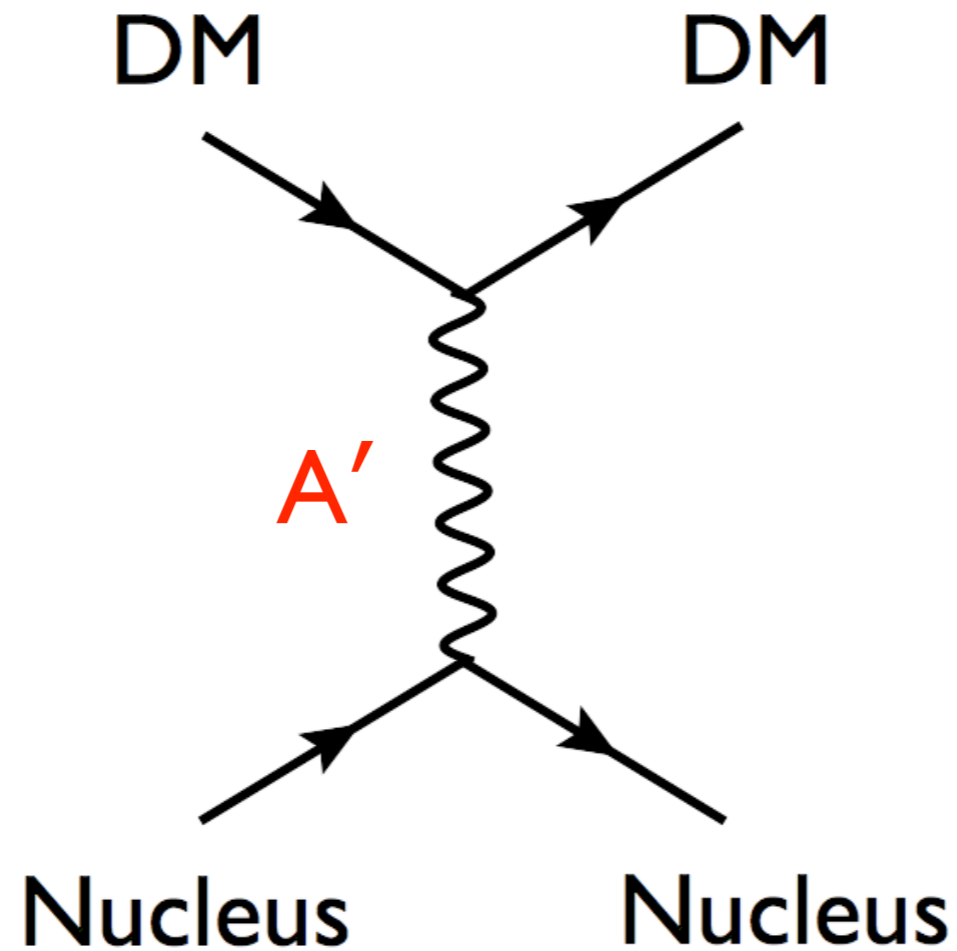


CoGeNT

see talk by J. Collar

direct detection anomalies

~ 10 GeV



e.g.

Arkani-Hamed et.al.;
Cheung, Ruderman, Wang, Yavin;
Morrissey, Poland, Zurek;
Cline, Frey;
Hooper, Weiner, Xue
RE, Kaplan, Schuster, Toro;

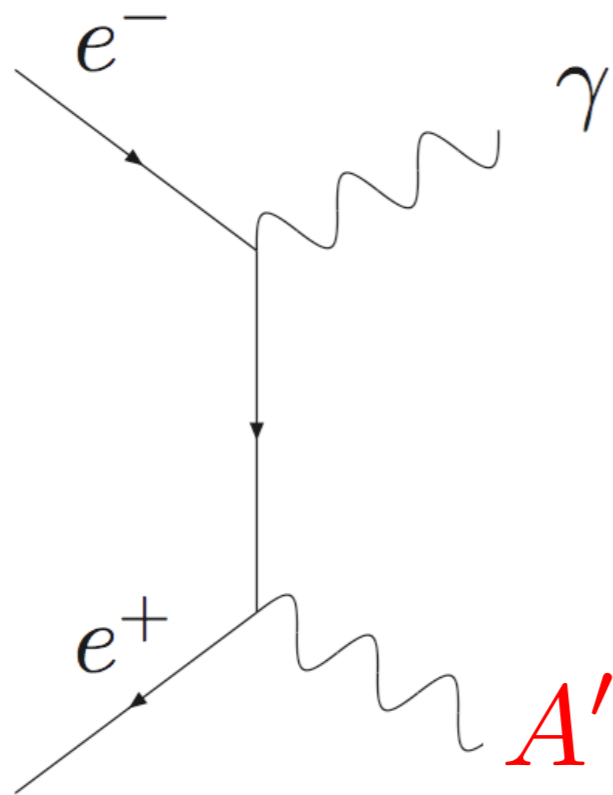
direct detection hints?

see talks by S.Andreas, D. Hooper

How look for A' with MeV-GeV mass?

