

Recent Measurements from the Yale Microwave Cavity Experiment (YMCE)

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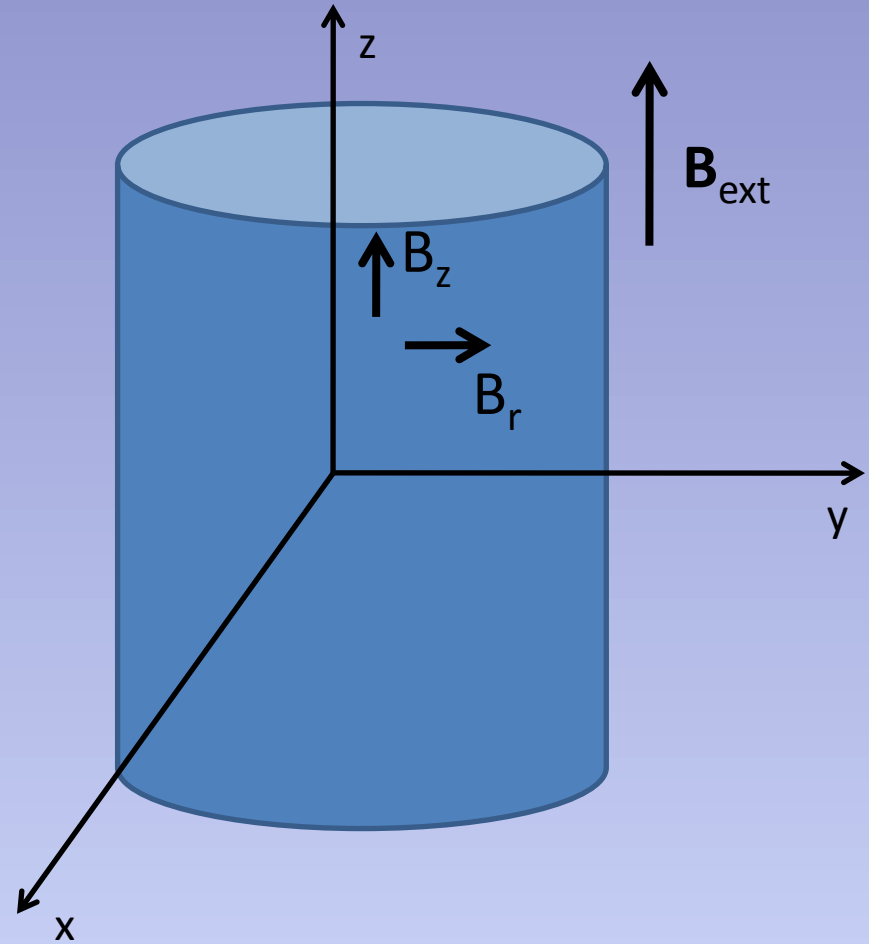
YMCE Measurements at 34 GHz

- Scalar ALPs
- Scalar Chameleons
- Hidden Sector Photons
- Pseudoscalar ALPs
with TM cavity



Resonant cavity searches

- Pioneered by Pierre Sikivie and ADMX collaborators in 1980s.
- ADMX: Single low-mode cavity in B-field to look for $\gamma\gamma$ coupling with local galactic halo axions (μeV).
- $\mathcal{L} = -g\phi \mathbf{B}_\gamma \cdot \mathbf{B}_{\text{ext}}$ (scalar ALP)
- $\mathcal{L} = -g\phi \mathbf{E}_\gamma \cdot \mathbf{B}_{\text{ext}}$ (pseudoscalar ALP)

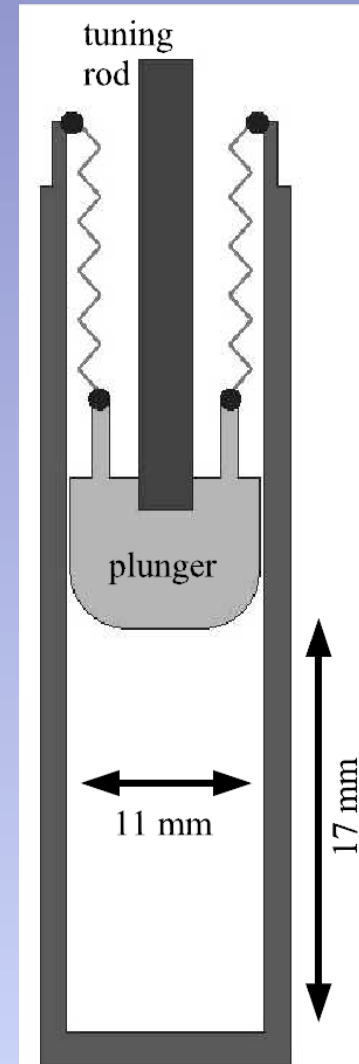


YMCE experiment at 34 GHz

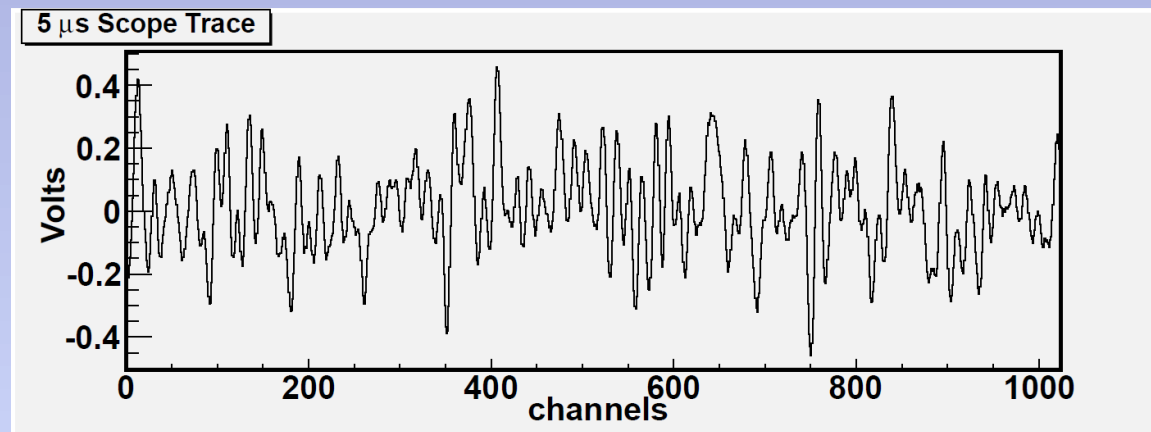
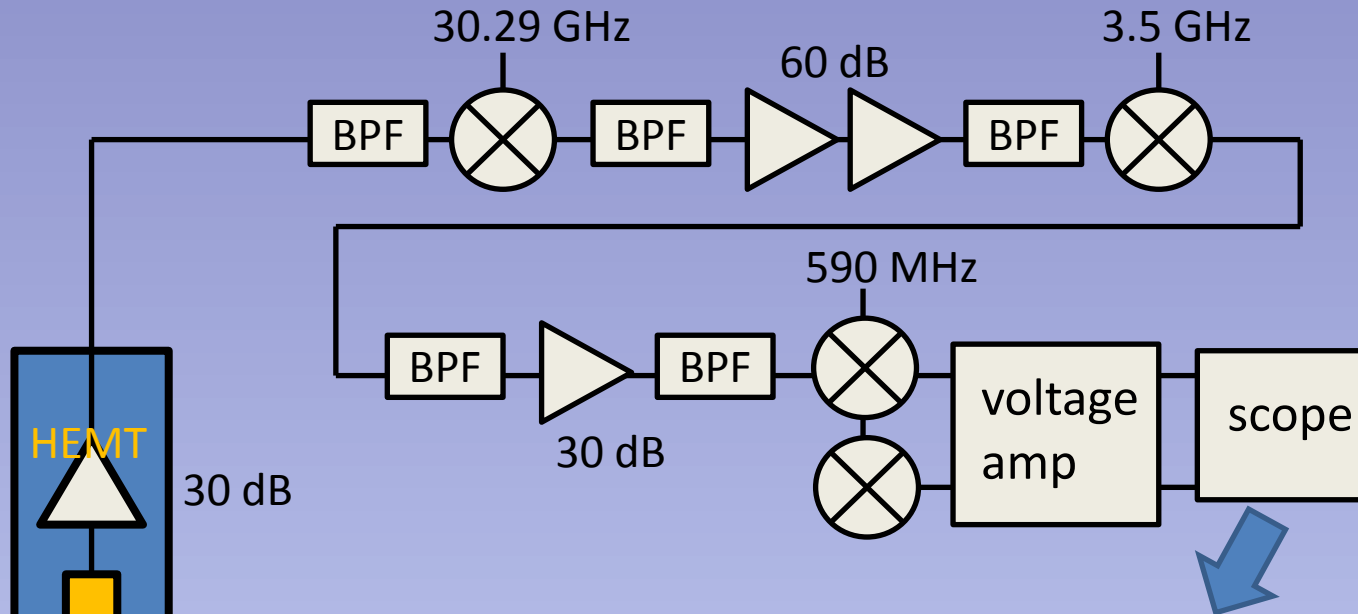
- Cu resonant cavity at 34 GHz, cooled to $T \sim 4$ K, tunable, TE₀₁₁ mode.
- From the Lagrangian,