RE, Schuster, Toro
Batell, Pospelov, Ritz
Reece, Wang
Borodatchenkova et.al.
Fayet

e^+e^- colliders



Rare meson decays

$$\phi \rightarrow \eta A'$$

$$\pi^0 \rightarrow \gamma A'$$

B-factories, Phi-factories

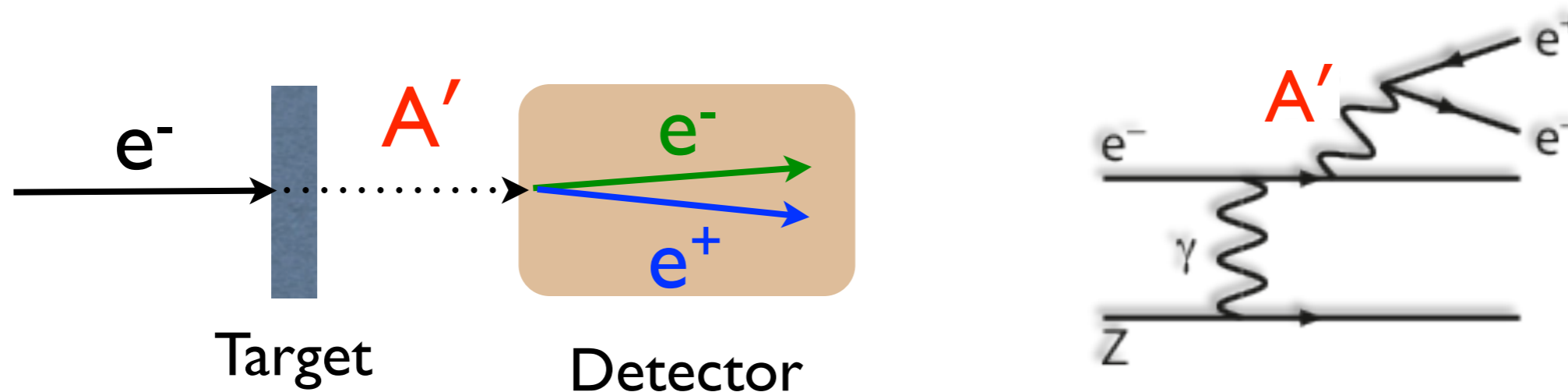
Searches ongoing

see talks by B. Echenard, Y. Liu

How look for A' with MeV-GeV mass?

Bjorken, RE, Schuster, Toro
Freytsis, Ovanesyan, Thaler
Reece & Wang

New e^- fixed target experiments



Experiments done/planned at

- Jefferson Lab (APEX, HPS, DarkLight)
- Mainz
- VEPP-3

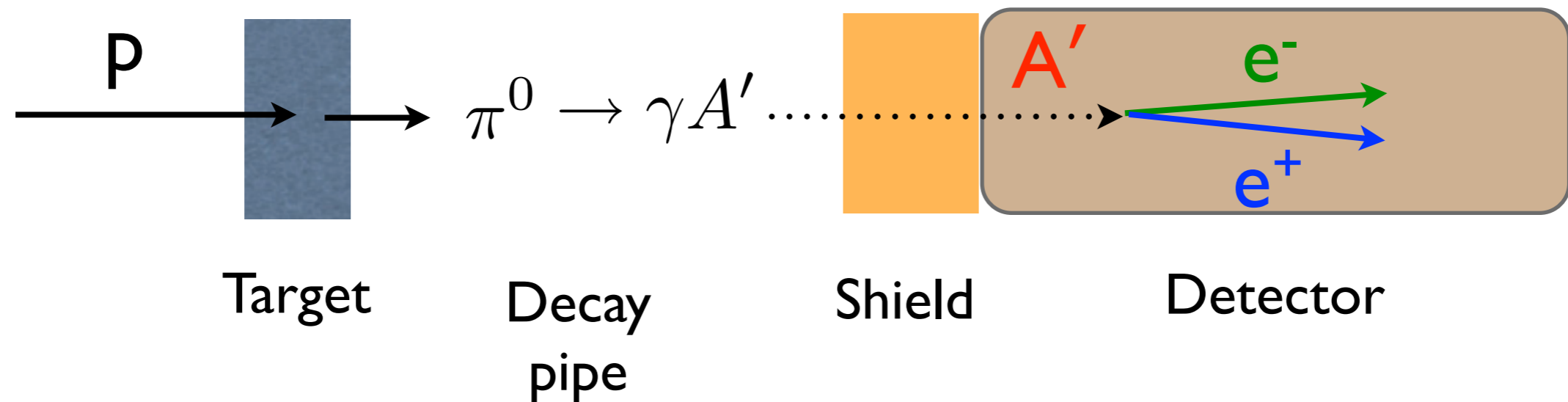
see talks by
J. Beacham (APEX)
P. Hanssen (HPS)
J. Boyce (JLab)

How look for A' with MeV-GeV mass?

Proton-beam fixed target experiments

Batell, Pospelov, Ritz
RE, Harnik, Kaplan, Toro

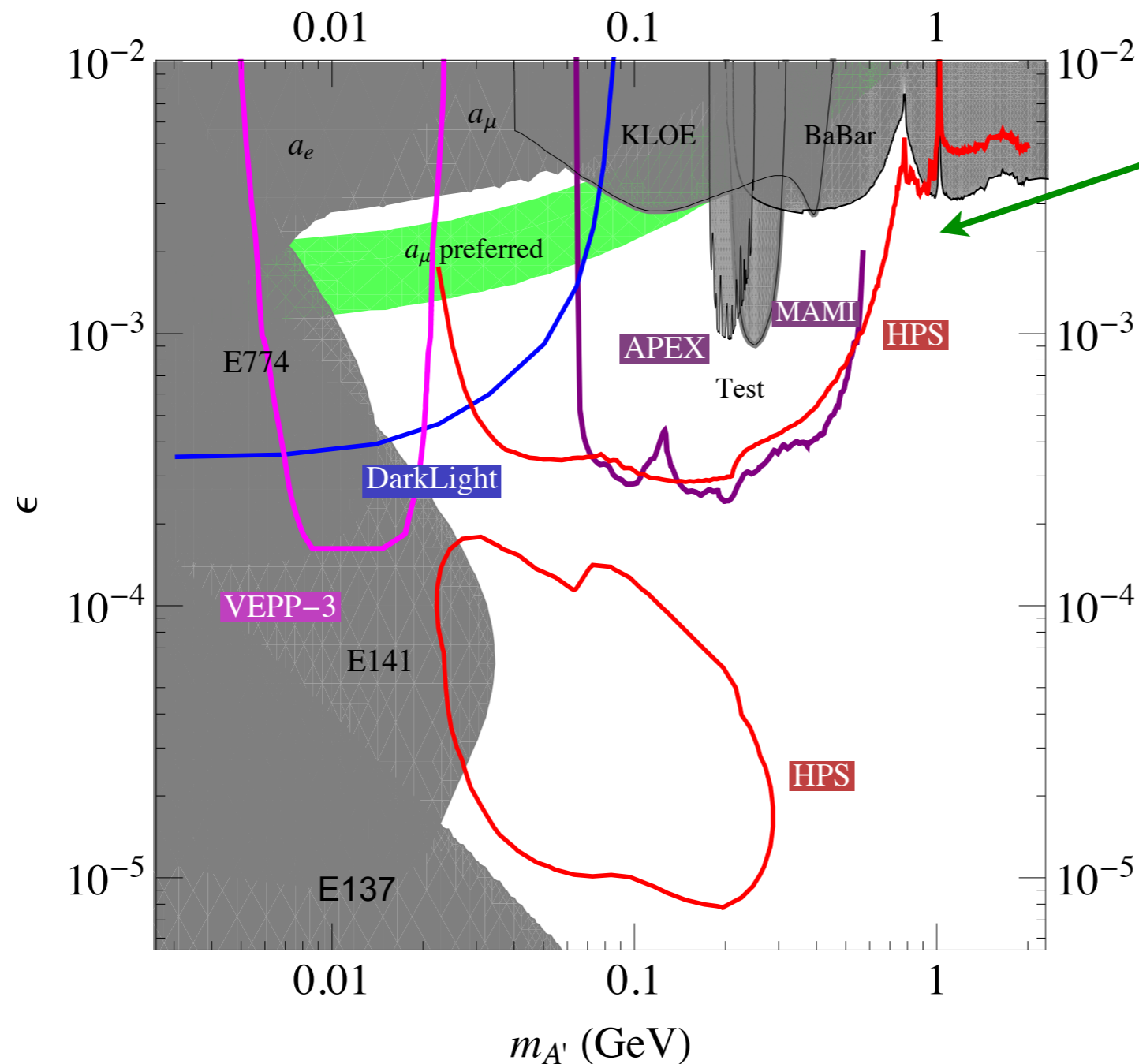
Example: produce A' from pion decays



e.g. LSND, MINOS, MiniBooNE, Project X

see talks by B. Batell, C. Wallace

MeV-GeV A' constraints & prospects



B/Phi-factories
can significantly
extend reach

MAMI has
plans to cover
similar region
to APEX

region motivated by theory, dark matter, muon $g-2$

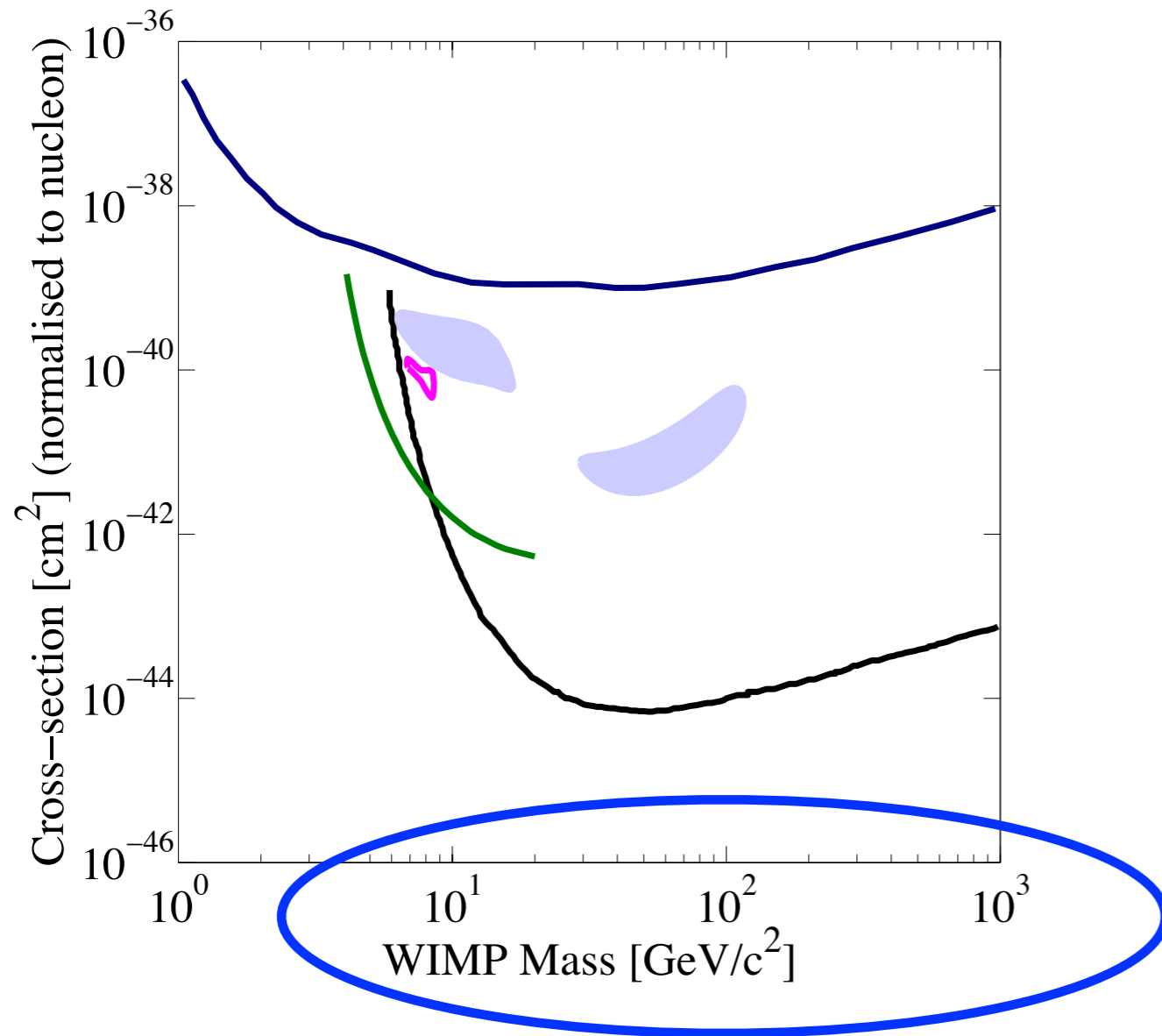
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see talk by A. Manalaysay