$$P_{S\gamma} = g_{S\gamma\gamma}^2 V B_{ext}^2 \rho_a C_{lmn} Q$$

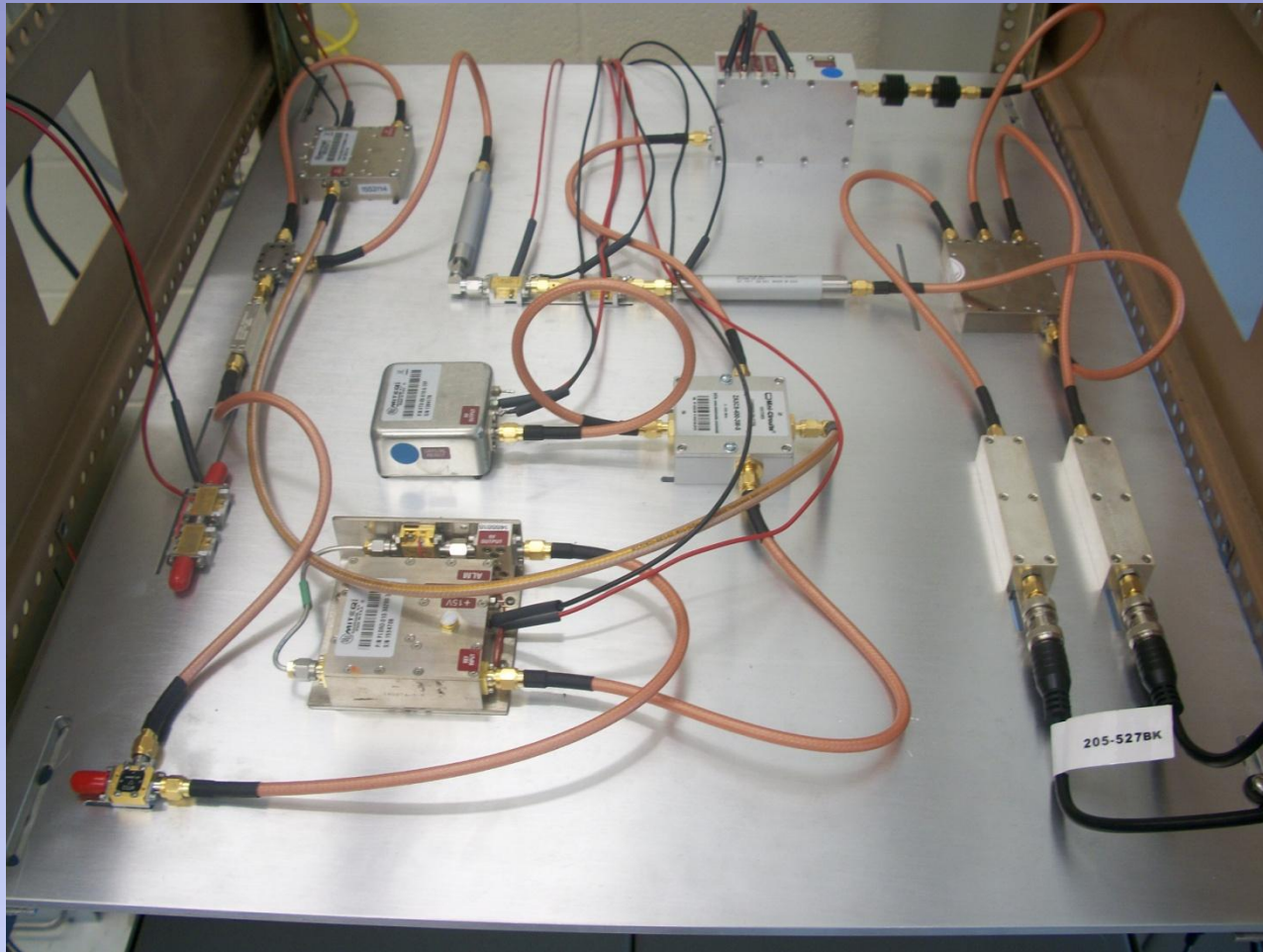
$$C_{lmn} \equiv \frac{\left| \int_V d^3x \mathbf{B} \cdot \hat{\mathbf{B}}_{ext} \right|^2}{V \int_V d^3x \frac{1}{\mu} |\mathbf{B}|^2}$$



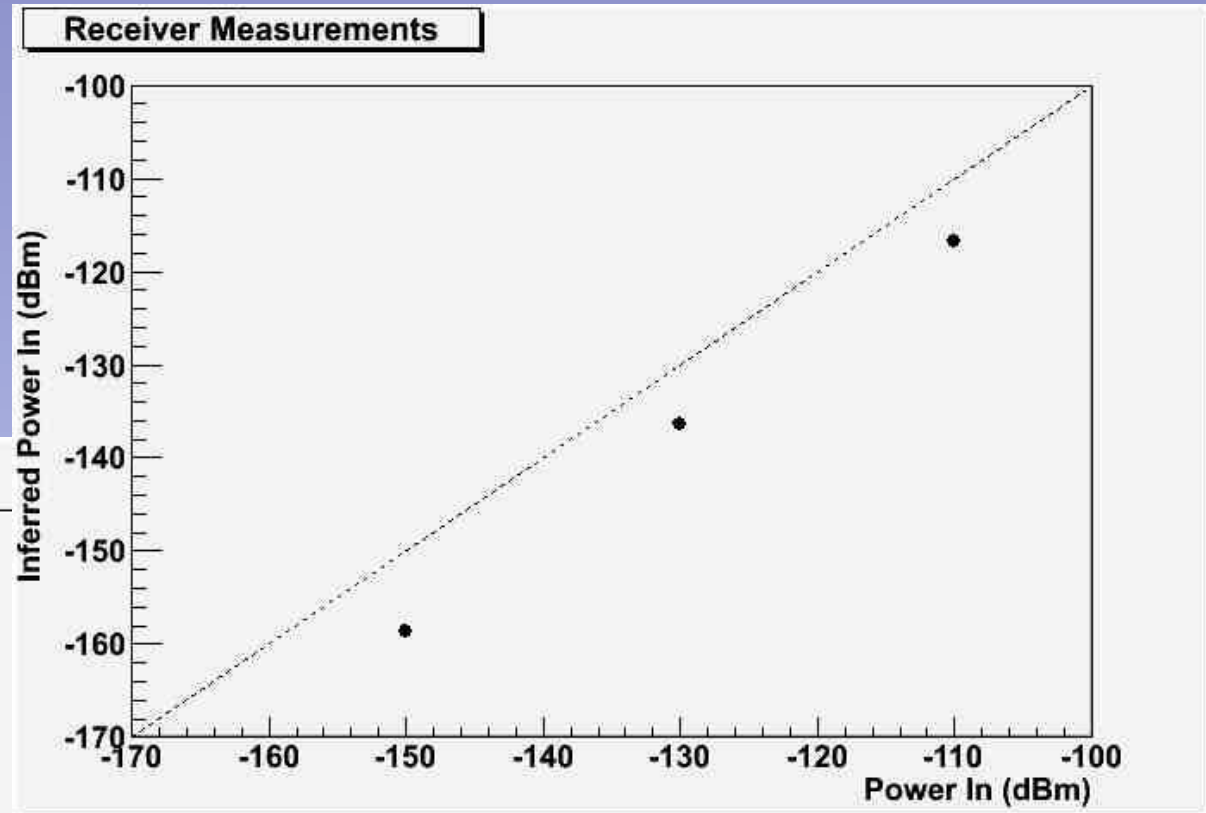
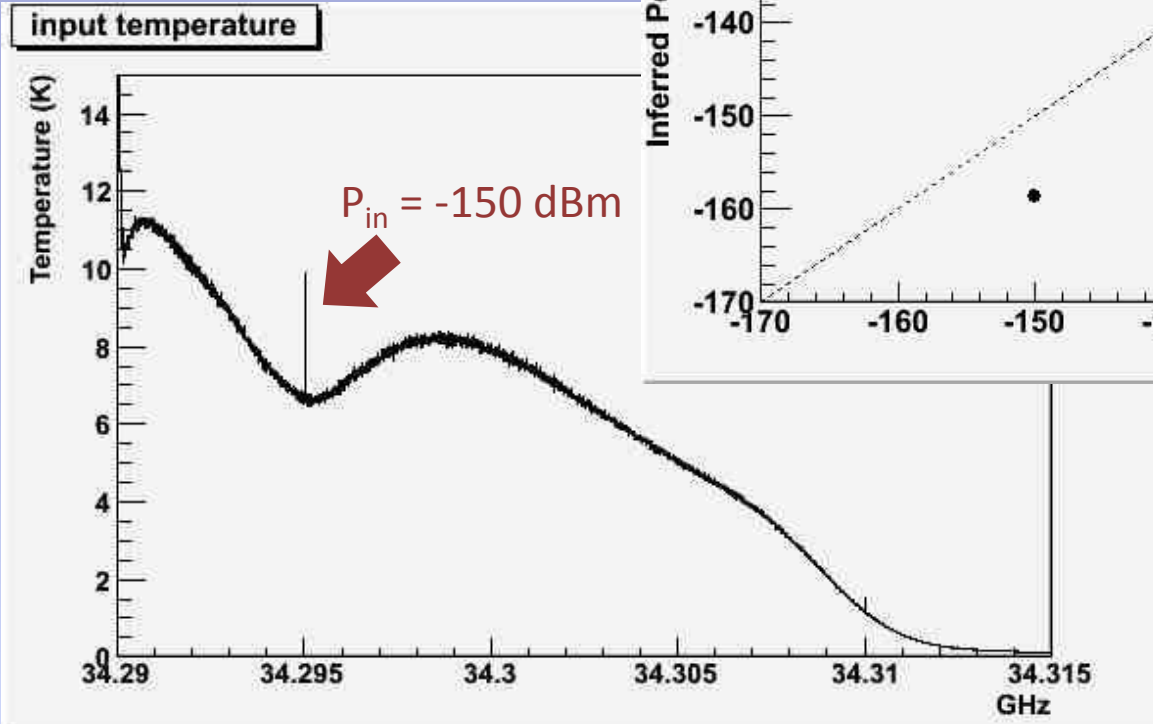
Microwave receiver