The usual DM direct detection parameter space

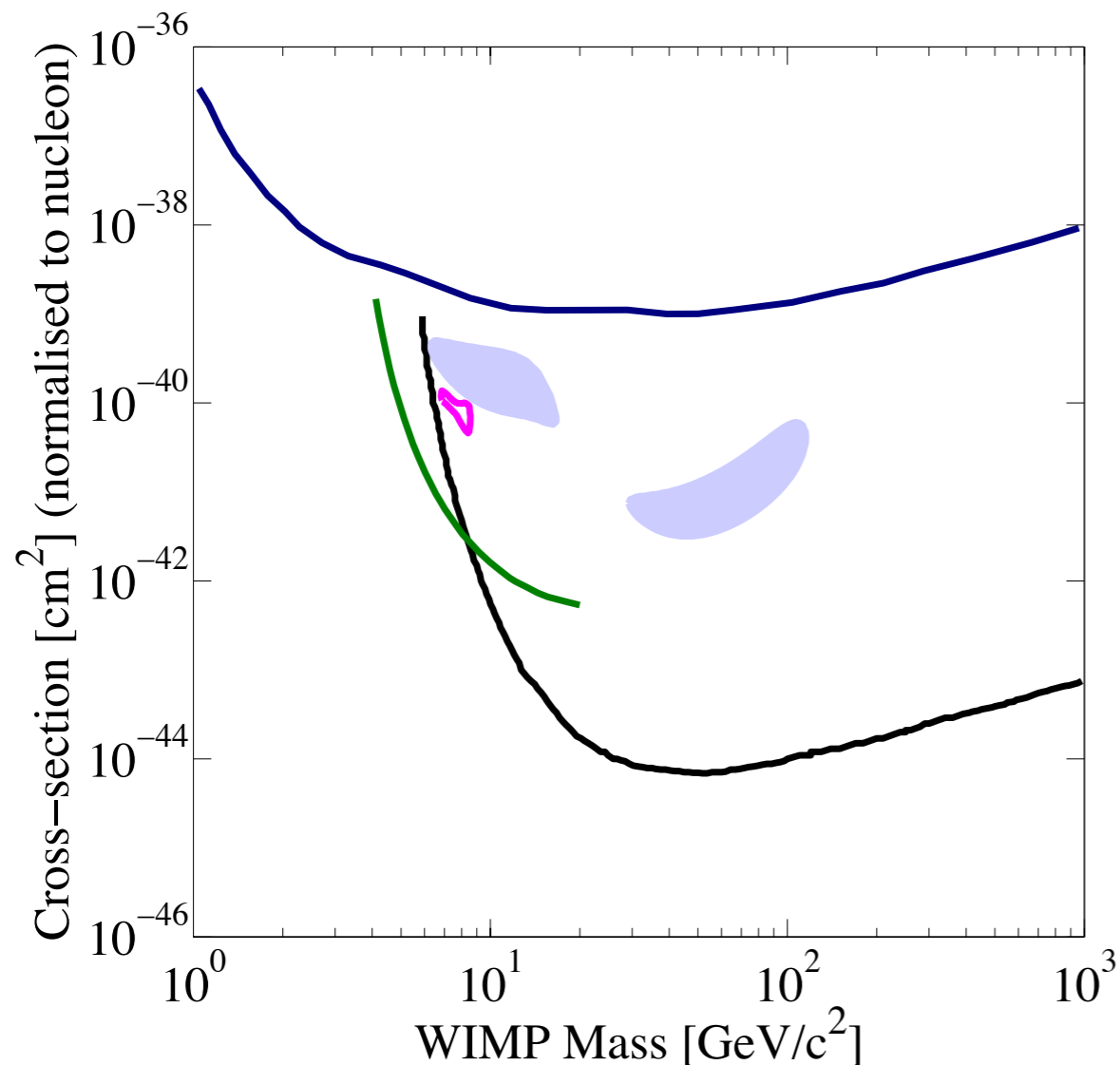


well-motivated by the
WIMP miracle

See *many* talks: J. Collar, P. Gondolo, W. Guo, C. Gustavo, K. Hiraide, B. Kilminster, V. Kozlov, B. Loer, D. McKinsey, R. Nelson, U. Oberlack, M. Schumann, E. Vazquez-Jauregui, U. Wichoski

Weak-scale

The usual DM direct detection parameter space

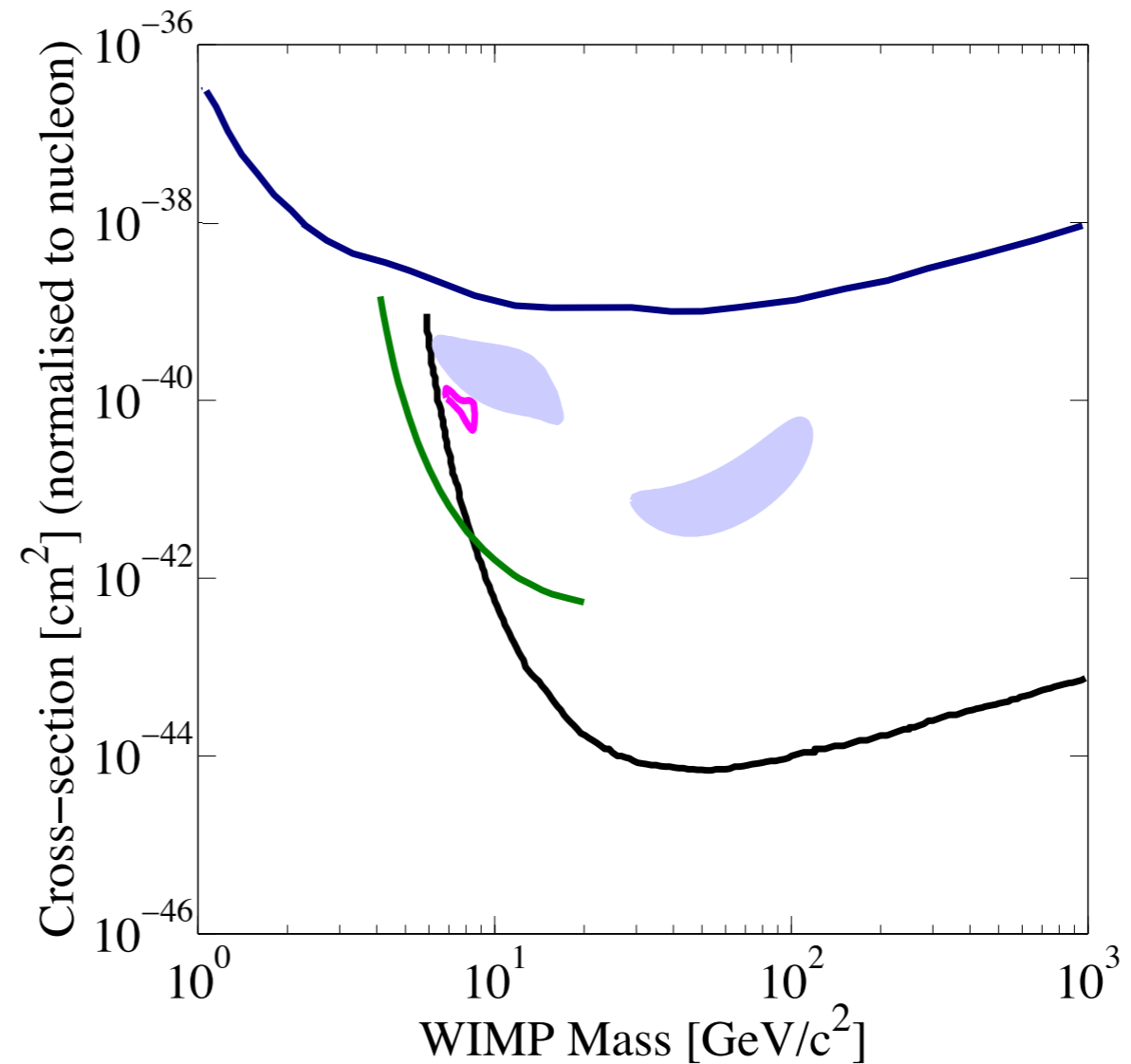


well-motivated by the
WIMP miracle

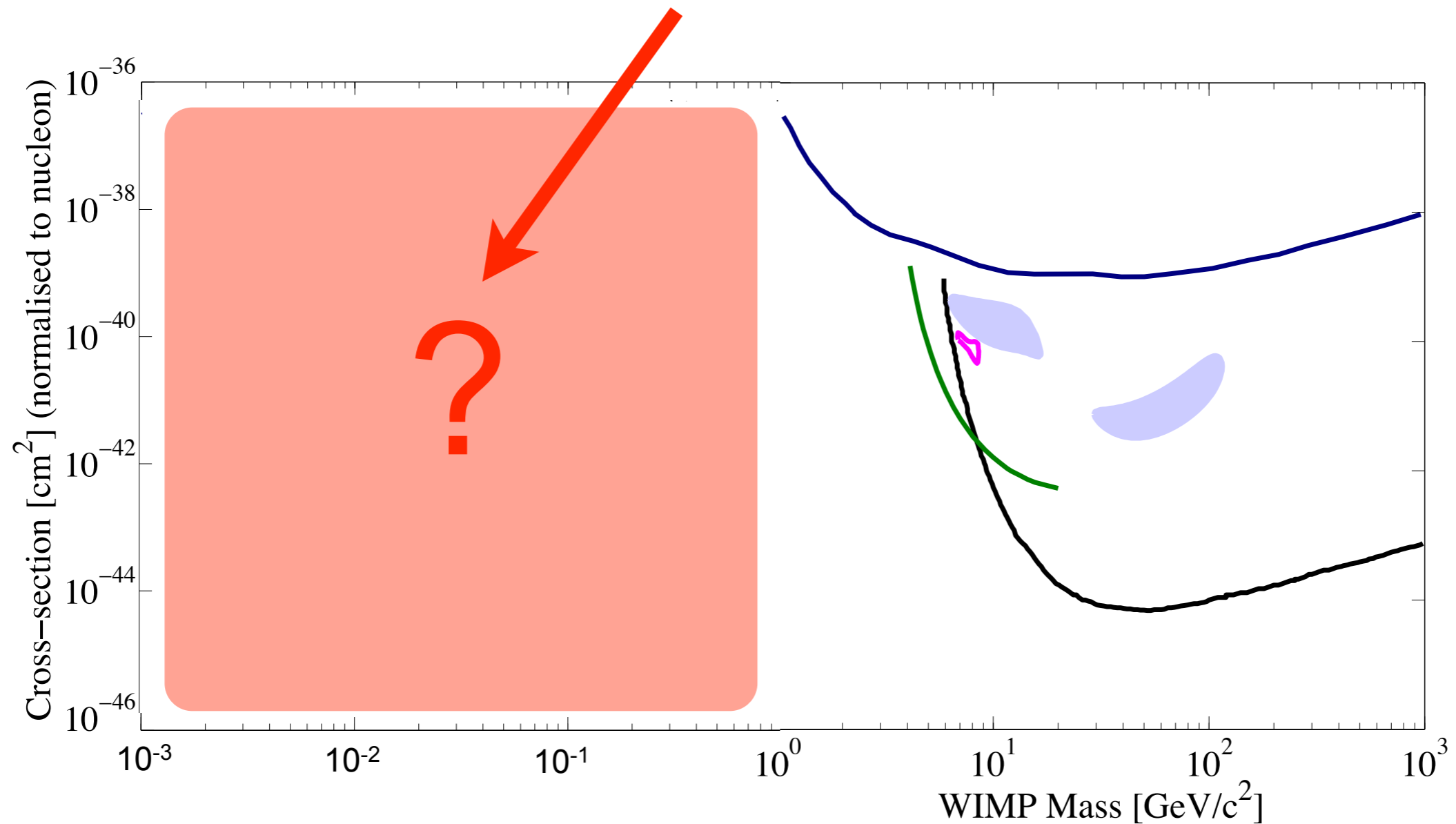
See *many* talks: J. Collar, P. Gondolo, W. Guo, C. Gustavo, K. Hiraide, B. Kilminster, V. Kozlov, B. Loer, D. McKinsey, R. Nelson, U. Oberlack, M. Schumann, E. Vazquez-Jauregui, U. Wichoski

BUT: don't let a paradigm blind you to other experimental opportunities!

Instead of considering only this...



What if DM is here?

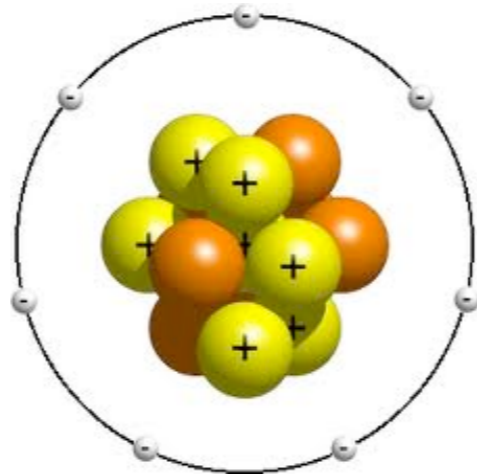


mass ~ MeV - GeV

Easy to build viable DM models
(e.g. interactions through an A'!)