Receiver layout

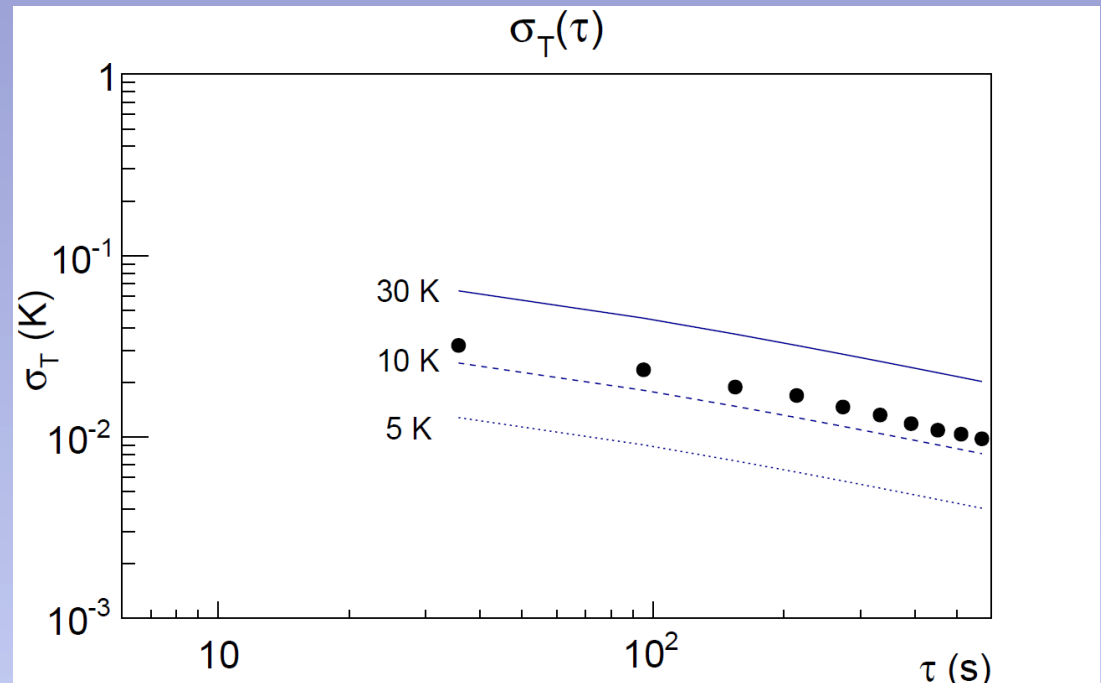


Receiver Tests

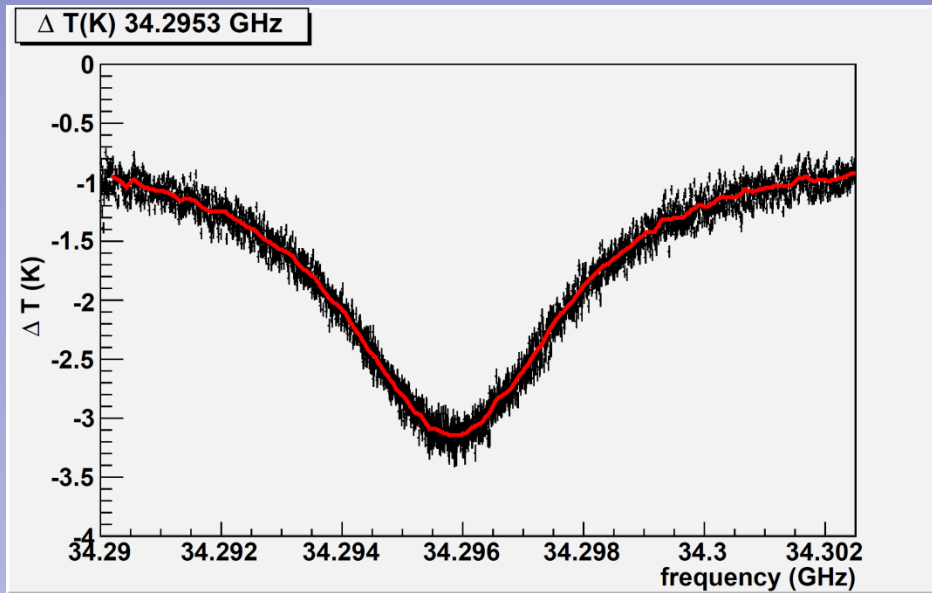


Noise floor of system

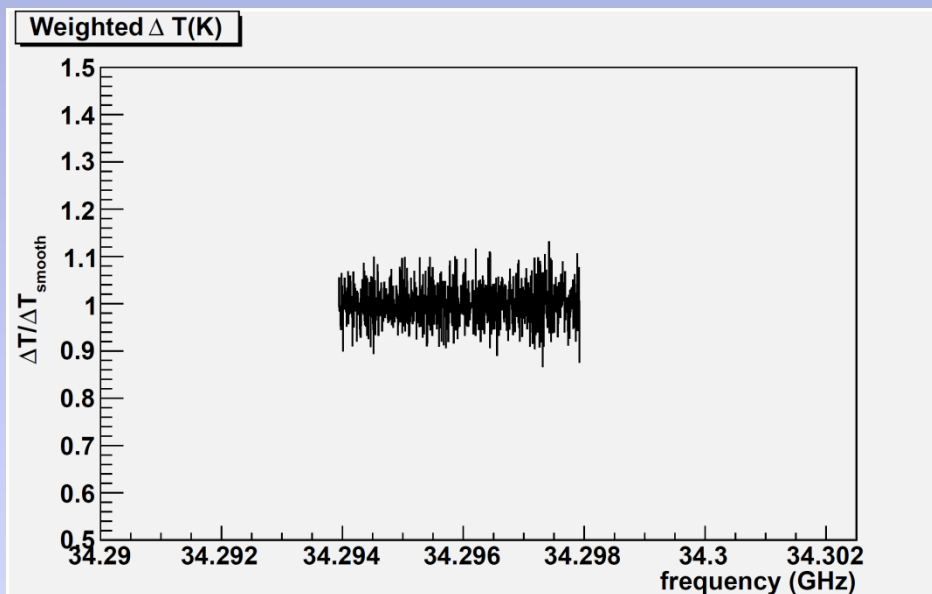
- $T_{\text{sys}} = T_{\text{cavity}} + T_{\text{HEMT}}$
 $\sim 15 \text{ K}.$
- $\sigma_T \sim 10 \text{ mK}$ for
 $\Delta\nu = 5 \text{ kHz}.$
- $P_{\text{min}} = k_B \Delta\nu T$
 $= 10^{-21} \text{ W}.$



One scan, cavity tuned to 34.295 GHz

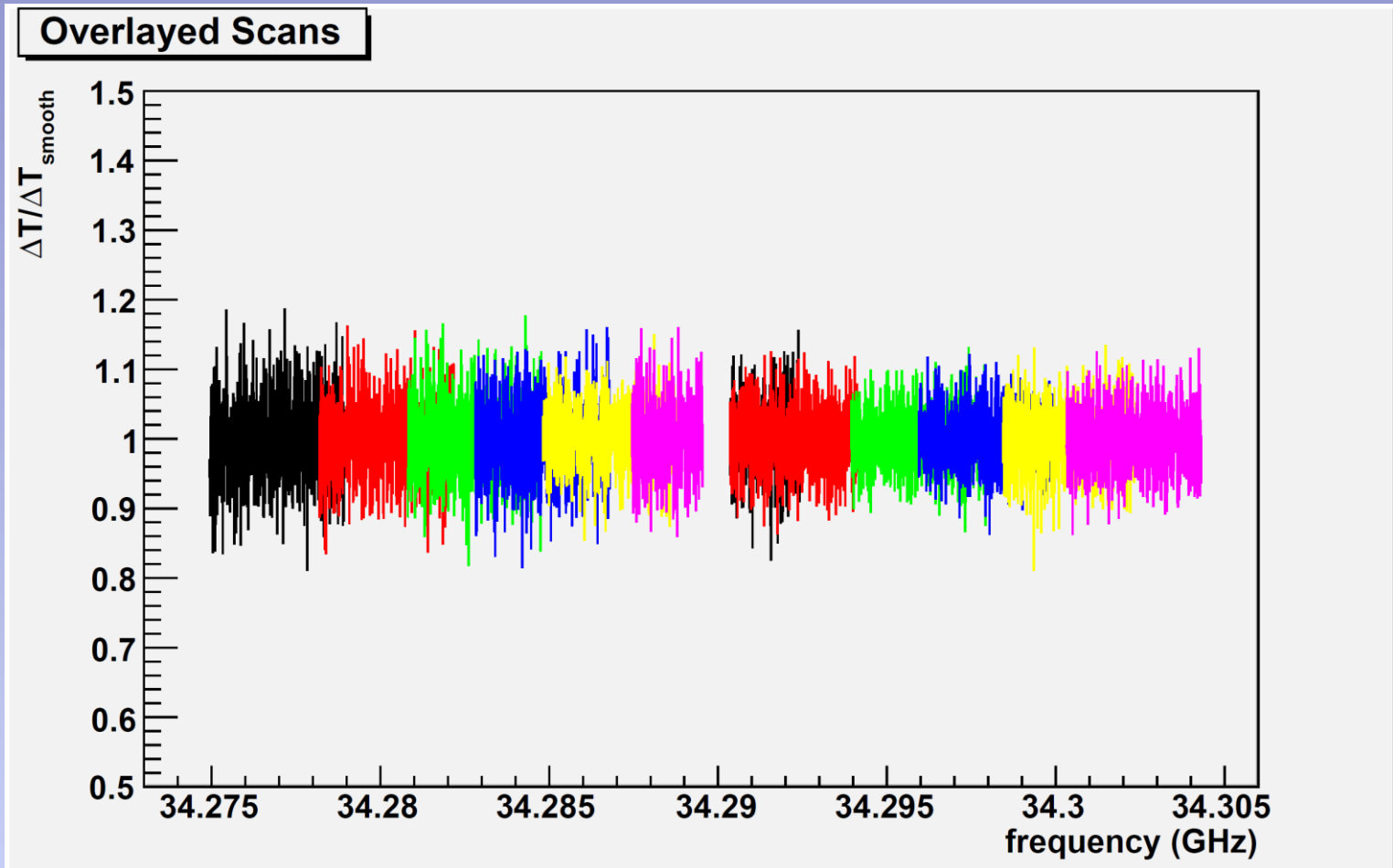


After baseline subtraction.

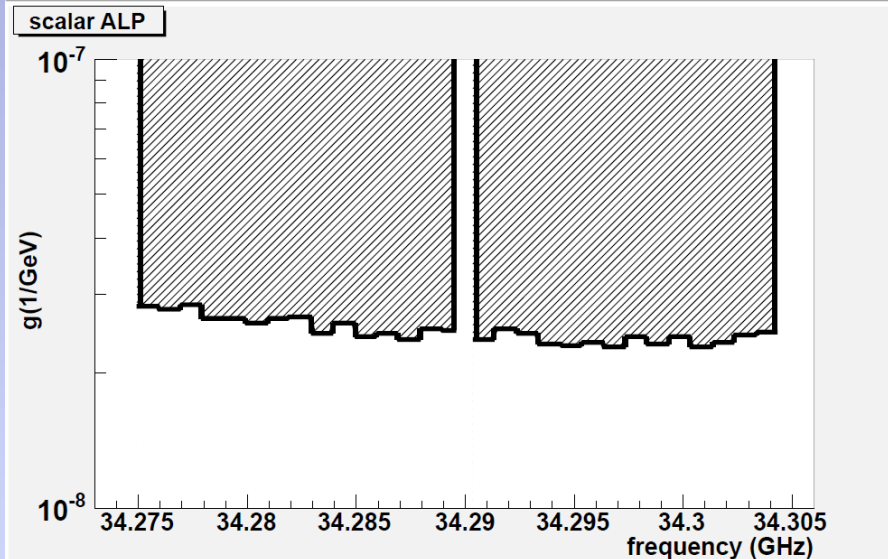
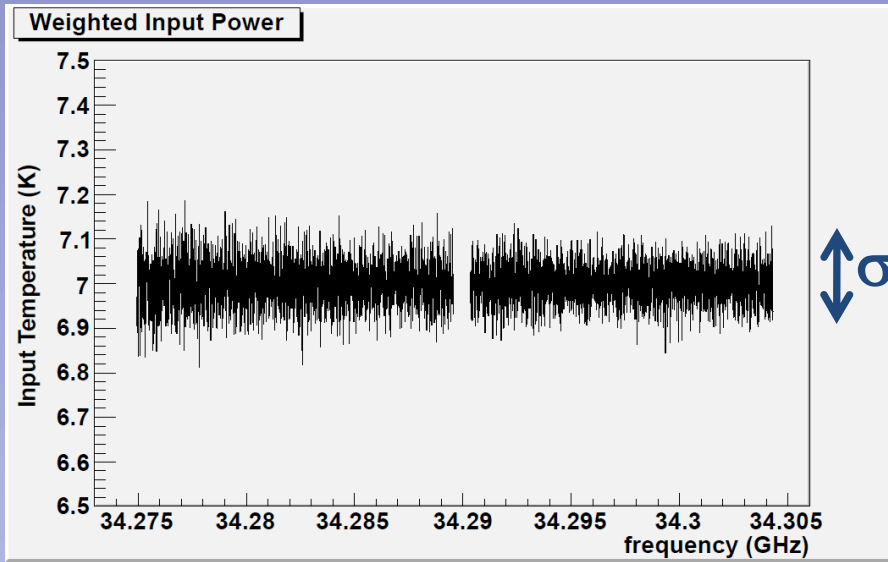


After baseline subtraction and weighting, cut on cavity.