Elastic nuclear recoils don't work

Heavy DM



Atom

Elastic nuclear recoils don't work

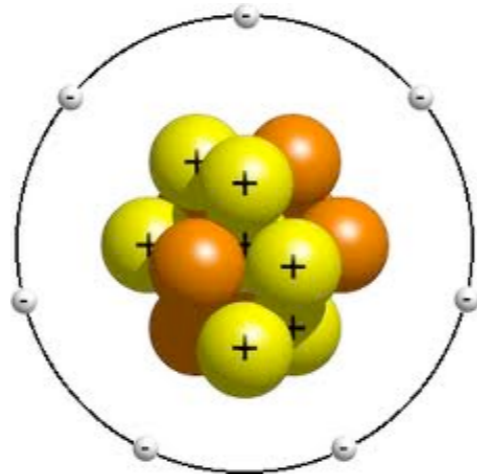
Heavy DM



large recoil...
“no problem”

Elastic nuclear recoils don't work

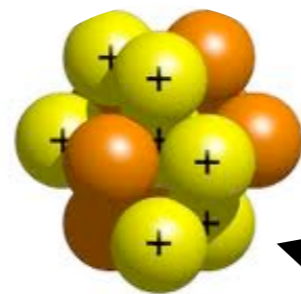
Light DM $\lesssim 1$ GeV



Atom

Elastic nuclear recoils don't work

Light DM $\lesssim 1$ GeV



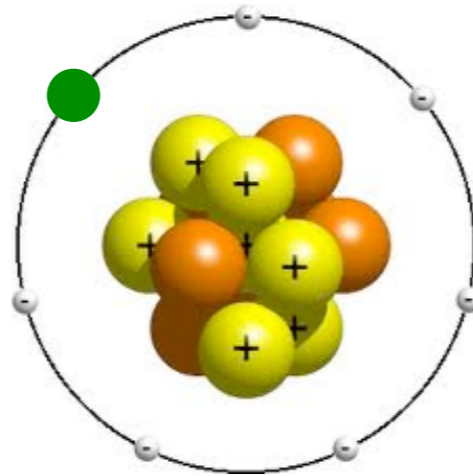
DM

Can't see
recoiling nucleus

But DM could scatter off an electron!

RE, J. Mardon, T. Volansky (1108.5383, PRD)

DM



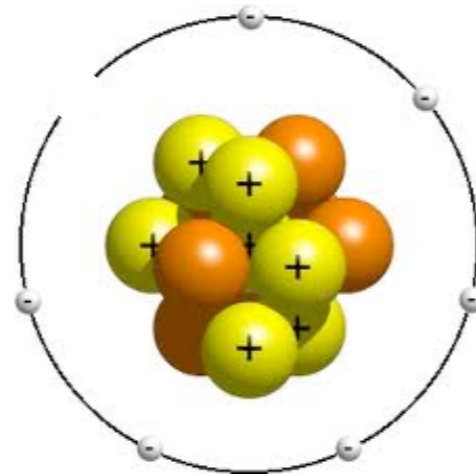
Atom

(only an example)

But DM could scatter off an electron!

RE, J. Mardon, T. Volansky (1108.5383, PRD)

DM



(only an example)



Atom

Ionization

Signal: single (or few) electron events

existing technologies can measure ionization,
even of a single electron !

A Proof of Principle

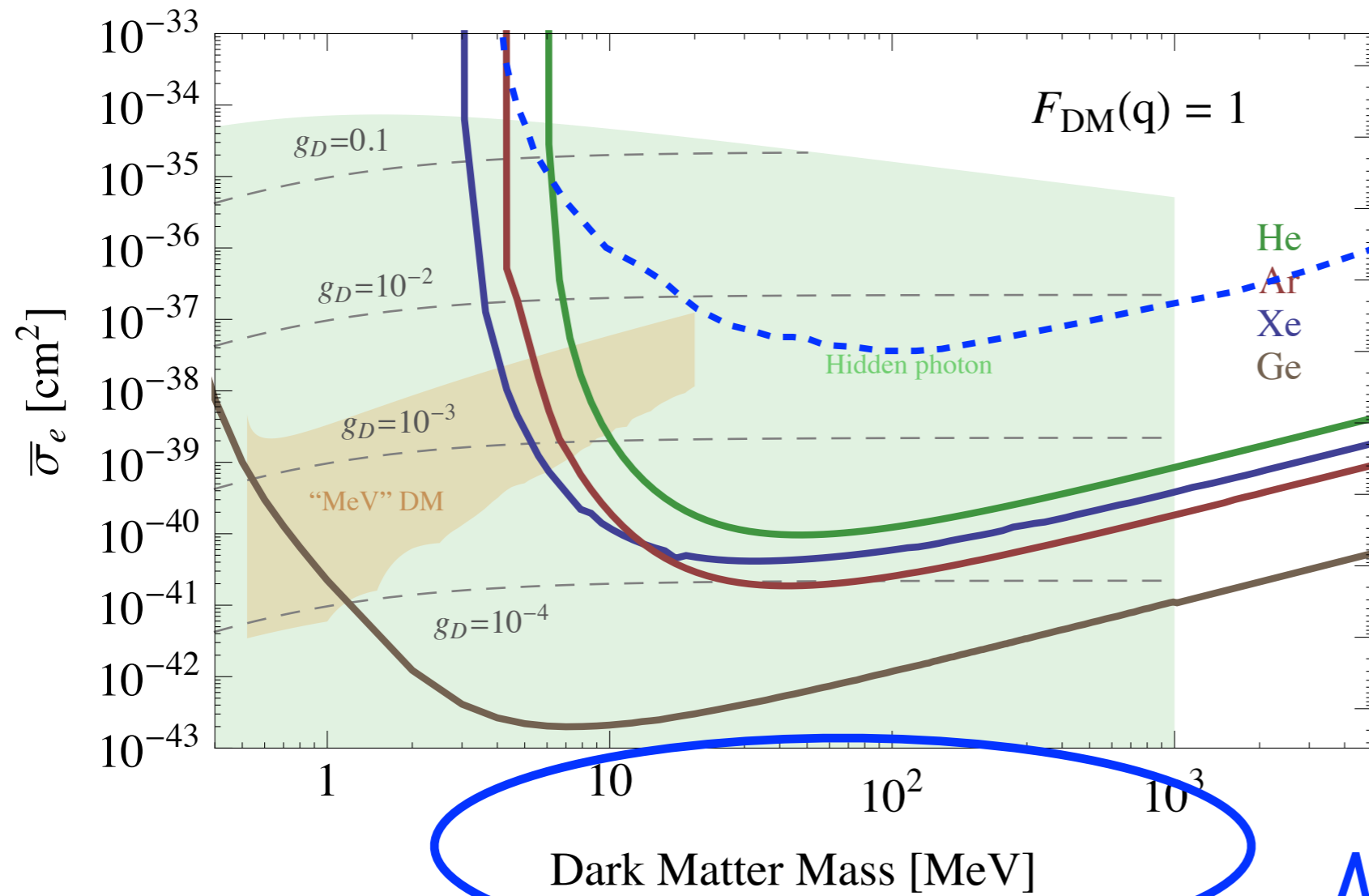
“First direct detection limits on
sub-GeV Dark Matter from XENON10”