Overlaid scans



Averaged scans and results

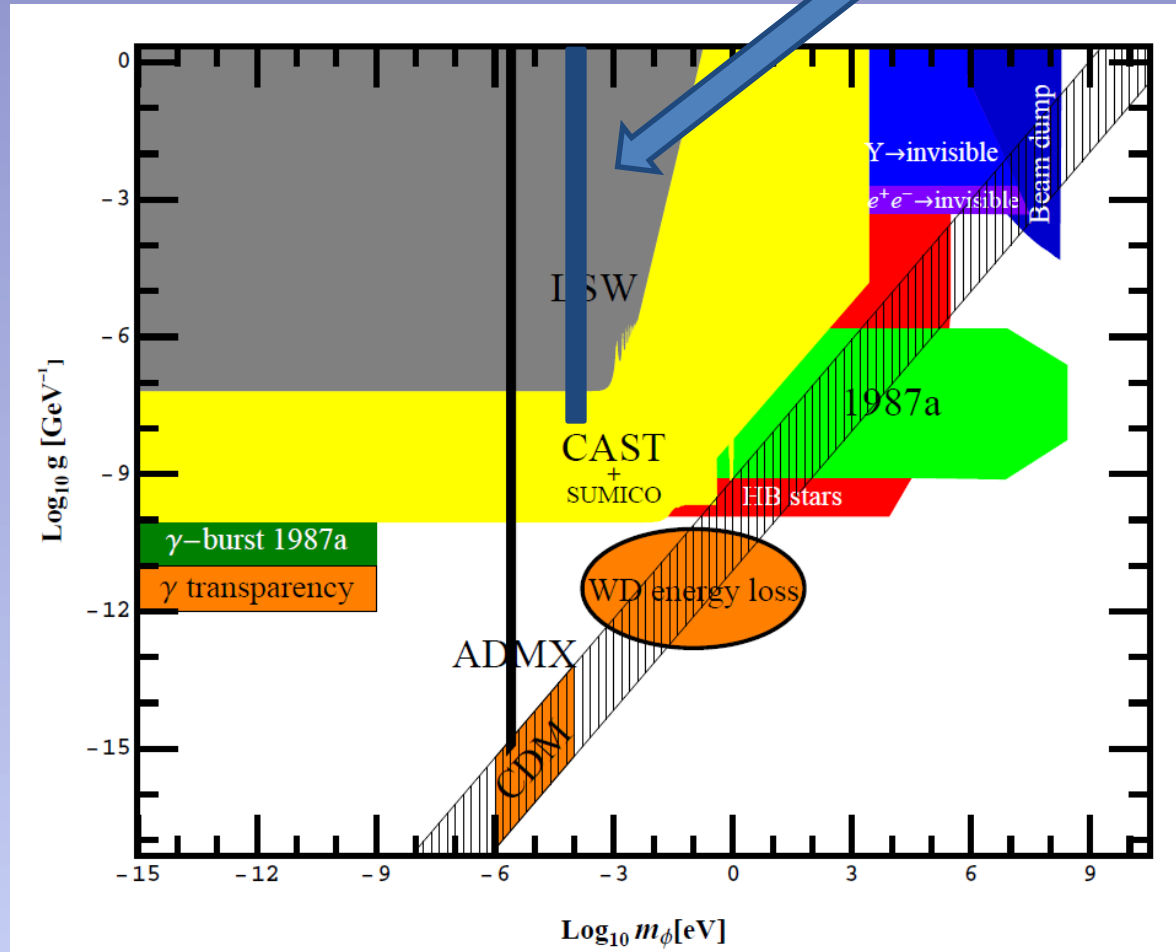


$$P_{S\gamma} = g_{S\gamma\gamma}^2 V B_{ext}^2 \rho_a C_{lmn} Q$$

$$C_{lmn} \equiv \frac{\left| \int_V d^3x \mathbf{B} \cdot \hat{\mathbf{B}}_{ext} \right|^2}{V \int_V d^3x \frac{1}{\mu} |\mathbf{B}|^2}$$

submitted for publication.

Status of axions and ALPs



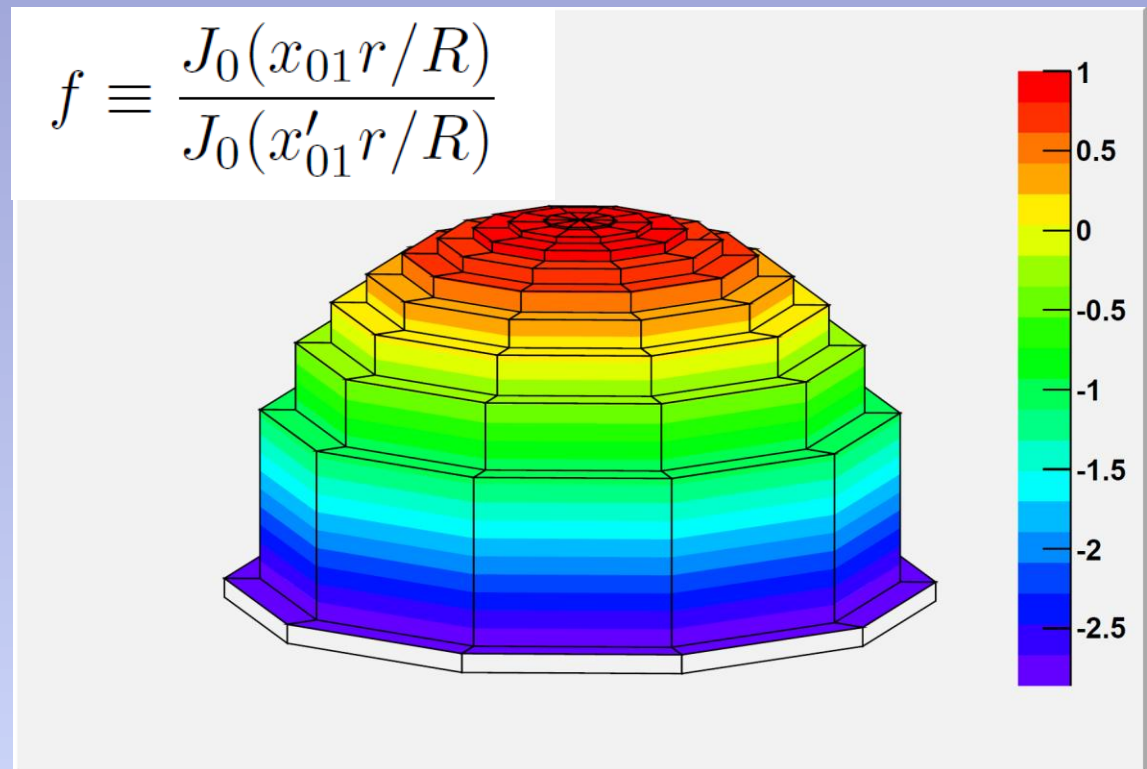
Scalar chameleons

- Dark energy candidate.
- m_{eff} depends on local energy densities $\rho_m(\mathbf{x})$ and $\rho_\gamma(\mathbf{x})$ with coupling strengths β and β_γ .^(*)
- After excitation chameleons could be trapped inside resonant cavity according to $\rho_m(\mathbf{x})$.
- Significant overlap between TE_{011} and 011 chameleon cavity modes.

(*) P. Brax, C. van de Bruck, A.-C. Davis, J. Khoury, and A. Weltman, Phys. Rev. D 70, 123518 (2004).

Overlap between TE_{mnp} mode and ϕ_{mnp} mode

- Apply TE_{mnp} and ϕ_{mnp} mode solutions to modified wave equations. (*)
- Overlap is f .
- Average of f over cavity volume is ~ 0.6 .



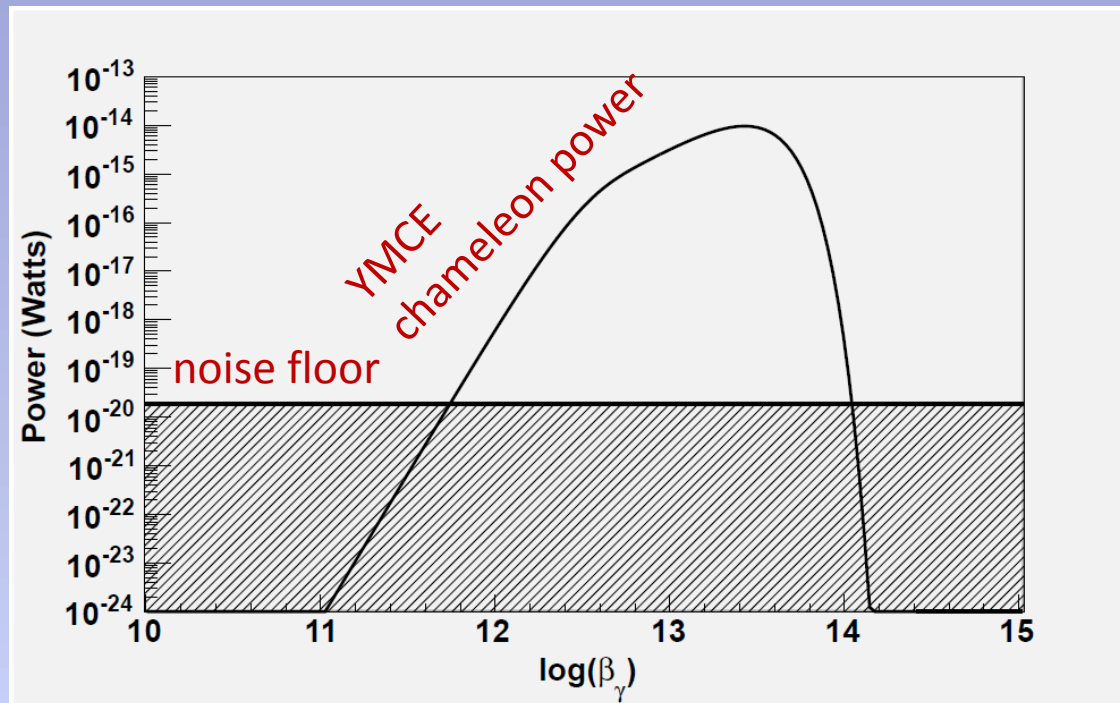
(*) G. Rybka, ADMX internal note: Scalar chameleon search, 2009.

Chameleon power in YMCE apparatus

$$P_{\text{out}} = P_{\text{in}} \frac{\pi\Gamma}{2b} (1 - e^{-\Gamma/2t_0})^2 e^{-\Gamma t} \quad (*)$$

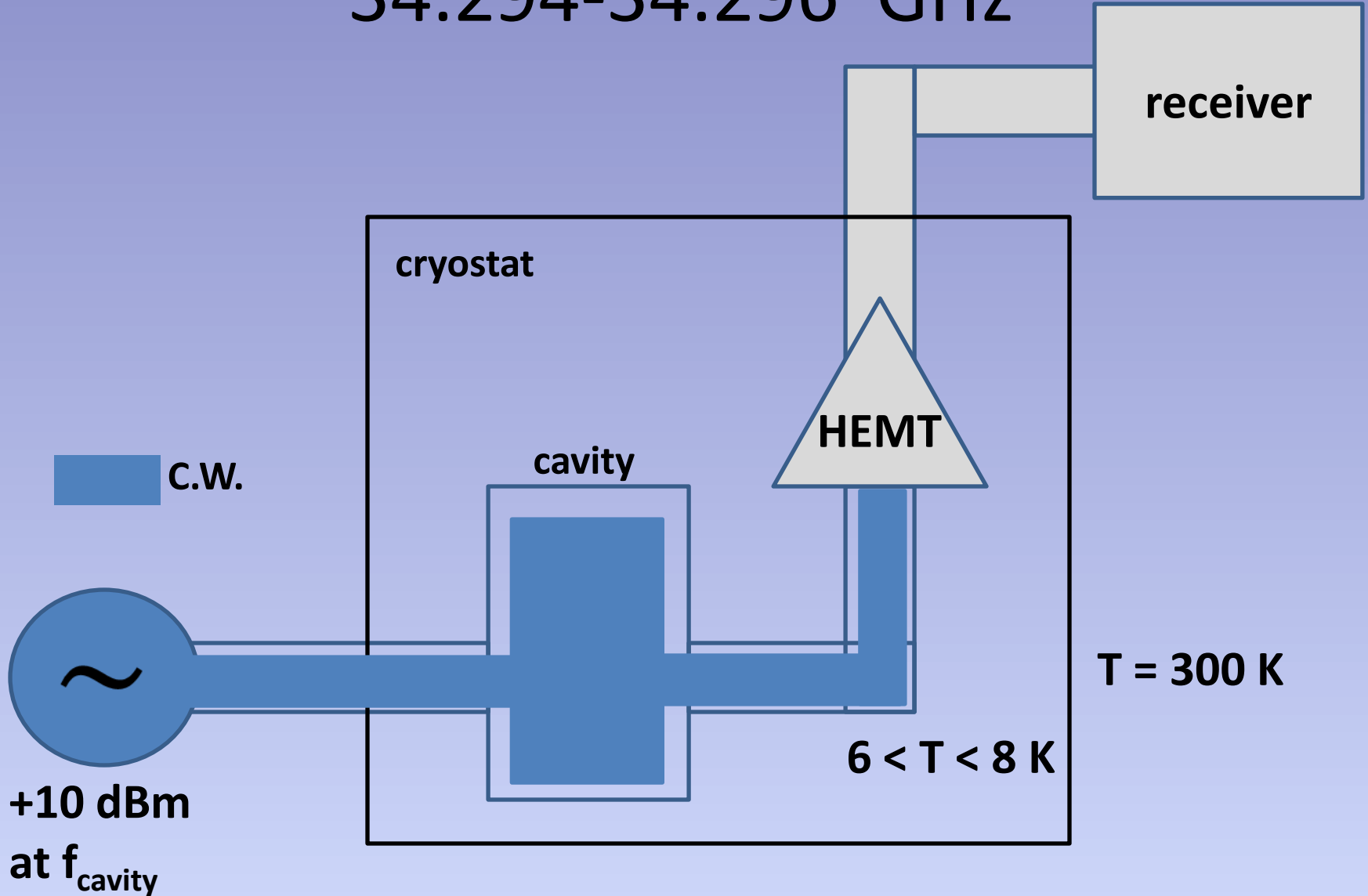
$$\Gamma = \frac{\beta_\gamma^2 f^2 B^2 Q k_{\text{tr}}^2}{M_{\text{pl}}^2 \omega^3} \quad (*)$$

$\ll \Delta v_{\text{RF}}$ in experiment.



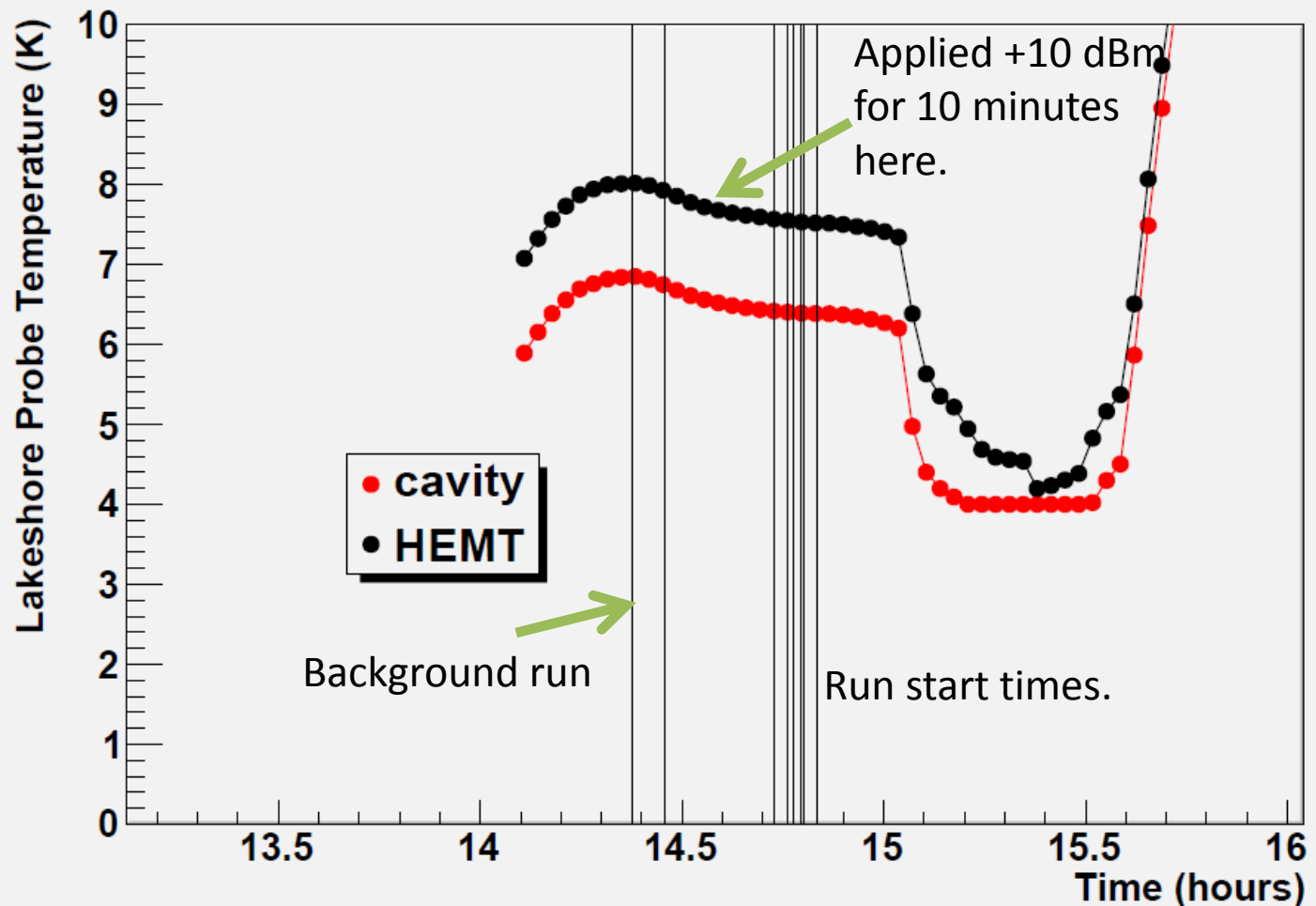
(*) G. Rybka et al., PRL **105**, 051801 (2010)
P. Sikivie Phys. Rev. D **32**, 2988 (1985)