RE, A. Manalaysay, J. Mardon, P. Sorensen, T. Volansky
(1206.2644, PRL)

see talk by A. Manalaysay

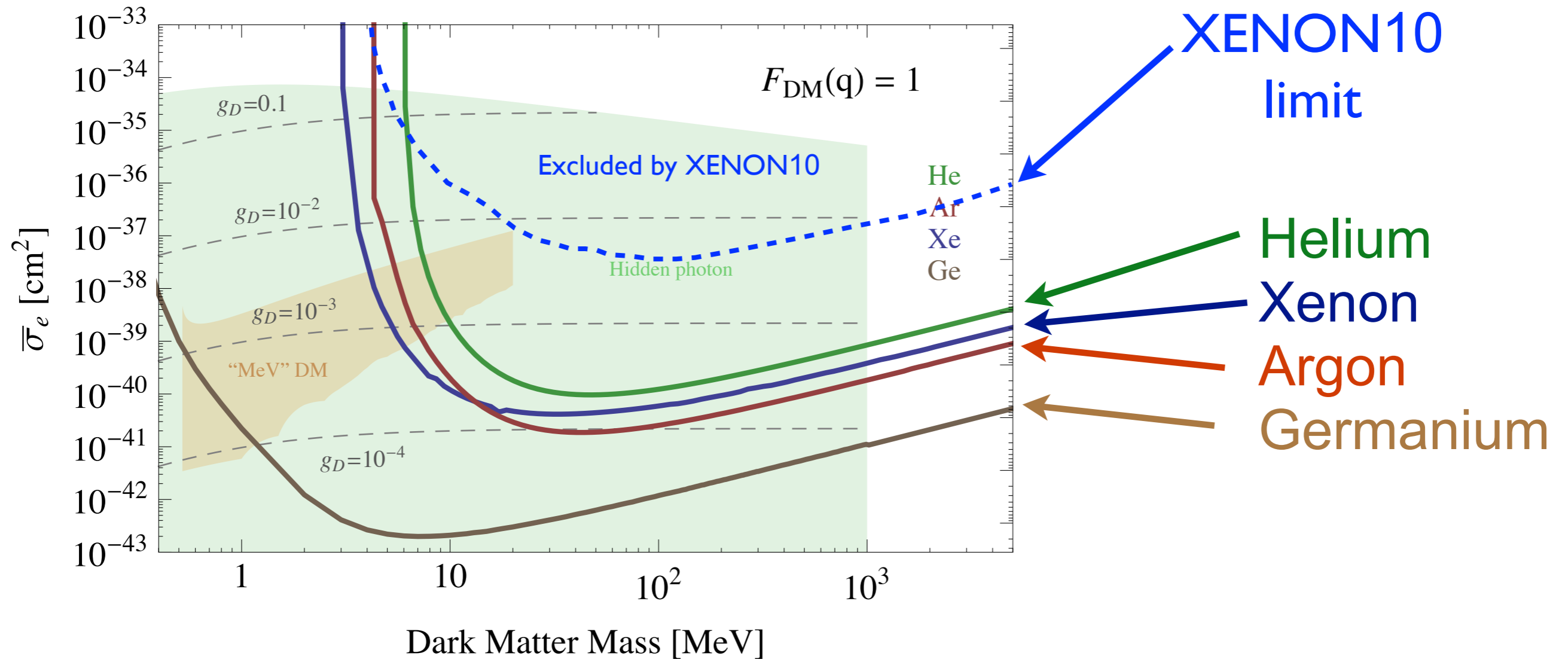
sub-GeV Dark Matter

Cross section Sensitivity and Event Rate (per kg·year)



sub-GeV Dark Matter

Cross section Sensitivity and Event Rate (per kg·year)



already have a limit down to a few MeV!

enormous potential for future experiments

Summary

- **Axions and WISPs**
 - motivated by strong CP, theory, DM, astro anomalies, ...
 - light-shining-through-walls, helioscopes, resonant cavities
- **Heavy Photons (> 1 MeV)**
 - motivated by theory, muon $g-2$, DM anomalies, ...
 - opportunities at e^+e^- colliders, e^- & p fixed target, LHC, ...
- **Direct detection of sub-GeV DM**
 - beware of WIMP or axion prejudice
 - much more work required to exploit experimental potential