+10 dBm for 10 min. at
34.294-34.296 GHz

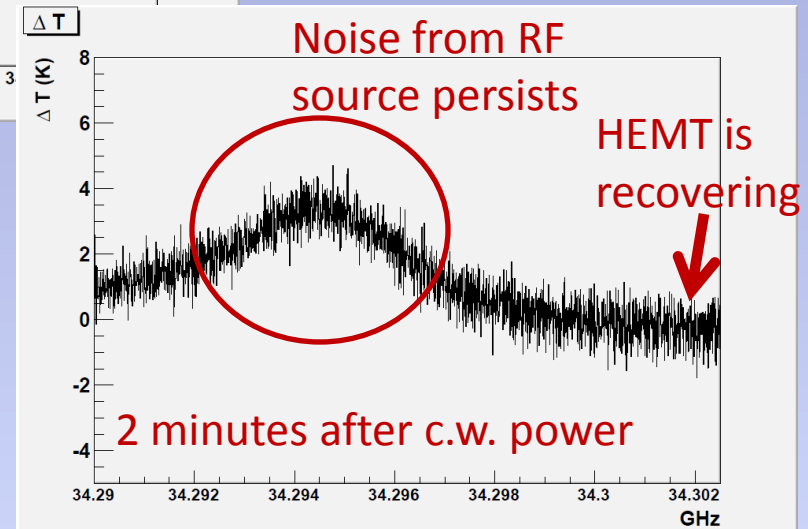
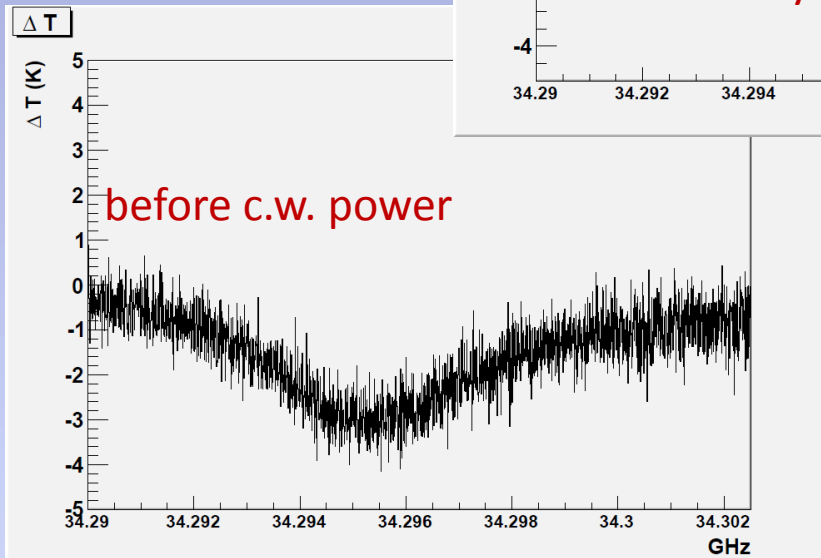
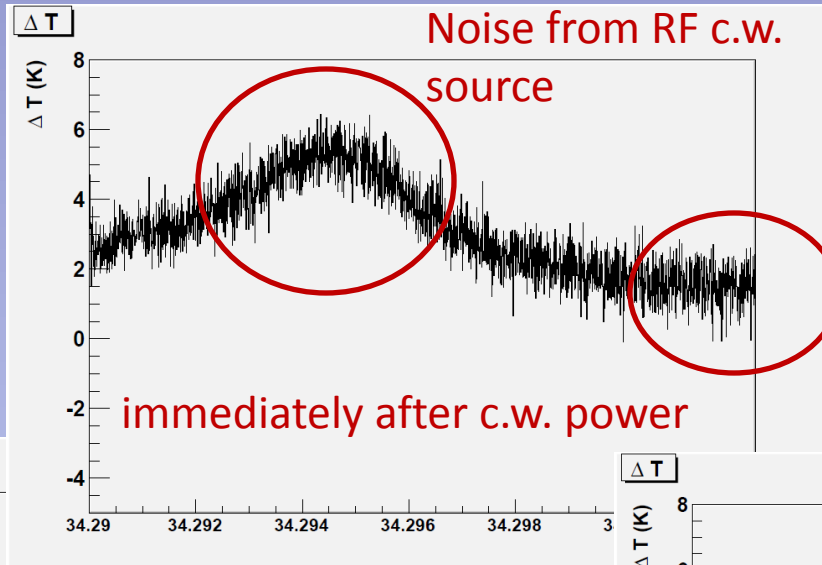


Temperatures inside cryostat

June 28, 2012



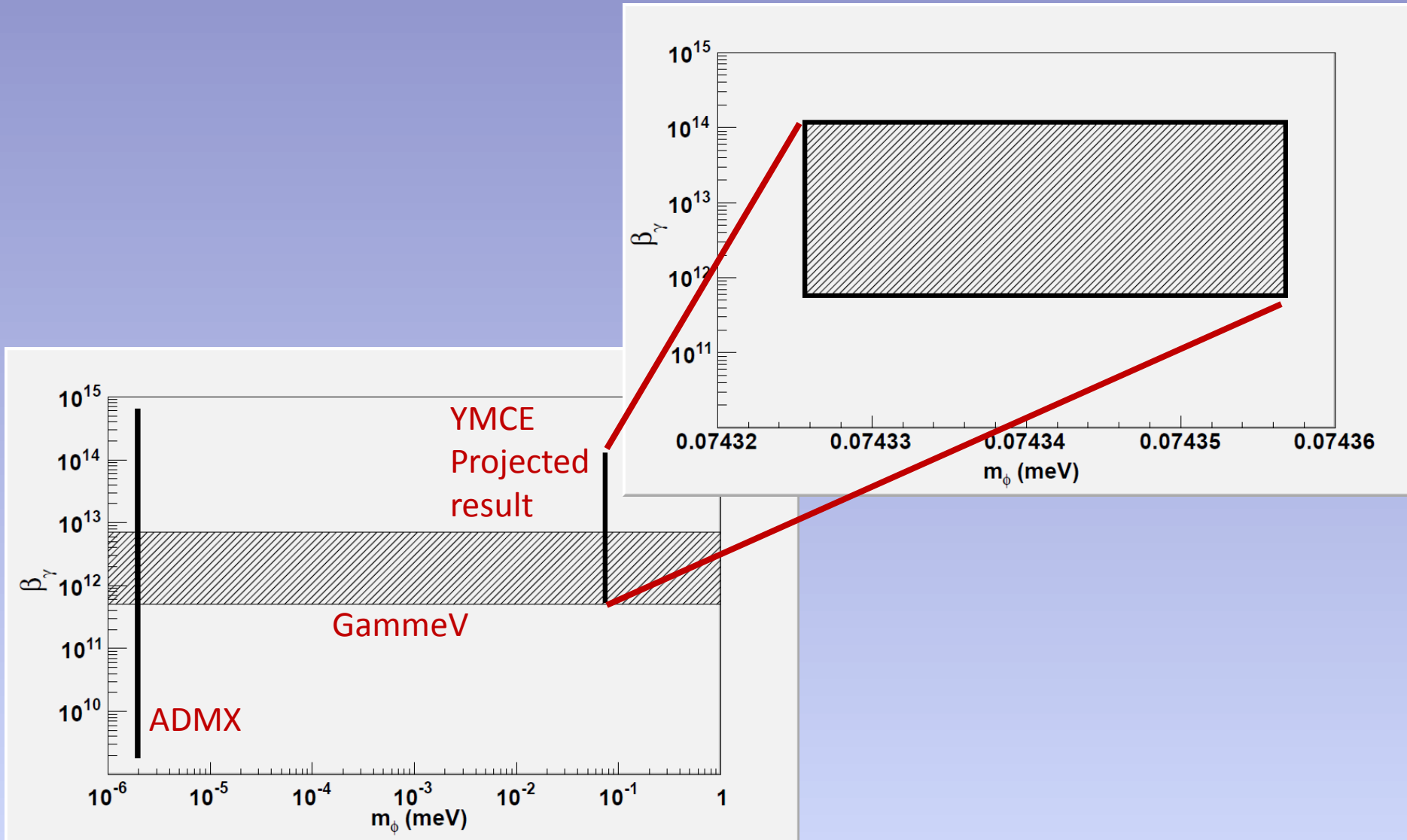
System shake-out for chameleon search



Status of chameleon search

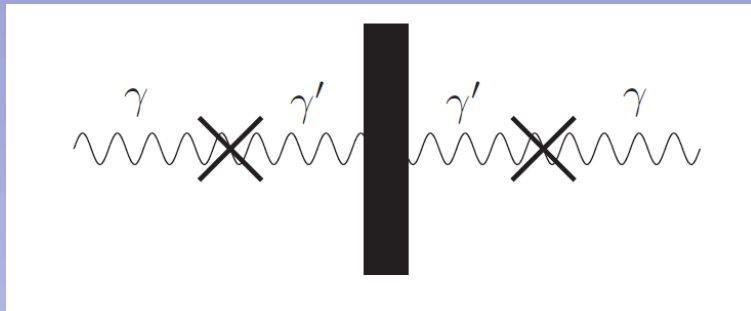
- HEMT behavior can be explained by traps in transistor semiconductor. Charge carriers are knocked out of traps by strong c.w. power. Time constant can be minutes.
- This effect should be reduced with lower c.w.
- RF leak can be eliminated with improved switching.

Chameleon searches



Hidden photons

For LSW-type experiment,
2 cavities are needed.



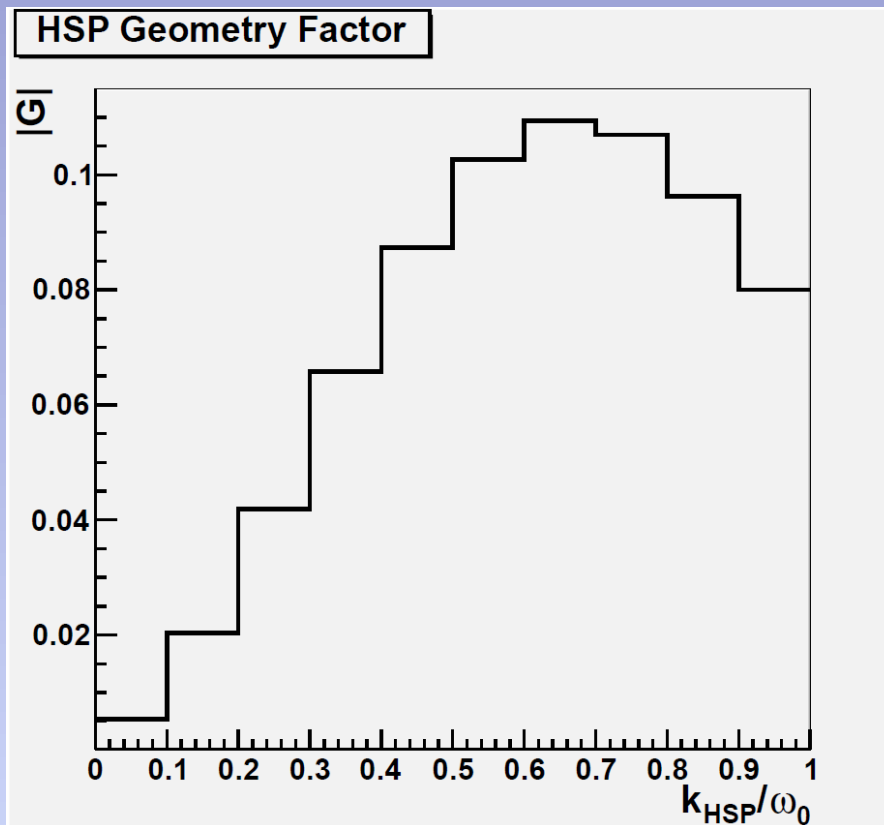
$$\mathcal{L} = \chi m_{\gamma'}^2 A_{\mu} X^{\mu}$$

$$P_{trans} = \chi^4 Q Q' \frac{m_{\gamma'}^8}{\omega_0^8} |G_{HSP}|^2.$$



Geometry Factor* – hidden photons

$$G_{HSP} \equiv \omega_0^2 \int_{V'} \int_V d^3\mathbf{x} d^3\mathbf{y} \frac{\exp(ik|\mathbf{x} - \mathbf{y}|) A(\mathbf{y}) A'(\mathbf{x})}{4\pi|\mathbf{x} - \mathbf{y}|},$$

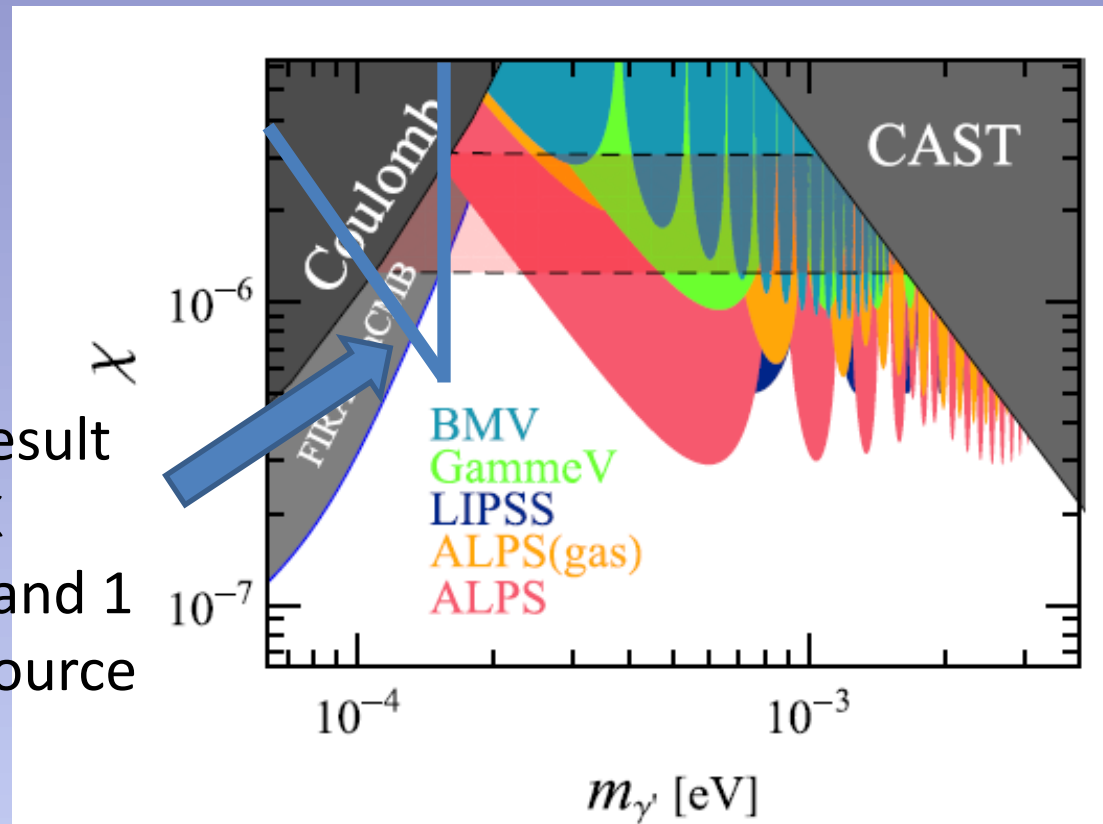


Separates **geometry information** (e.g. cavity fields and their overlap) from the remainder of the calculation.

* J. Jaeckel and A. Ringwald, Phys. Lett. B 659 (3) 509, 2008.

Sensitivity to hidden photons

Expected result
with 10 mK
resolution and 1
Watt c.w. source



* K. Ehret et al, Physics Letters B, Volume 689, Issue 4-5, p. 149-155.

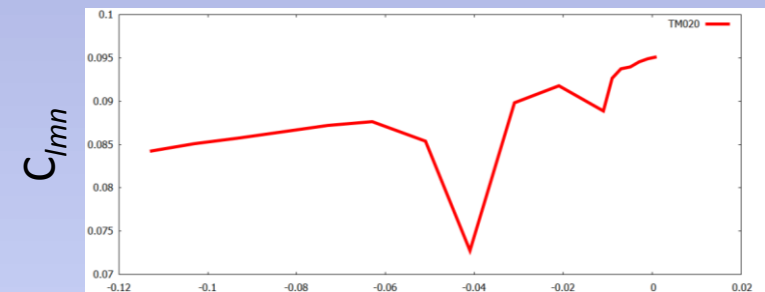
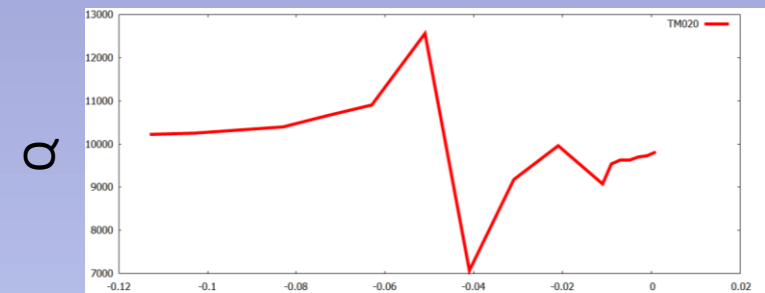
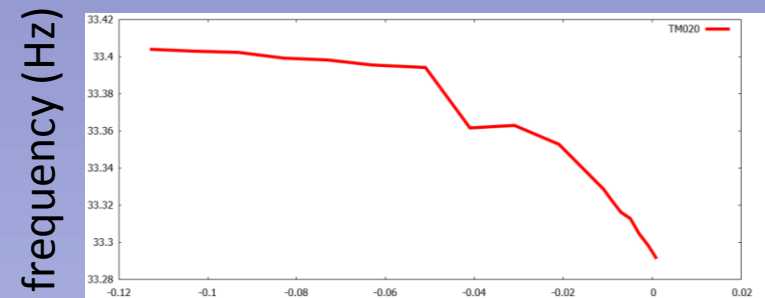
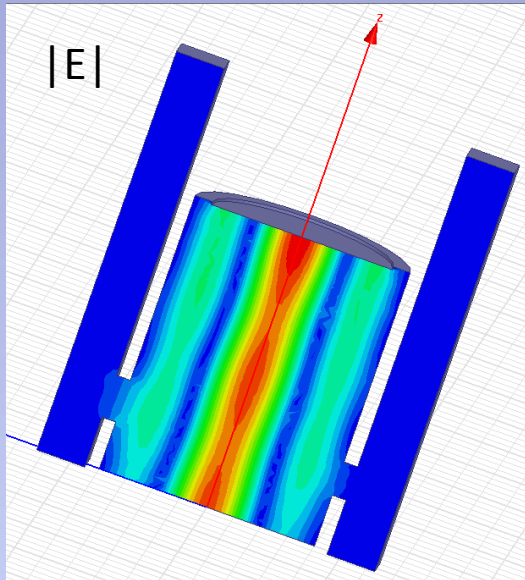
Working prototype of TM_{020} cavity

- Sensitive to pseudoscalar $\alpha\gamma\gamma$.
- Bench tests underway.
- Limited tunability with dielectric post and spring-loaded contacts.
- Cavity performance is modeled in HFSS software.



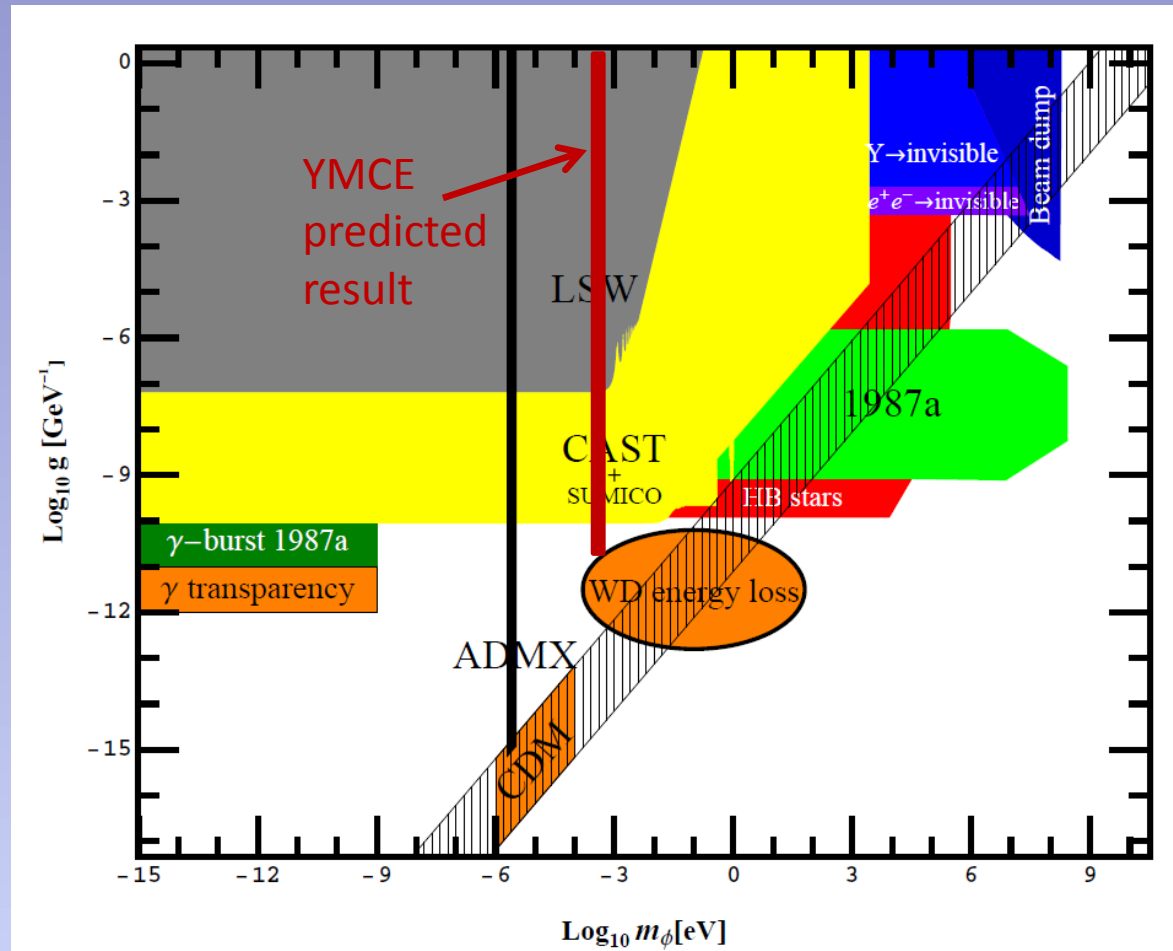
HFSS simulations for TM_{020} cavity

- Tuning range of 100 MHz.
- $C_{l_{mn}}$ is 0.1 .



post height (inches)

1 cm TM_{020} cavity with HEMT, $\Delta\nu_{RF} = 5$ kHz and $\sigma_T = 10$ mK



Summary

- Scalar ALP measurement is the first narrow band exclusion of $S^0_{\gamma\gamma}$ at 0.1 meV .
- Chameleon search at 0.1 meV is underway.
- Drive cavity for HSP search is now complete and ready to be installed into the experiment.
- Bench tests with early TM cavity have begun.

Auxiliary slides

TE modes in cylindrical cavity with chameleons

$$\nabla^2 B_z - \frac{\partial^2}{\partial t^2} B_z = \frac{\beta_\gamma}{M_P} \mathbf{B} \left(\frac{\partial^2}{\partial z^2} \phi - \frac{\partial^2}{\partial t^2} \phi \right)$$
$$\nabla^2 \phi - \frac{\partial^2}{\partial t^2} \phi - m_\phi^2 \phi = \frac{\beta_\gamma}{M_P} B_z \mathbf{B}$$

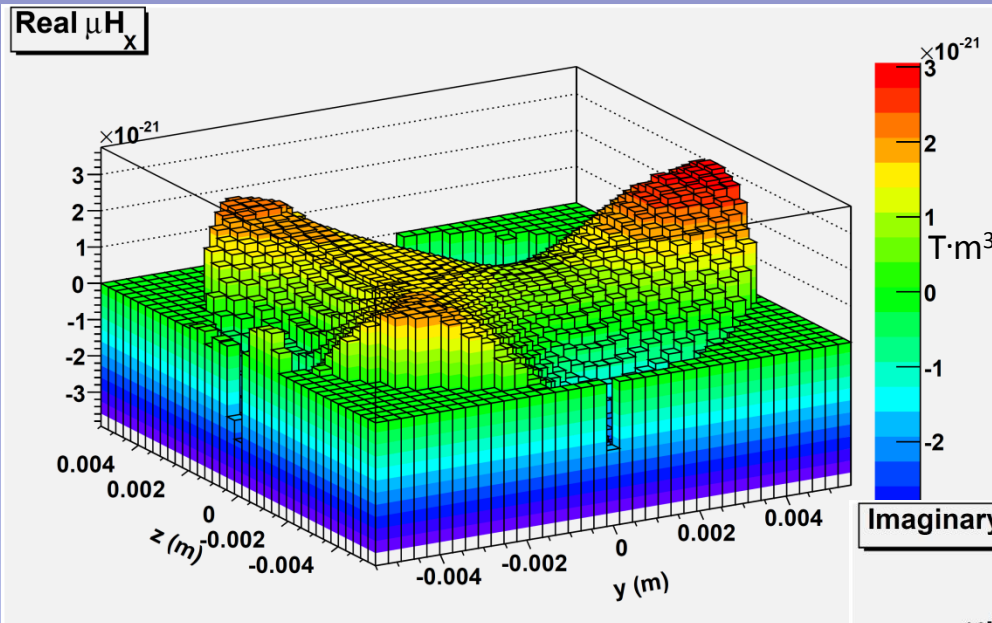
Wave equations in vacuum with scalar chameleons ϕ , in cylindrical cavity with longitudinal \mathbf{B} field. (*)

Wave function shapes of TE and ϕ mnp modes in cavity of radius R and height d . (*)

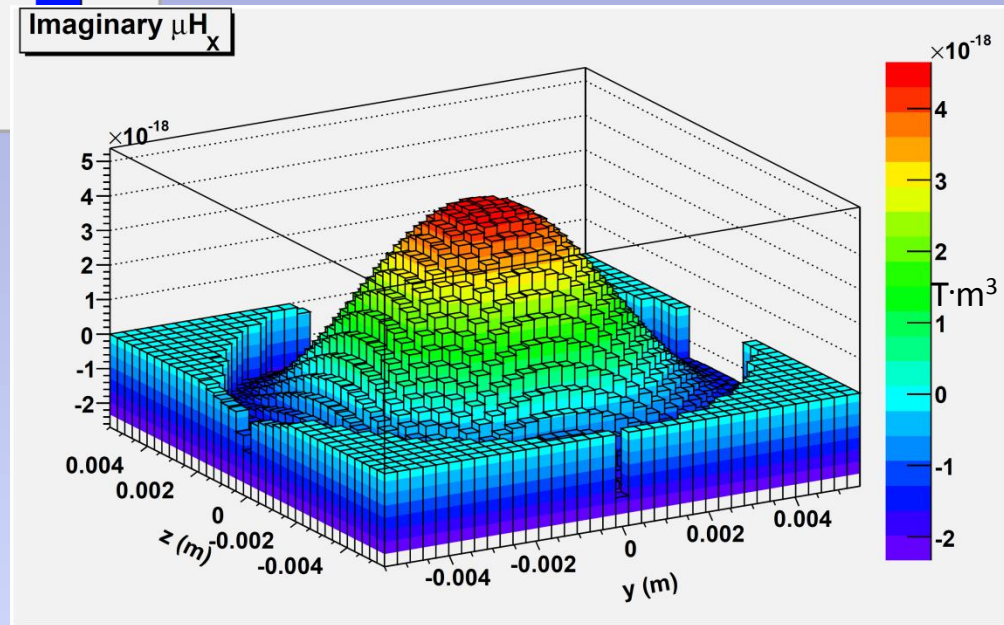
$$B_z = B_0 J_m \left(\frac{x'_{mn} r}{R} \right) \sin \left(\frac{p\pi z}{d} \right) e^{\pm im\phi}$$
$$\phi = \phi_0 J_m \left(\frac{x_{mn} r}{R} \right) \sin \left(\frac{p\pi z}{d} \right) e^{\pm im\phi}$$

(*) G. Rybka, ADMX internal note: Scalar chameleon search, 2009.
P. Sikivie, Phys. Rev. Lett. 48 (1982), 1156

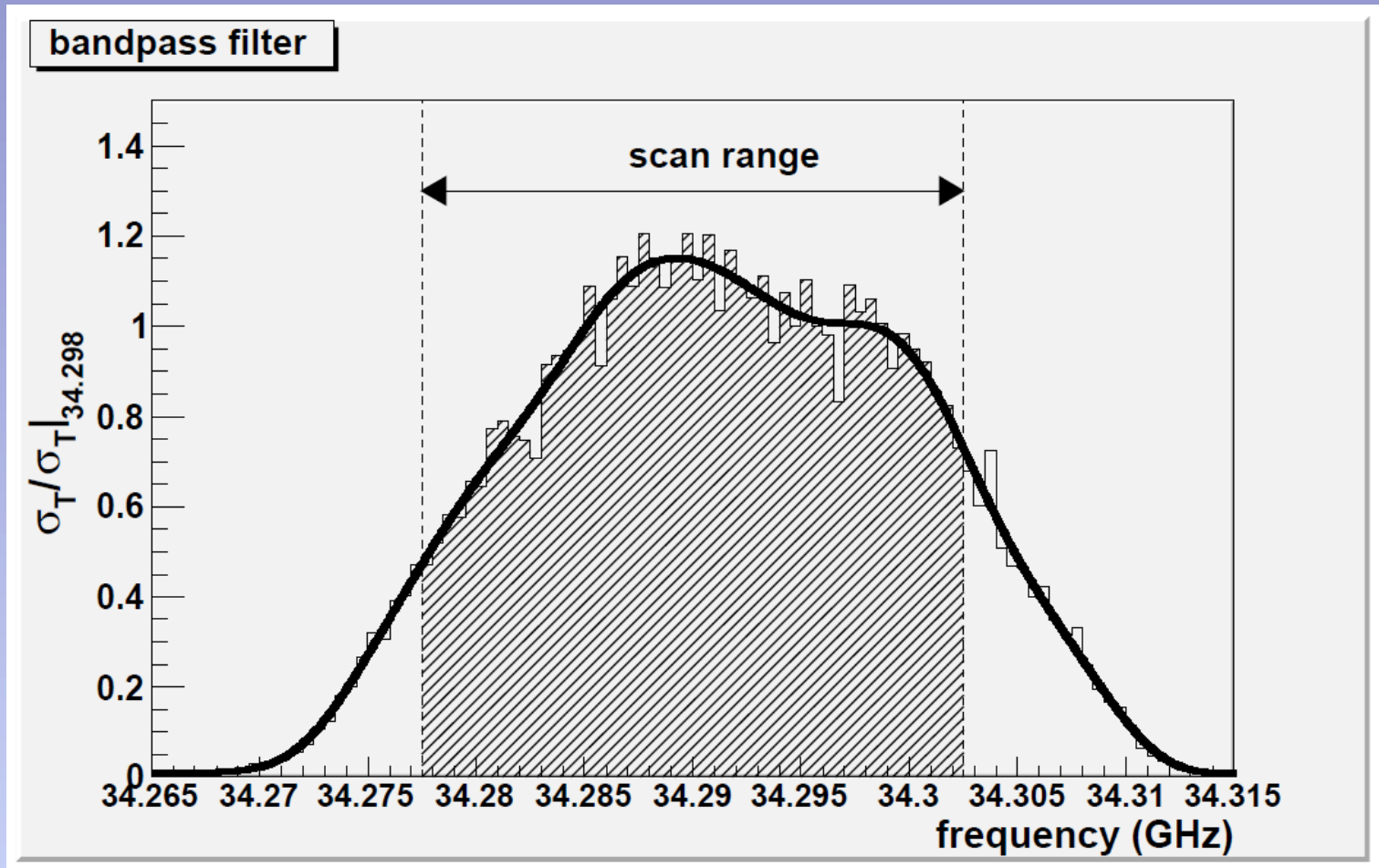
Simulated axial B fields



➔ $C_{lmn} = O(10^{-6})$



Receiver passband



Ideal signal to noise ratio

